

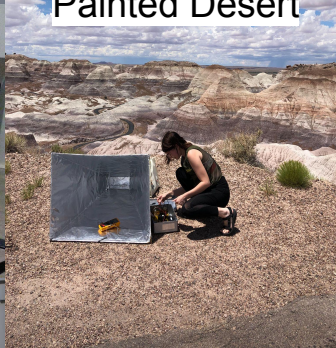
Signal Processing

CRASIES 5/26

HERA, Karoo Desert
South Africa



CHART
Observing Trip,
Painted Desert



Libby

- Current TSI postdoctoral fellow
- PhD at ASU in Exploration Systems Design (Instrumentation)
- B.S in Astronomy & Astrophysics from UChicago

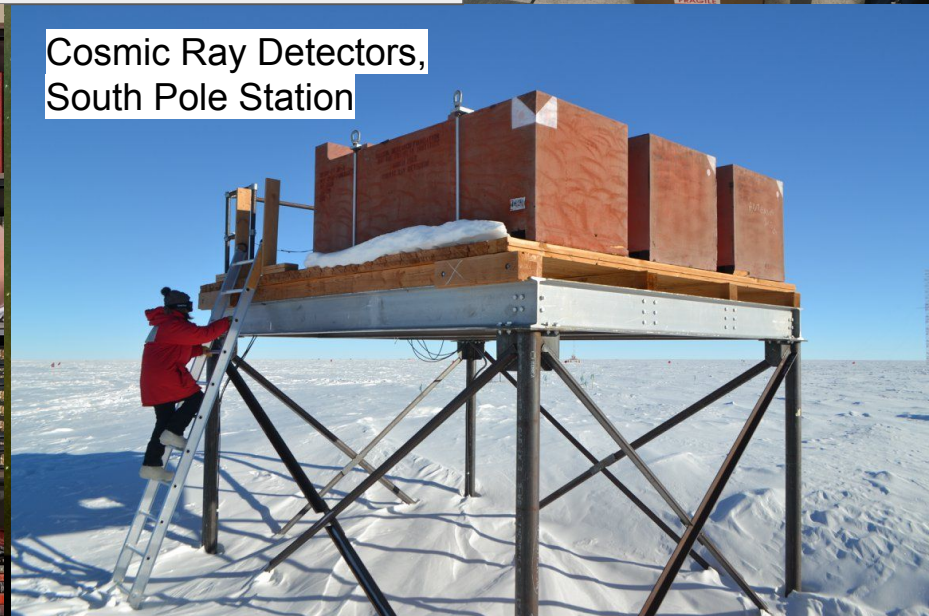


