

# Shielding/Water Tank

November 13, 2025

Hugh Lippincott

(borrowing liberally from Sally Shaw, Brian Lenardo, Scott Haselschwardt, Gabriel Orebi Gann, among many others)

Also see [Teena John on cosmic veto](#) from yesterday



# Why Shielding/Water Tank



- Background Mitigation

# Why Shielding/Water Tank

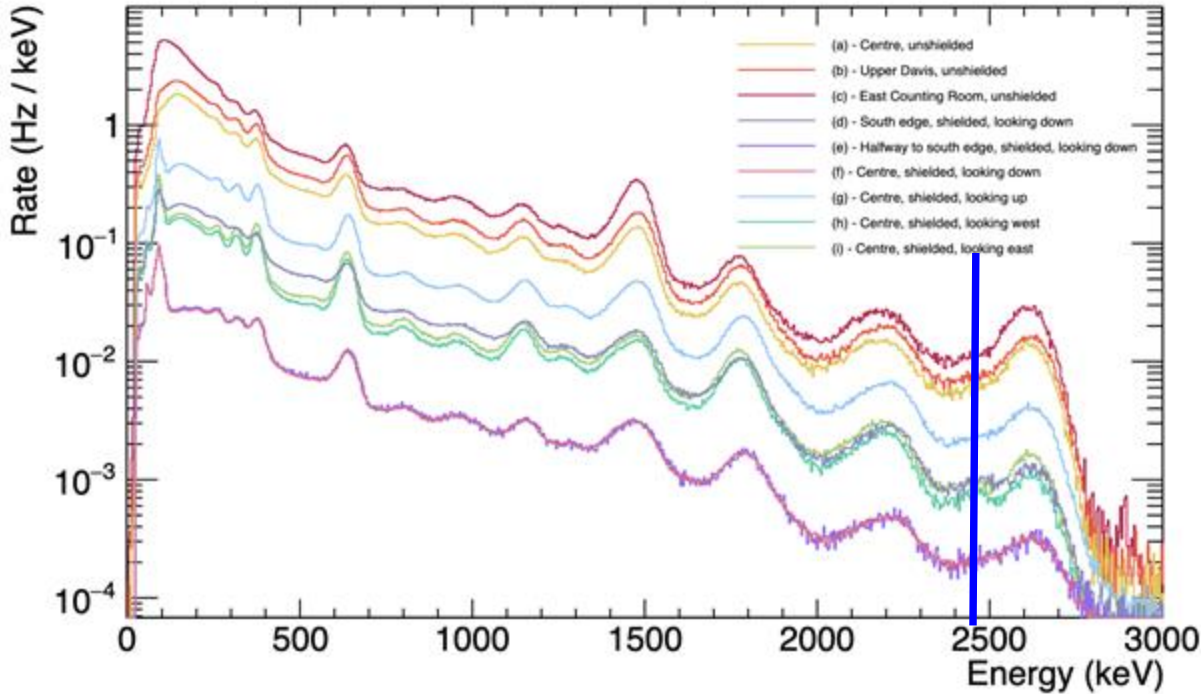


- Background Mitigation
  - a. Shielding - mainly from the lab environment
    - Both gammas and neutrons

# Gamma mitigation



- E.g. Gamma ray backgrounds in Davis Cavern at SURF (measured with 1.6 kg NaI)
- Top line is unshielded
- Bottom line is with 8" of lead shielding
- Radon in air present throughout
- To reach nEXO sensitivity, need something like  $\sim 10^{10}$  reduction at Q value



From [1904.02112](#)

# Neutron mitigation



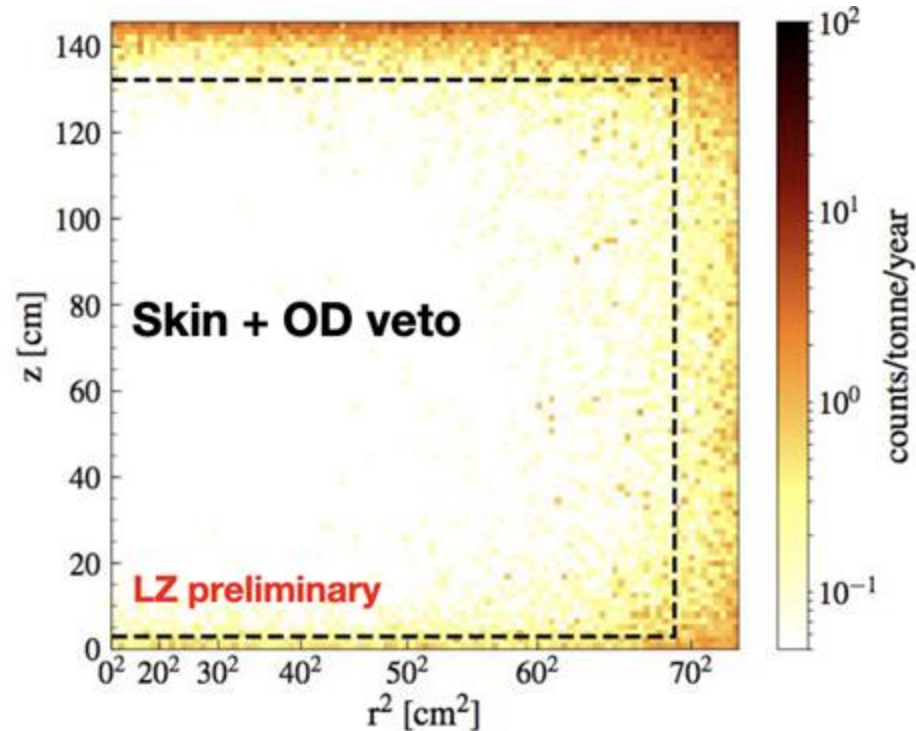
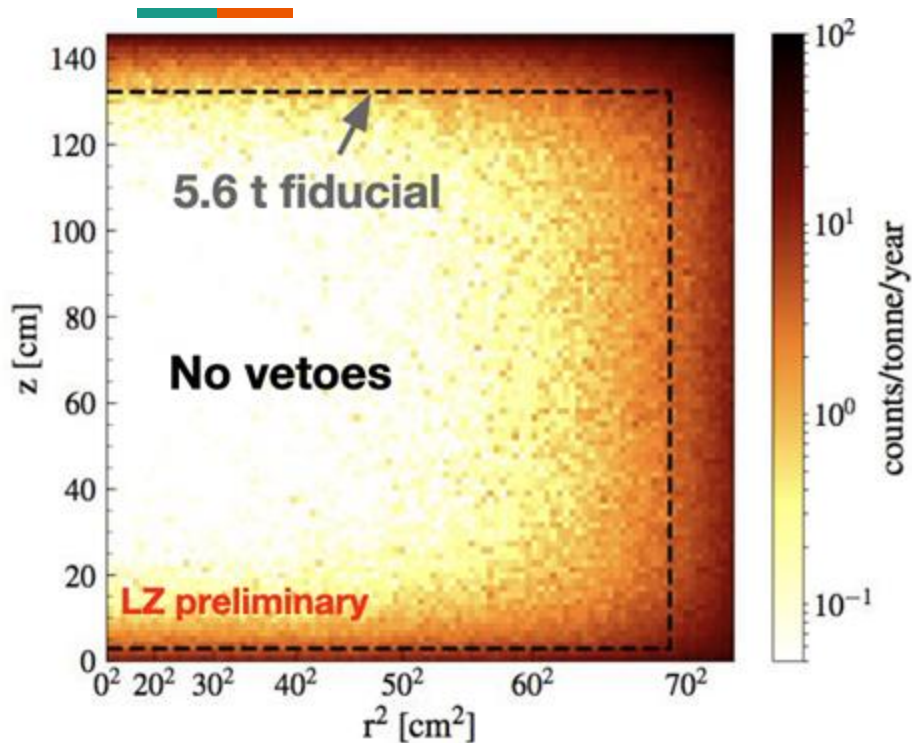
- E.g. Neutron flux in SNOLAB is  $\sim 4000$  fast and  $4000$  thermal neutrons/m<sup>2</sup>/d\*
  - (This is just a lot of neutrons, although 1000 times less than surface)
- For DM experiments, goal is  $< 1$  neutron interaction per exposure.
  - E.g. XENONnT has a surface area of  $10 \text{ m}^2$
  - $\sim 1\text{e}10$  neutrons traversing the detector per year (if it were in SNOLAB)
- Thermal neutrons lead to capture induced gammas
  - E.g.  $^{16}\text{O}(n, \gamma)^{17}\text{O}$  with 3.2 MeV gamma
  - E.g.  $^{16}\text{O}(n, p)^{16}\text{N}$  with 6.17 gamma
  - Iron, nickel, chromium (and Gd) have capture lines in 7-9 MeV range

# Why Shielding/Water Tank



- Background Mitigation
  - a. Shielding - mainly from the lab environment
    - Both gammas and neutrons
  - b. Active veto
    - Tag events that scatter in the central detector

# Active Veto



- Example: LZ expands fiducial volume by 2.3 tonnes (out of 5.5) with active veto of gamma rays and particularly neutrons

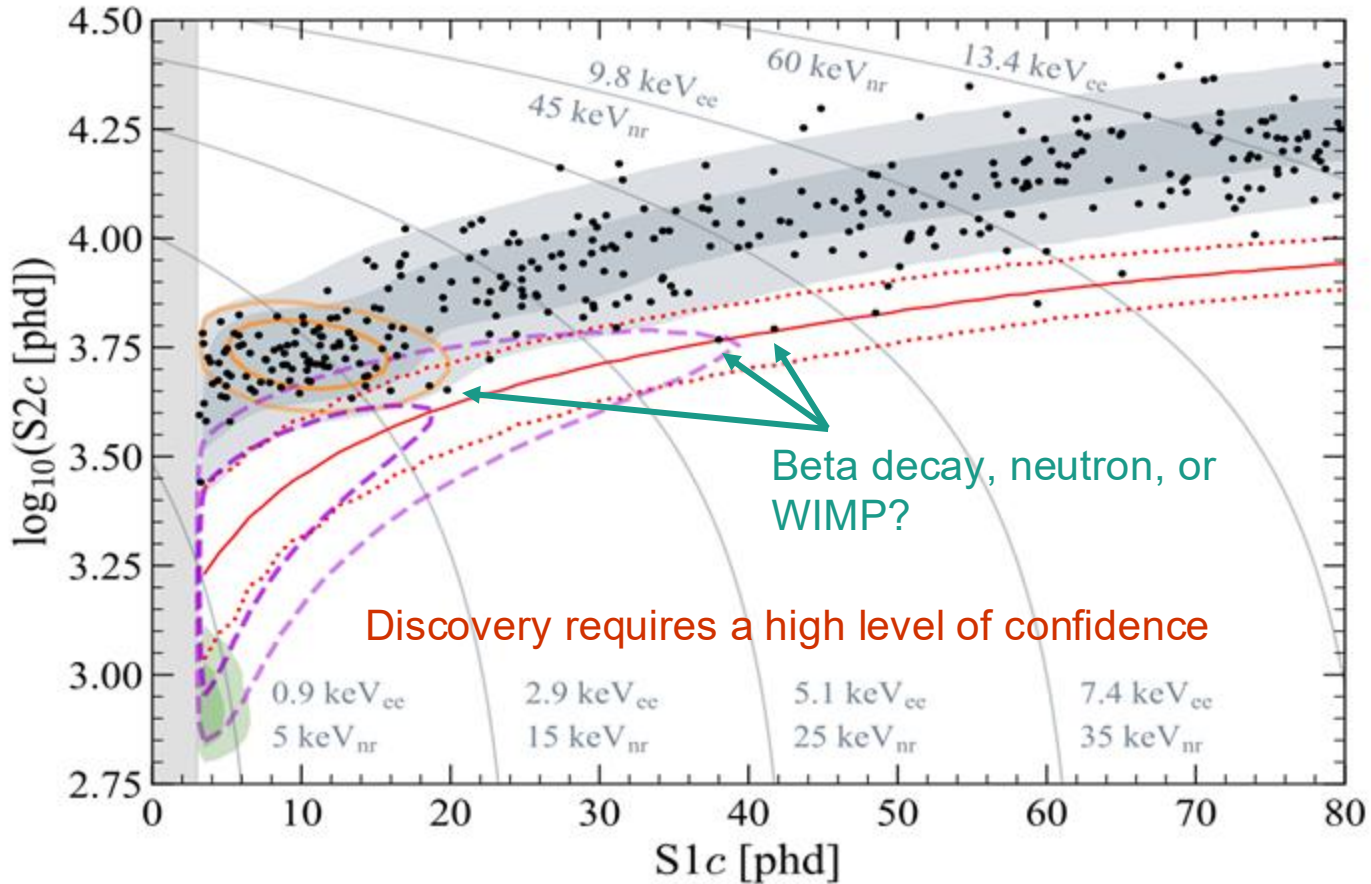
# Why Shielding/Water Tank



- Background Mitigation
  - a. Shielding - mainly from the lab environment
    - Both gammas and neutrons
  - b. Active veto
    - Tag events that scatter in the central detector
  - c. Vibes...