Goldstone Deep Space
Communications Center

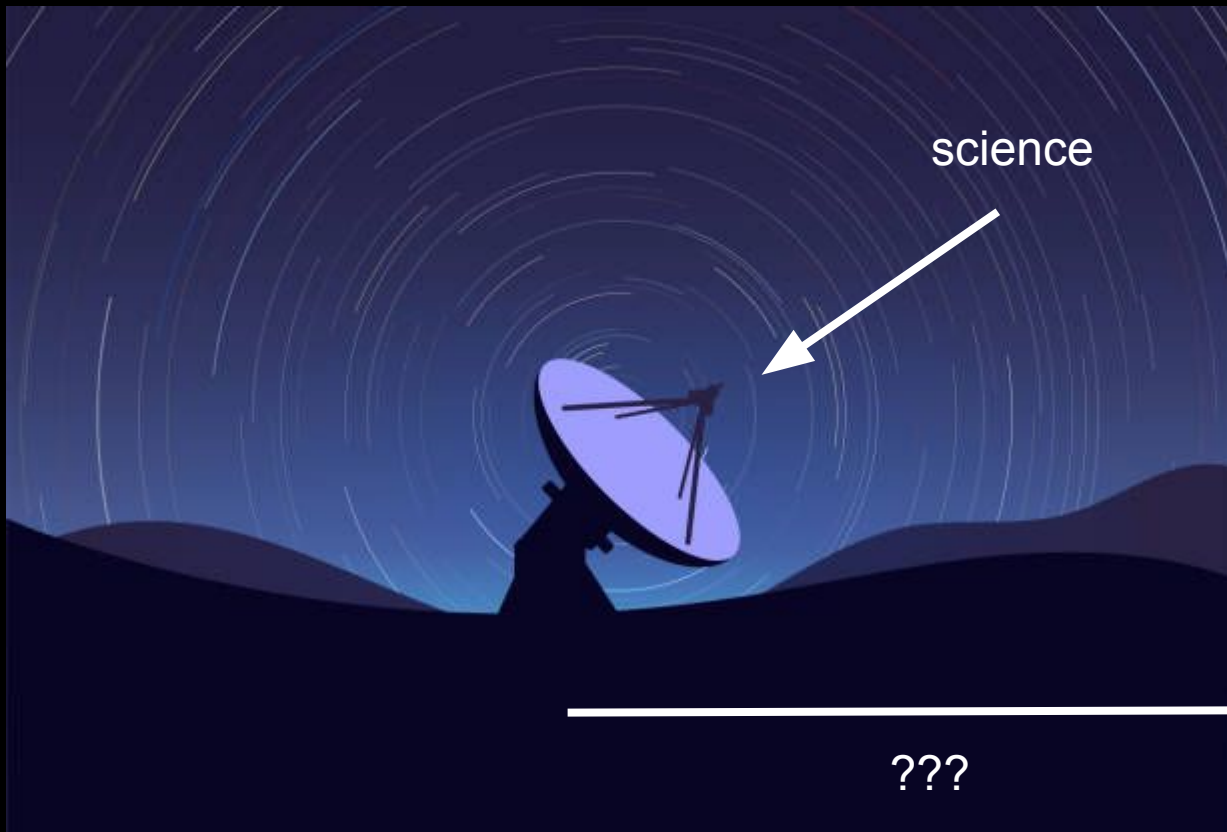


Cat

Cosmic Ray Detectors,
South Pole Station



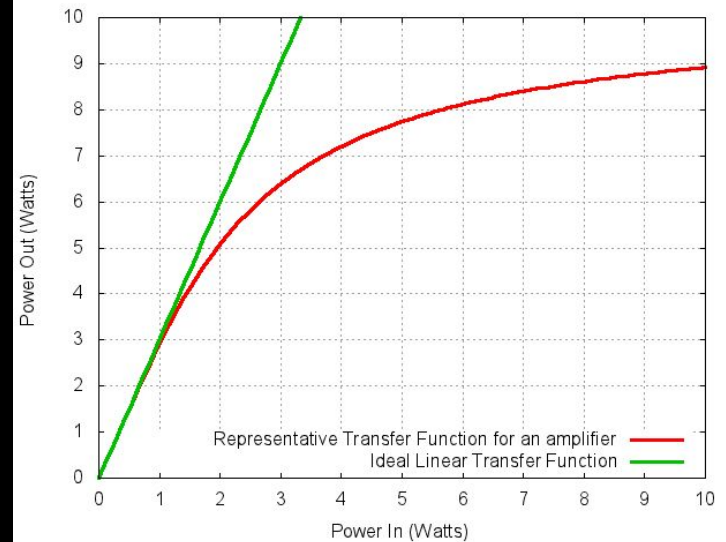
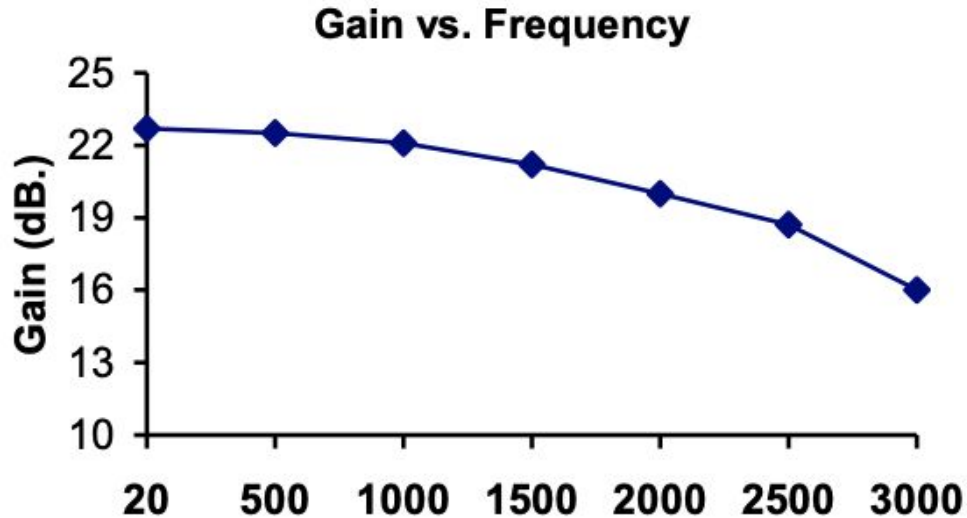
- Primarily work on instrumentation and signal processing
- Do cool stuff with cool equipment in cool places



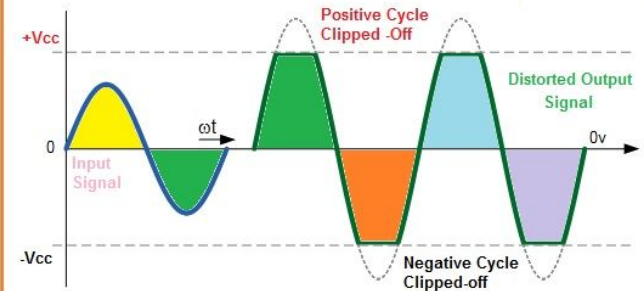
Analyzing science data



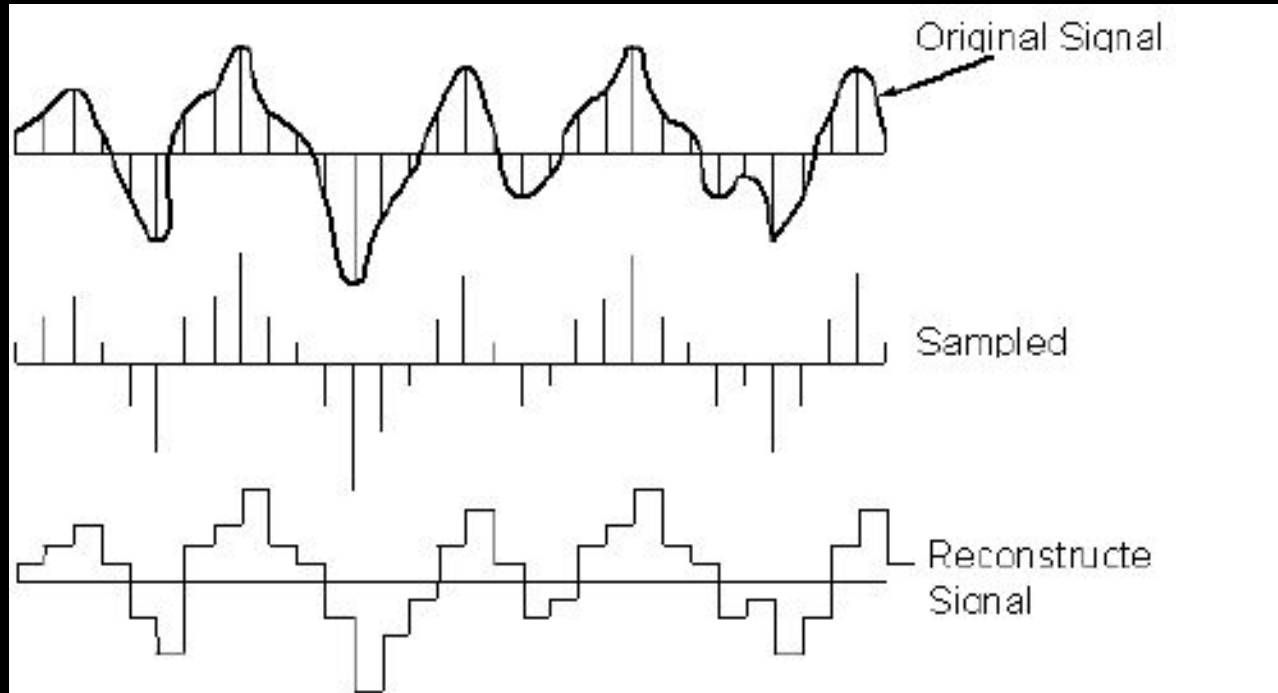
Amplifier response curve



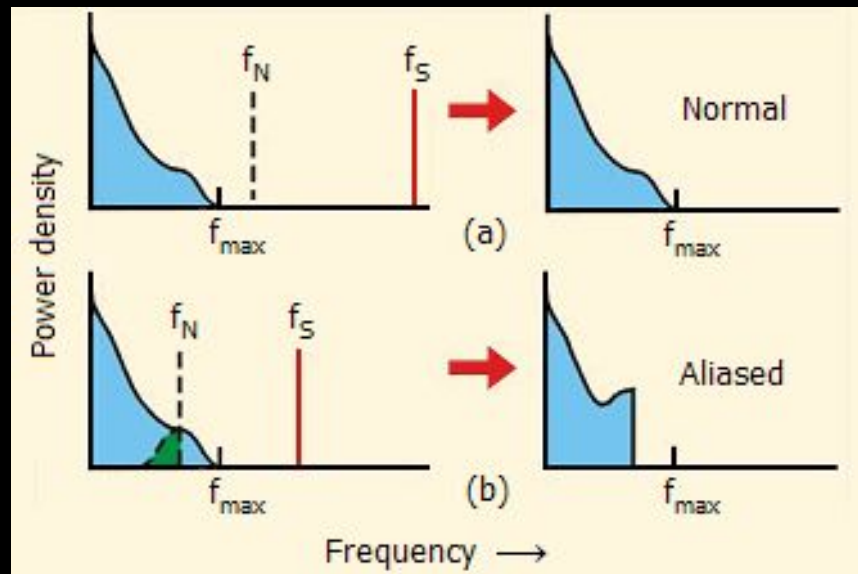
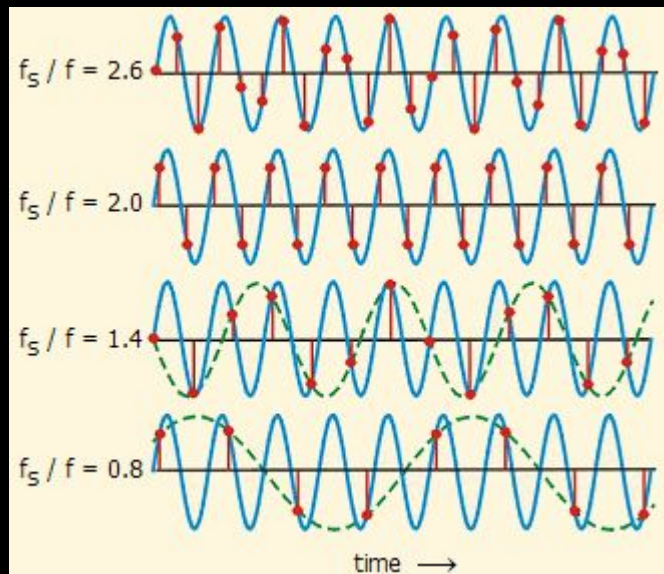
Amplitude Distortion due to Clipping