# Vibes...



# Elements of Shielding/Water Tank

- Passive shielding - e.g. Super CDMS ([Jack Nelson, DPF 2021](#))

## Design of Shield

### Outer neutron shield

- 60 cm of water in stainless steel tanks + high density polyethylene base
- Reduces MeV neutrons from cavern wall by  $10^6$

### Gamma shield

- 23 cm lead. Inner layer is cleaner
- Reduces MeV gammas from cavern wall by  $10^5$

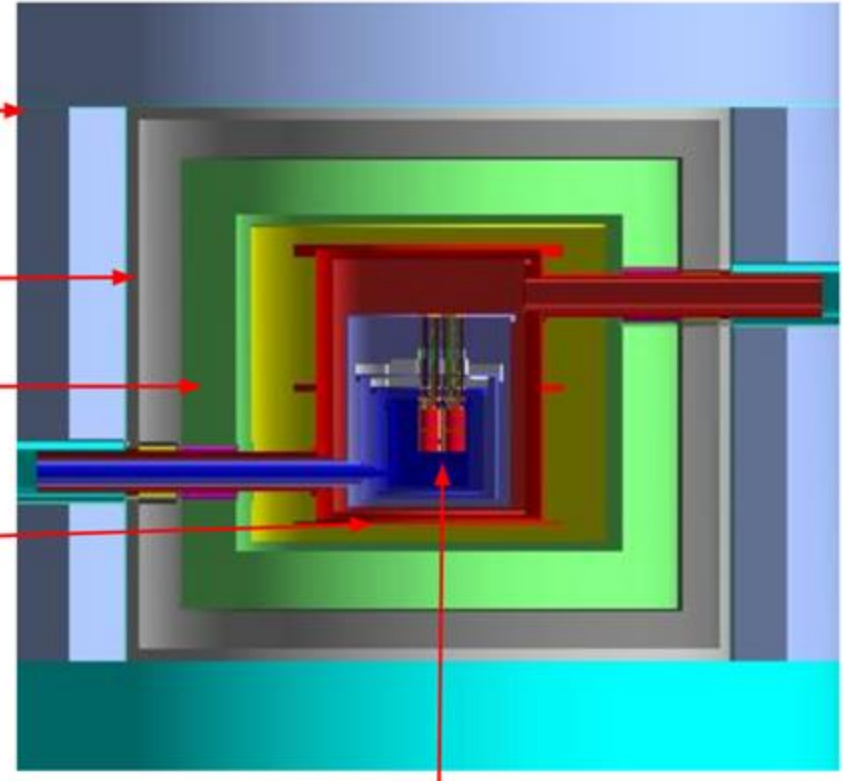
### Inner neutron shield

- 40 cm high density polyethylene
- Reduces GeV neutrons induced by muons in cavern or shield by 100

### Cryostat

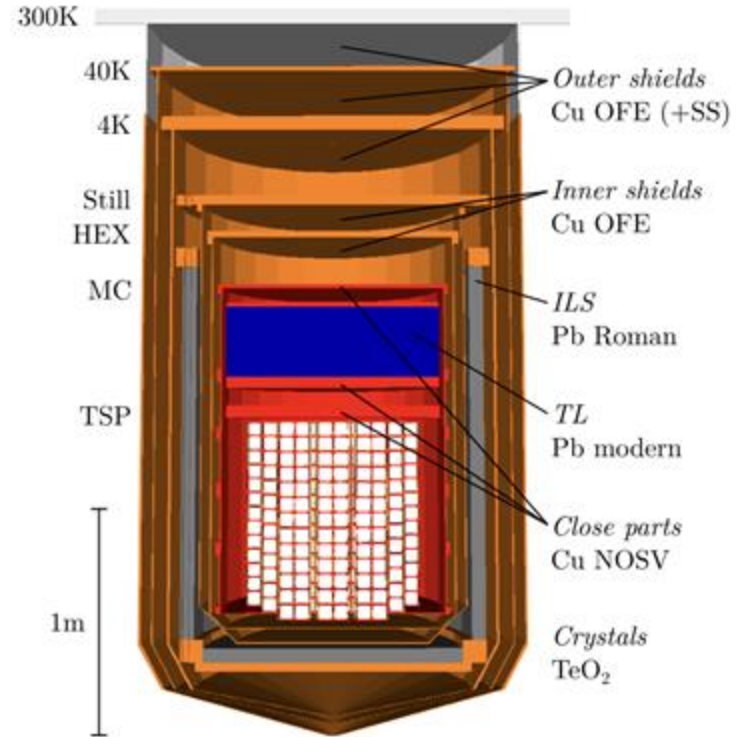
- Shielding is secondary function, but few cm Cu provides modest gamma reduction

Layered design provides complimentary shielding from environmental neutrons and gammas



# Elements of Shielding/Water Tank

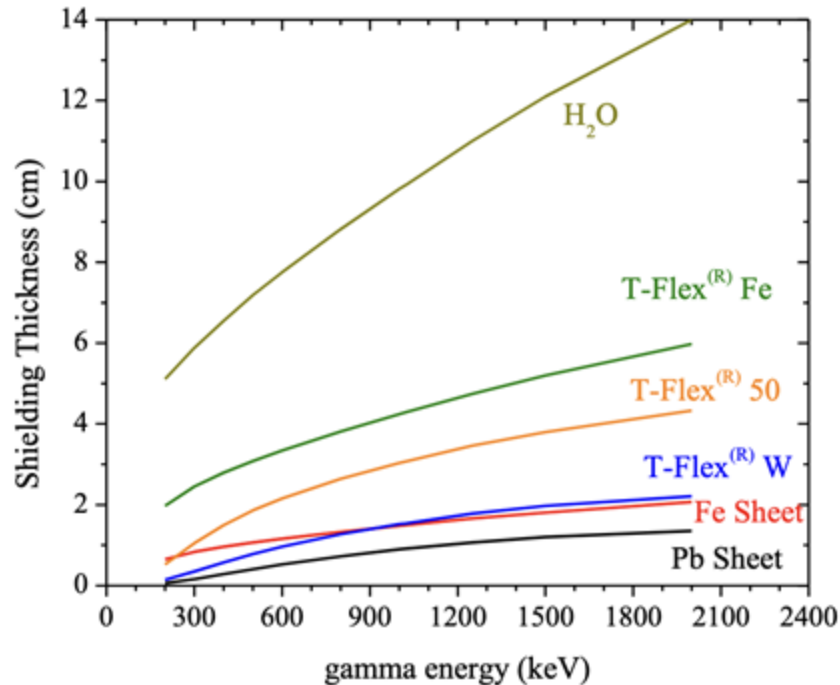
- Passive shielding - e.g. CUORE ([PRD 110, 052003](#))
  - External lead shield and poly+boric acid
  - Copper
  - Lead (modern and ancient
  - Copper again (also for thermal shielding)



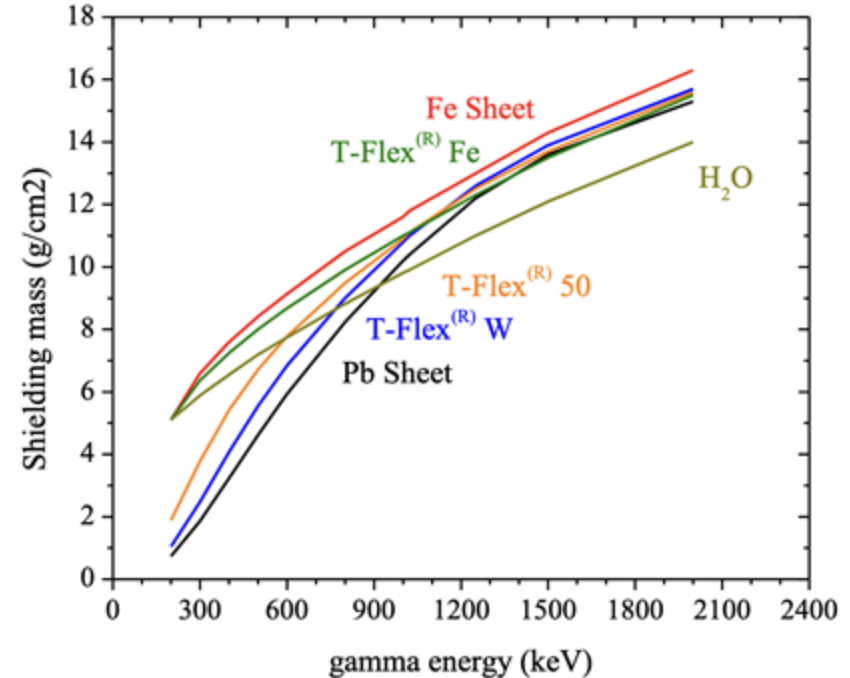
# Elements of Shielding/Water Tank

- Passive shielding - why not just water?

Shielding Thickness (cm) Required to Achieve 50% Attenuation



Mass of Shielding (g/cm<sup>2</sup>) Required to Achieve 50% Attenuation



# Elements of Shielding/Water Tank

- If using water, why not instrument it? -> Active shielding
  - E.g. DEAP-3600 in its water tank
  - Note the outward facing PMTs



# Elements of an Active Veto



- Detection medium
  - a. Water (e.g. many)
  - b. Water with Gd for thermal neutron capture (e.g. XnT)
  - c. Water based liquid scintillator
  - d. Liquid scintillator or LS
  - e. Gd-LS (e.g. LZ)
  - f. Liquid xenon (“skin”)
  - g. Liquid argon (e.g. LEGEND, DarkSide)
  - h. ...

# Elements of an Active Veto



- Containment
  - a. Steel
  - b. Steel with tyvek or other reflector
  - c. Acrylic (for LS)
  - d. ...
- Light collection
  - a. PMTs most common
  - b. SiPMs (e.g. DS-20k)
  - c. Wavelength shifting fibers to SiPMs (e.g. LAr in LEGEND)
  - d. ...