Sampling Theorem

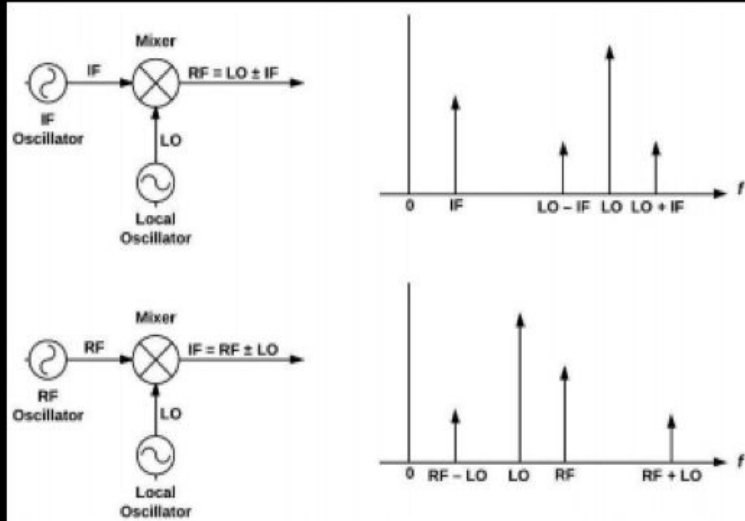


Aliasing

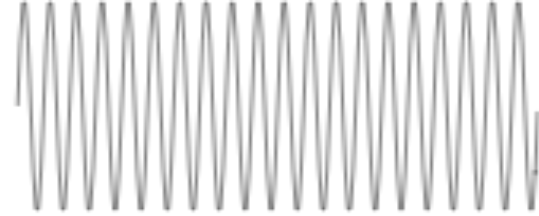


Mixing/Tuning

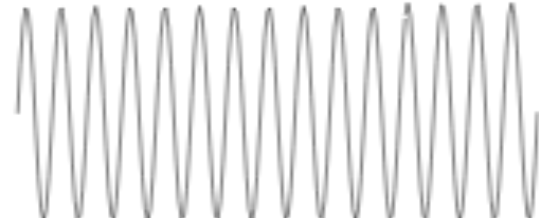
What happens if I want to sample a really high frequency, but low bandwidth?