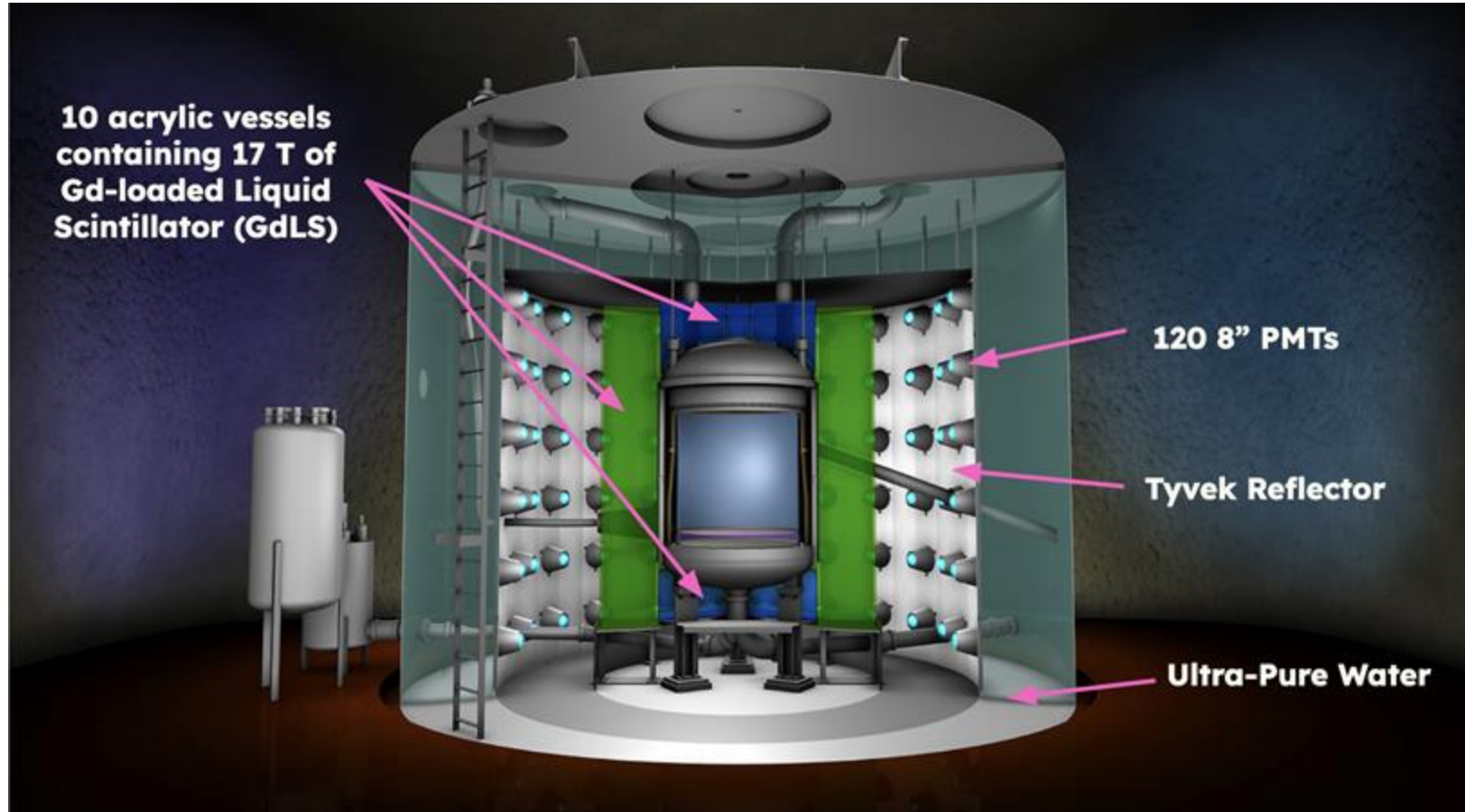
# Considerations



- Size
  - a. Factor  $1e10$  reduction in  $\sim 2.5$  MeV gammas takes  $\sim 5$  meters of water in all directions
  - b. Immediately drives a  $\sim 12 \times 12$  m water tank (need space for the actual detector)
- Ease of fabrication
  - a. Big steel tanks are relatively straightforward
  - b. Acrylic tanks less so (especially if they are sealed)
- Medium vs light collection
  - a. Trade light yield (e.g. LS vs. water) for additional light sensors
  - b. How much to instrument?
  - c. Reflector?

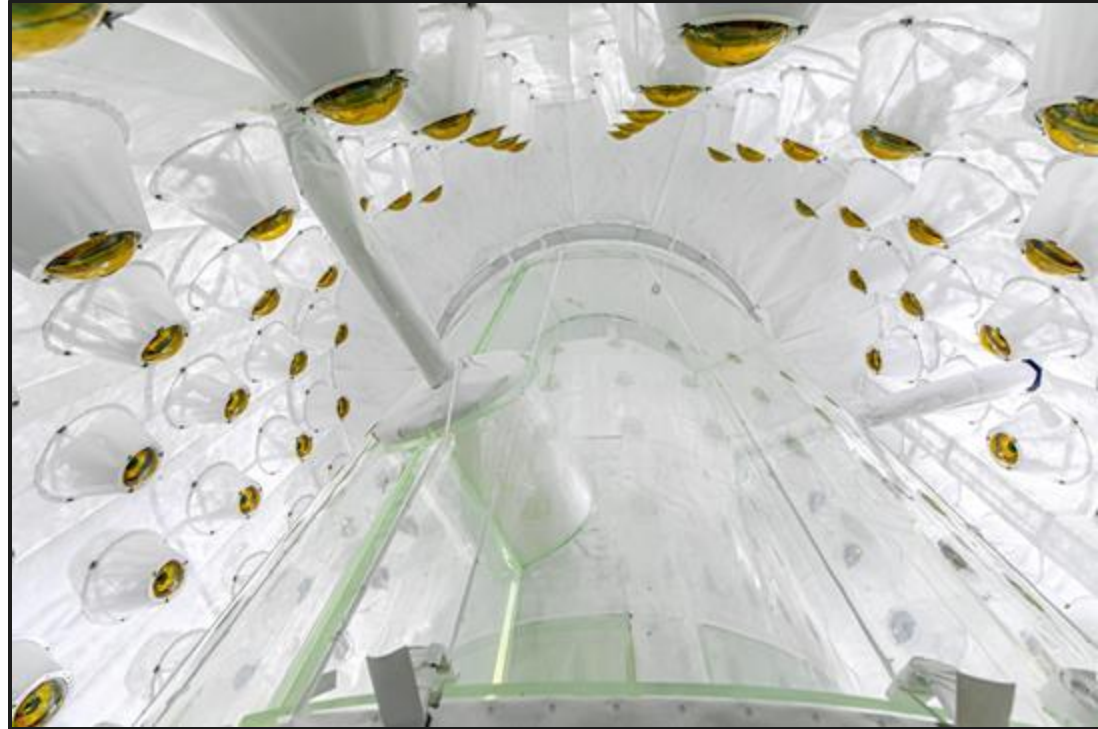
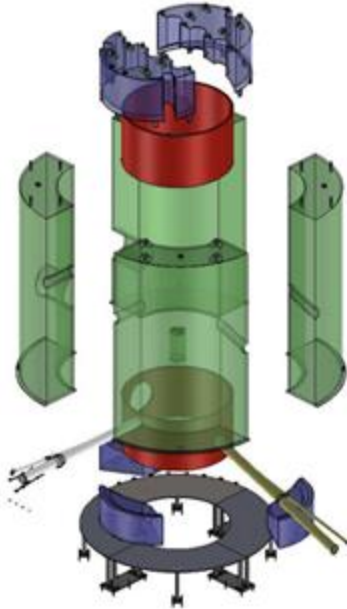


# Some designs - LZ



# Some designs - LZ

- LZ water tank is a bit small (7.6 m x 5.9 m with 238 tonnes of water)
- 10 acrylic vessels that fit together around the detector
  - a. Driven by cost considerations