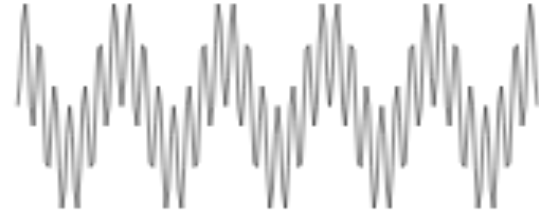
Input 1



Input 2



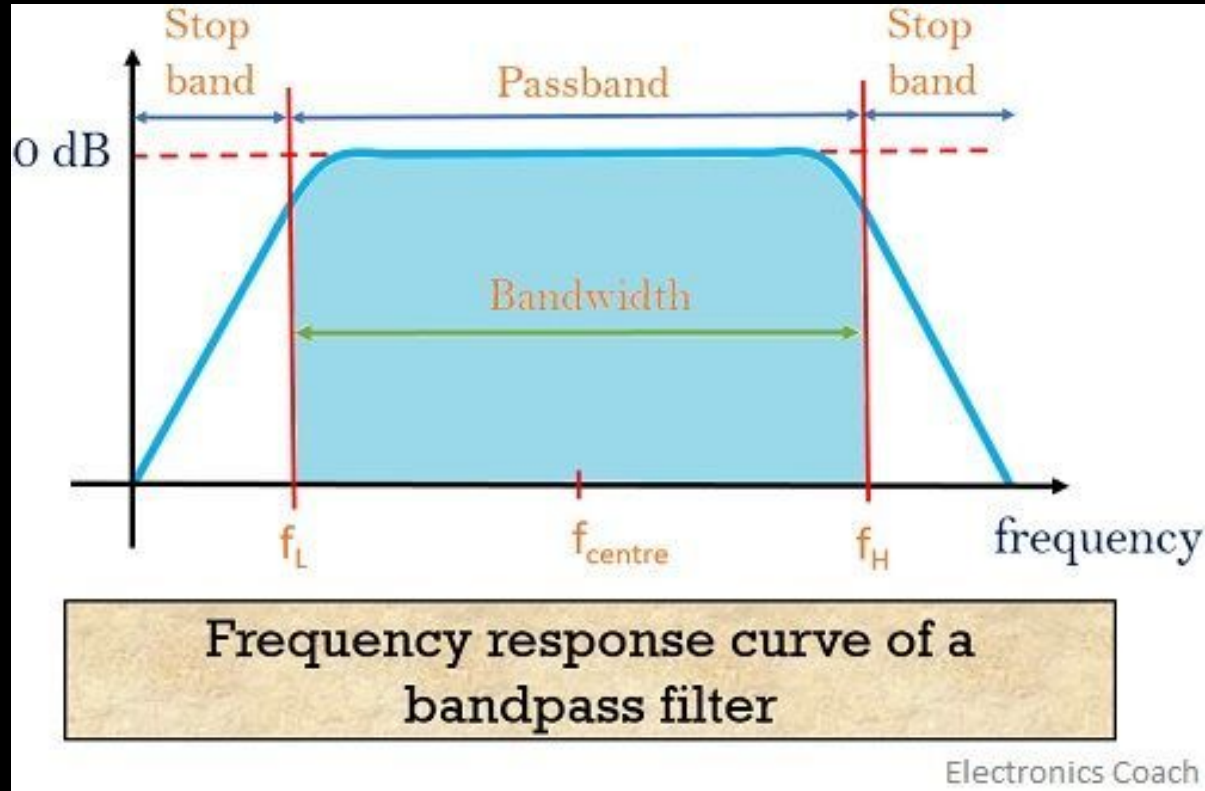
Output



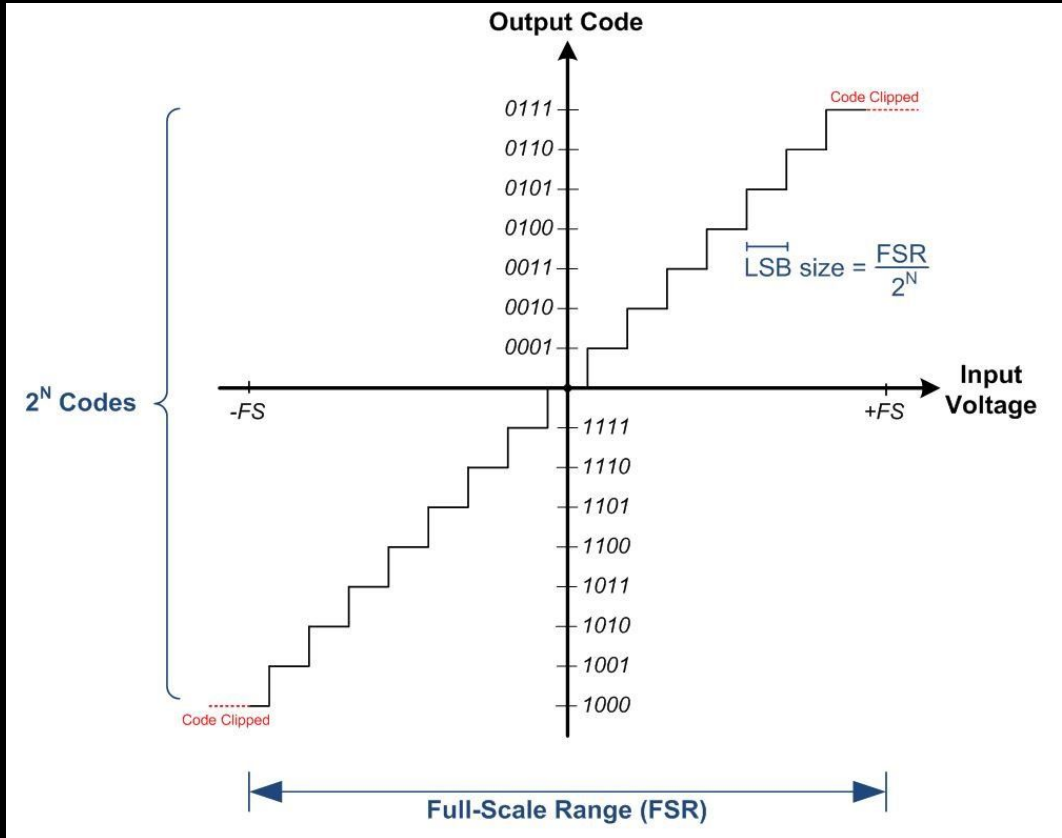
<https://coppermountaintech.com/rf-mixer-characterization/>