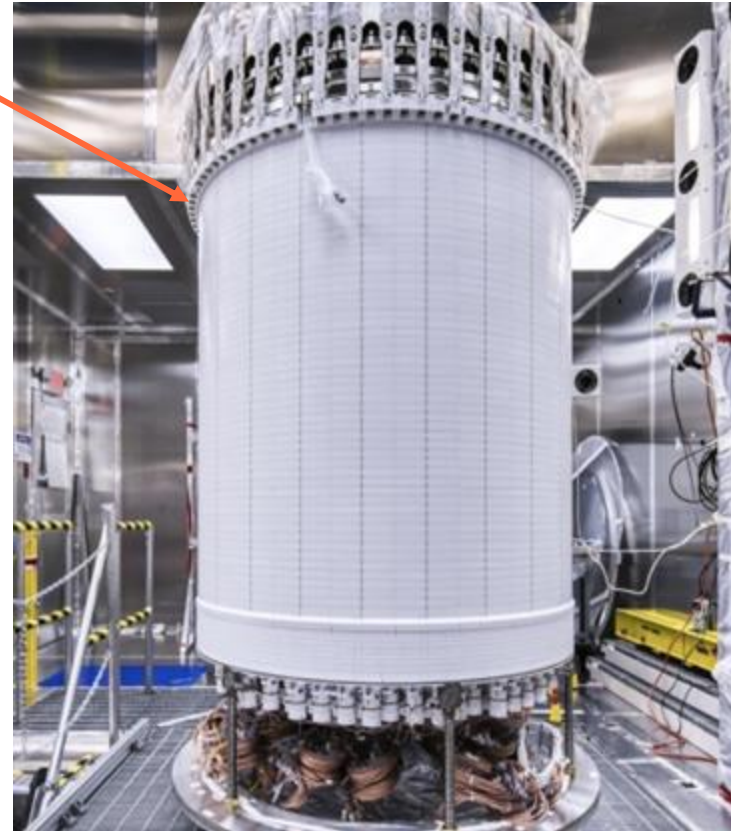
# Some designs - LZ

- 2 tonnes LXe skin

93 1" PMTs

PTFE reflector

38 2" PMTs

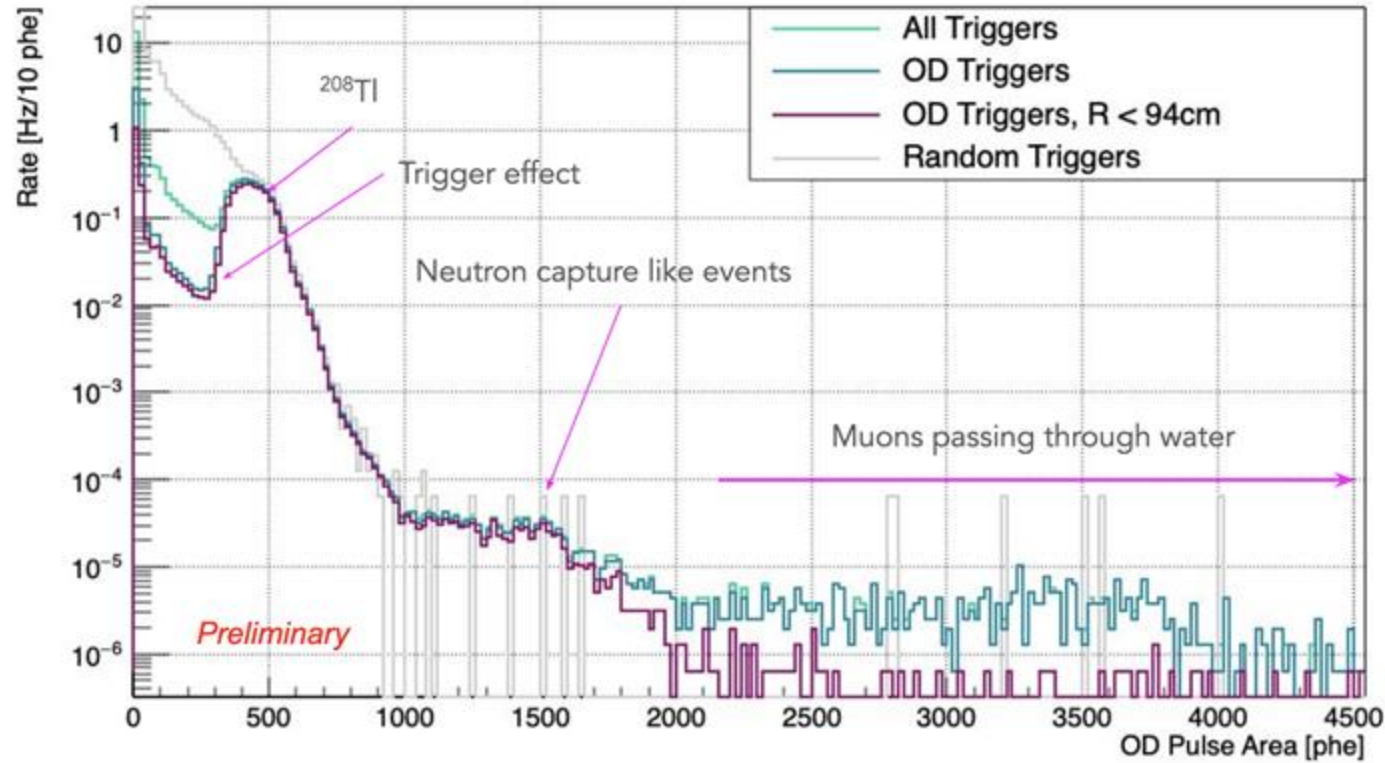


# Some designs - LZ



- Low veto threshold required ( $\sim 150$ - $200$  keV in OD, 100 keV in skin)
  - Maximize veto efficiency - the vessels are too thin to contain 2.2 MeV gammas let alone 8.5 MeV endpoint of Gd capture
  - In turn requires high light yield ( $>80$  pe/MeV)
- Minimize OD background rate ( $<100$  Hz) to minimize livetime loss in WIMP search
  - Accidental veto in OD from backgrounds is deadtime
  - Drives standoff from steel tank to Tyvek curtain
  - Drives GdLS radiopurity requirement
  - Drives PMT/support structure radiopurity requirement

# Some designs - LZ



Higher than expected (x100) n capture rate

Partly (at least 50%) due to cavern neutrons capturing on water tank walls

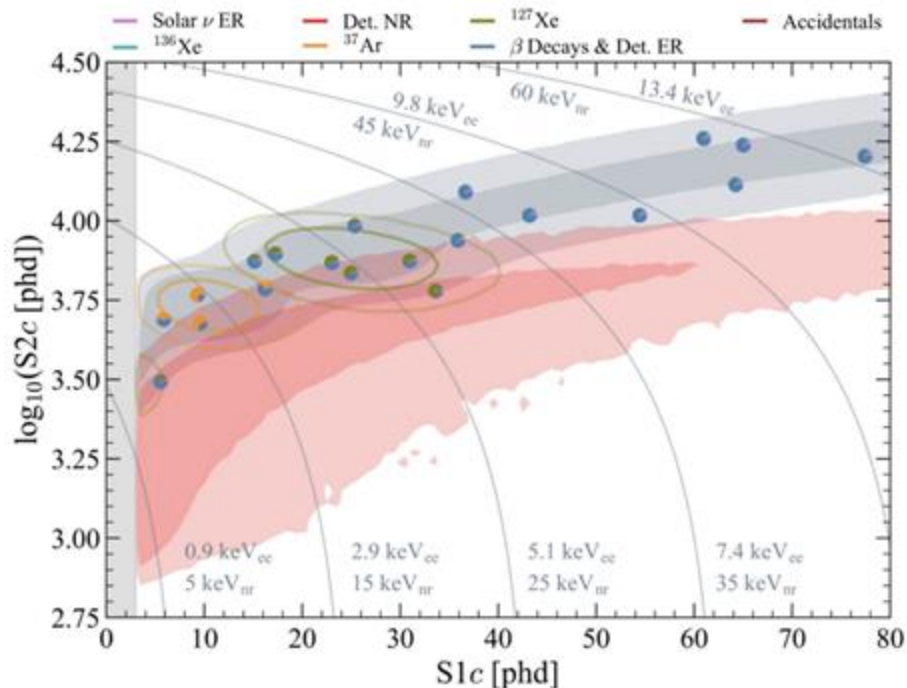


# Some designs - LZ

- OD+skin used to directly constrain neutron rate as a sideband in WIMP search
- Measured ~92% neutron tagging efficiency with 3% dead time

Source	Pre-fit Expectation	Fit Result
$^{214}\text{Pb}$ $\beta$ s	$743 \pm 88$	$733 \pm 34$
$^{85}\text{Kr} + ^{39}\text{Ar}$ $\beta$ s + det. $\gamma$ s	$162 \pm 22$	$161 \pm 21$
Solar $\nu$ ER	$102 \pm 6$	$102 \pm 6$
$^{212}\text{Pb} + ^{218}\text{Po}$ $\beta$ s	$62.7 \pm 7.5$	$63.7 \pm 7.4$
Tritium+ $^{14}\text{C}$ $\beta$ s	$58.3 \pm 3.3$	$59.7 \pm 3.3$
$^{136}\text{Xe}$ $2\nu\beta\beta$	$55.6 \pm 8.3$	$55.9 \pm 8.2$
$^{124}\text{Xe}$ DEC	$19.4 \pm 2.5$	$20.4 \pm 2.4$
$^{127}\text{Xe} + ^{125}\text{Xe}$ EC	$3.2 \pm 0.6$	$2.7 \pm 0.6$
Accidental coincidences	$2.8 \pm 0.6$	$2.6 \pm 0.6$
Atm. $\nu$ NR	$0.12 \pm 0.02$	$0.12 \pm 0.02$
$^8\text{B} + \text{hep}$ $\nu$ NR	$0.06 \pm 0.01$	$0.06 \pm 0.01$
Detector neutrons	$0.0^{+0.2}$	$0.0^{+0.2}$
40 GeV/c <sup>2</sup> WIMP	–	$0.0^{+0.6}$
Total	$1210 \pm 91$	$1202 \pm 41$

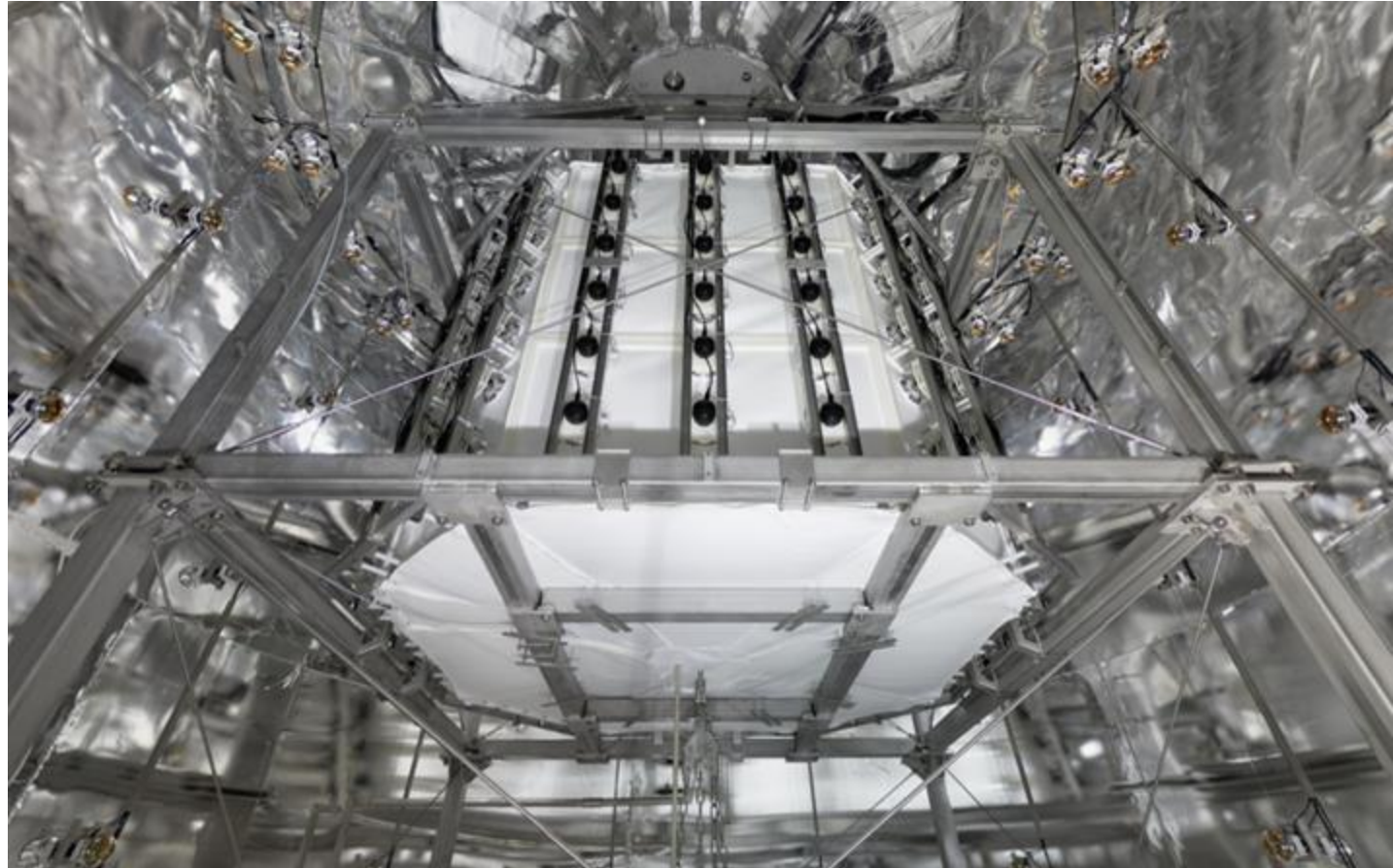
<sup>a</sup> The expected number of neutron events results from a fit to the sample of veto detector-tagged events. This expectation is not explicitly used in the final combined fit as this sample is included directly in the likelihood, as described in the text.