<https://academo.org/demos/wave-interference-beat-frequency/>

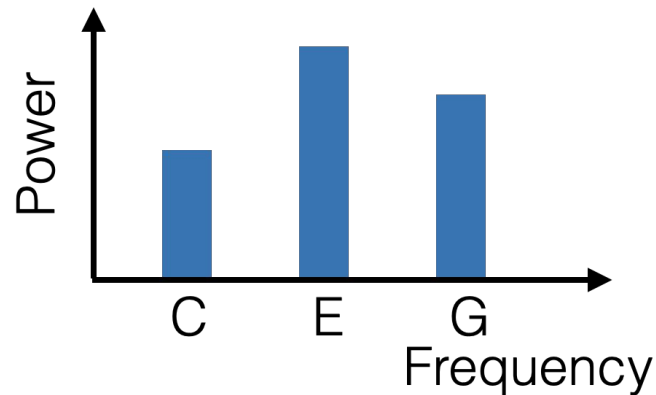
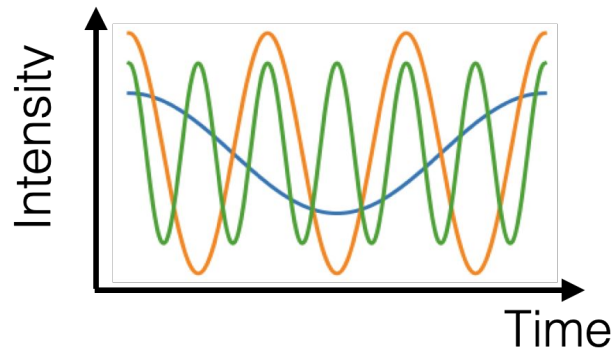
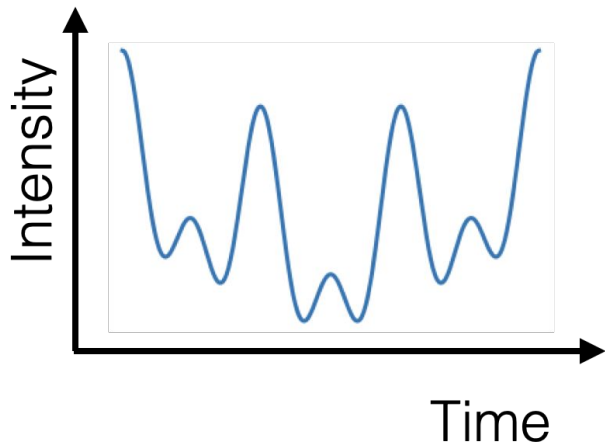
Filtering



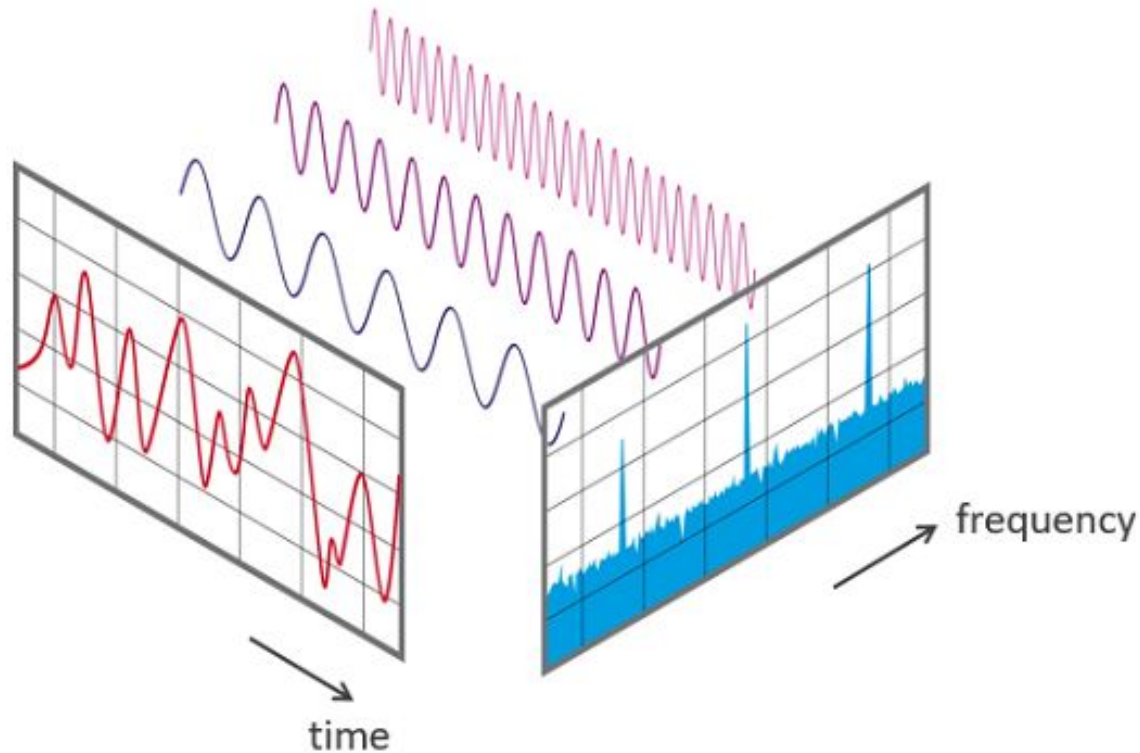
Analog to Digital Conversion



Fourier Transform



Fourier Transform to Spectra



Fourier Series

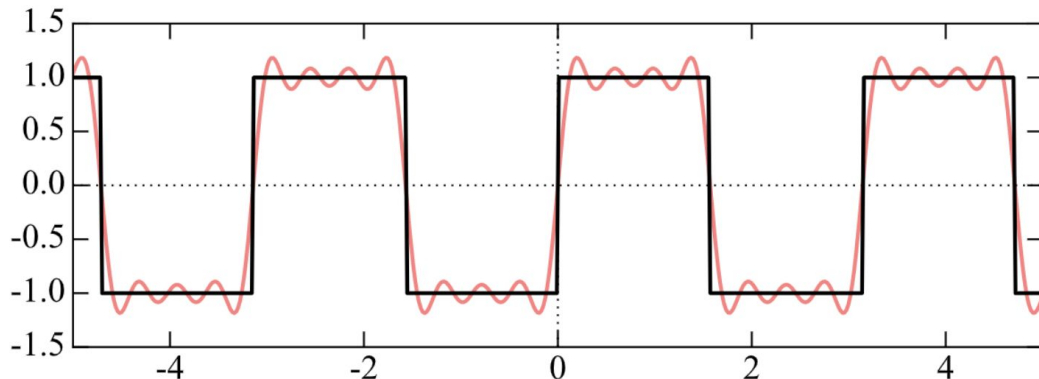


Figure 1: A square wave (black) approximated by a sum of cosines (red).

If you'd like some intuition for the Fourier Series:

https://phet.colorado.edu/sims/html/fourier-making-waves/latest/fourier-making-waves_en.html

Fourier transform

$$\hat{f}(\xi) = \int_{-\infty}^{\infty} f(x) e^{-i2\pi\xi x} dx, \quad \forall \xi \in \mathbb{R}.$$

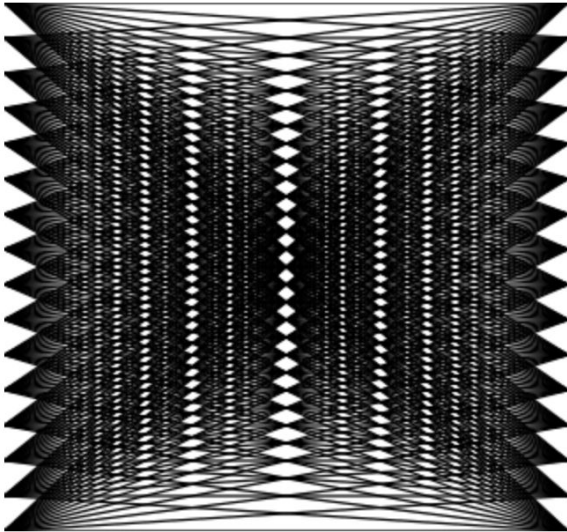
Discrete Fourier transform

$$X_k = \sum_{n=0}^{N-1} x_n \cdot e^{-i2\pi \frac{k}{N} n}$$

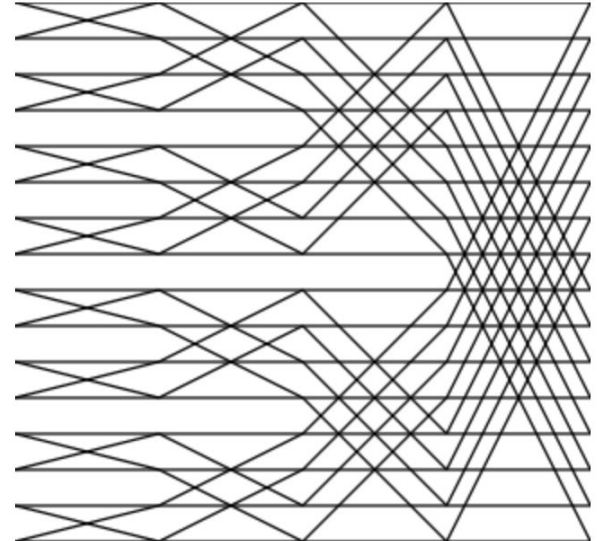
Fast Fourier Transform

- Efficient technique to calculate DFT
- Computations are faster compared to DFT definition
- “Divide and Conquer” → compute DFT of smaller chunks of data and combine them ⇒ lower number of computations
- Most commonly used numerical tool to compute DFT of real signals

DFT, size 16



FFT, size 16



Integration Time

- CHORD will produce 4 + 4 bit samples, at 3.2 GSPS, over 512 dual polarization antennas = 3.3 terabytes/sec
- We cannot store this for extended periods of time, so we either have to:
 - Trigger on events for short timescale capture (i.e, FRBs)
 - Integrate the data down (cosmology)
- We also need to integrate for many science applications to get the required Signal to Noise ratio



SNR: 40 dB



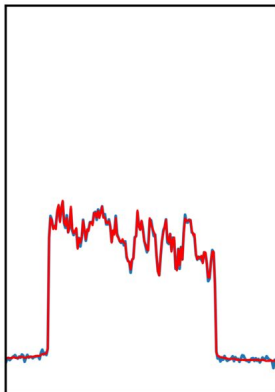
SNR: 20 dB



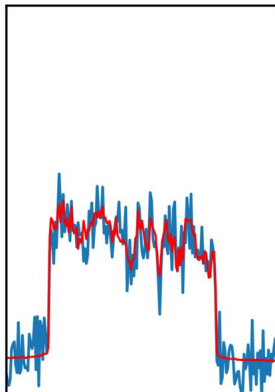
SNR: 10 dB



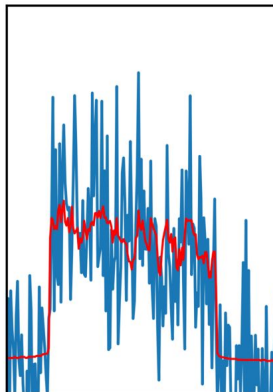
SNR: 0 dB



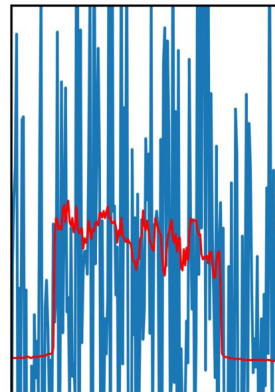
$S/\sigma = 100$



$S/\sigma = 10$



$S/\sigma = 3$



$S/\sigma = 1$

How long do I integrate for?

- Do you want to survey the whole sky, or stare at a single location for a long time?
- Can your telescope point, or does it just survey the sky as it goes overhead?
- What is the required SNR for your science target?
- What is your telescope's field of view?
- Do you need to be able to exclude certain data (i.e, RFI, jackknife testing)
- Etc...
 - For a cosmology telescope that wants to survey a large period of the sky, and does not track (driftscan), often you need hundreds/thousands of hours to meet science goals
 - Sky motion through the beam limits feasible integration to second timescales, and most integration is done offline