# Some designs - XENONnT

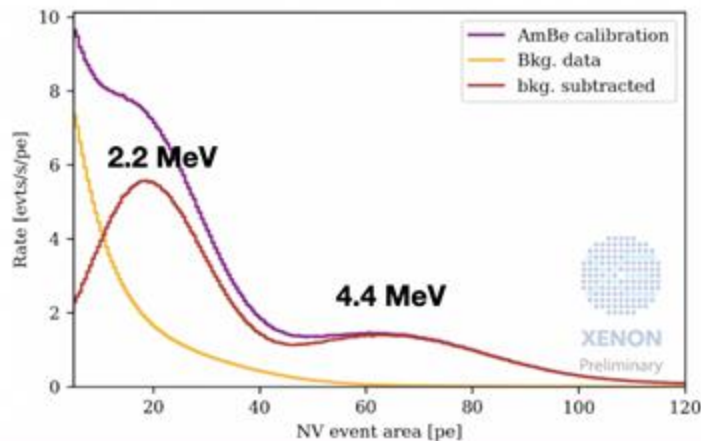


- 9.6 x 10.2 m (700 tonnes total)
- 33 tonne inner water region around cryostat separated by PTFE
- Now loaded with Gd



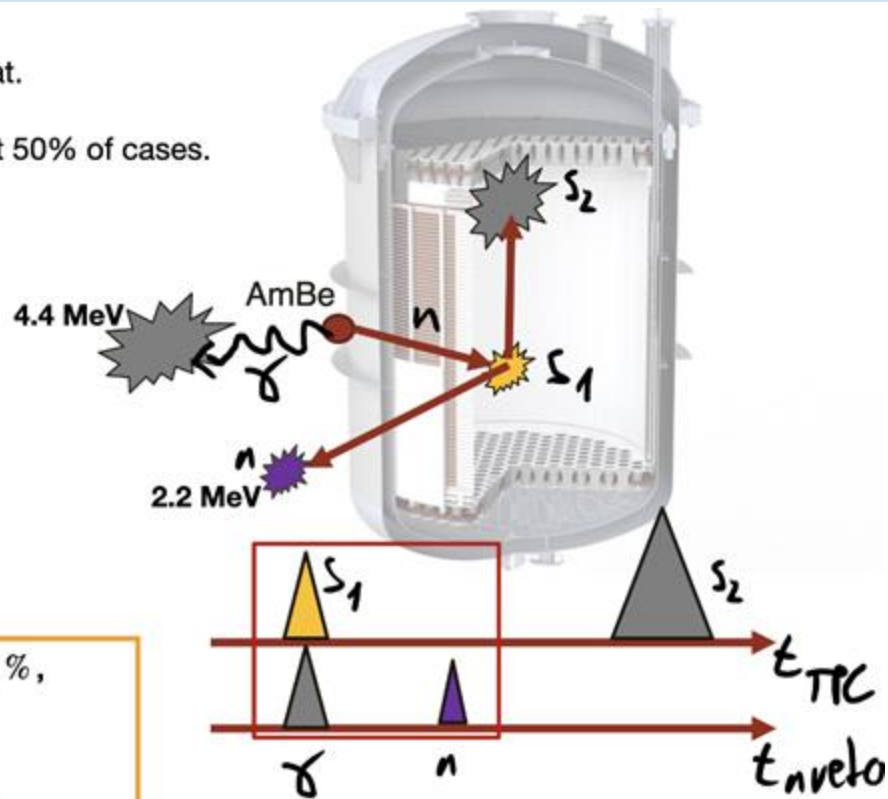
# Some designs - XENONnT

- Neutron calibration with AmBe source placed close to the cryostat.
- AmBe emits a 4.4 MeV gamma together with the neutron in about 50% of cases.



The neutron tagging efficiency in a 250 us window is  $(53 \pm 3) \%$ ,  
with a live-time loss is 1.6%  
(the best ever obtained in a water Cherenkov detector)

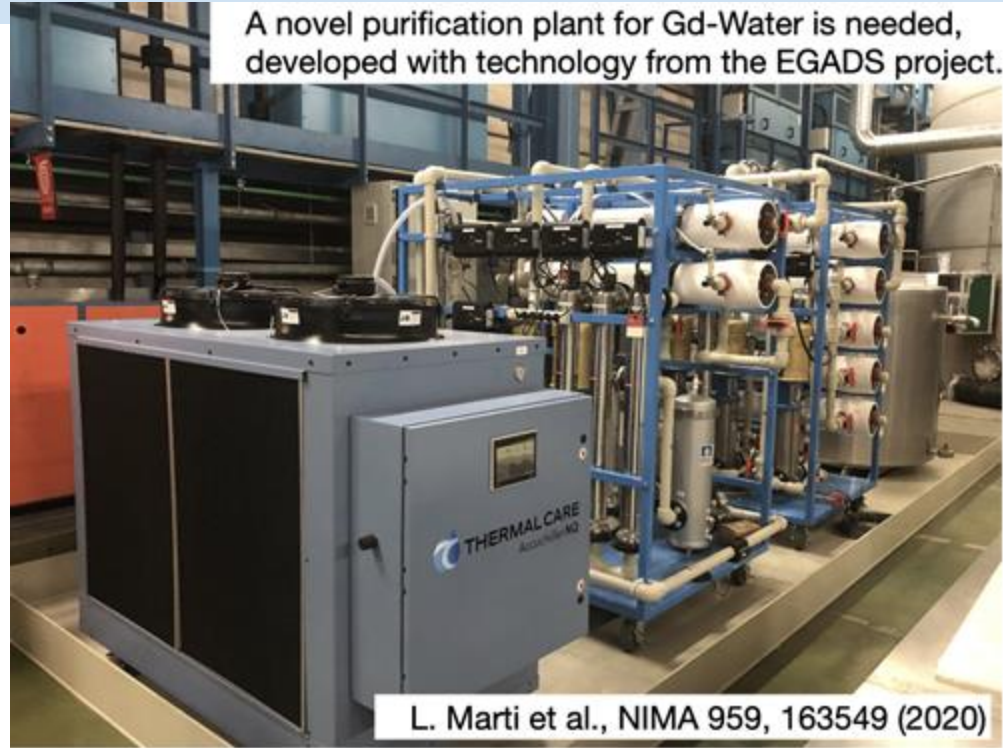
Paper submitted to EPJ-C, <https://arxiv.org/abs/2412.0526>





# Some designs - XENONnT

- Add Gd-Sulphate-Octahydrate salt at 0.02% Gd concentration
- Improve neutron tag by factor of 2
- See [M. Selvi, DRD2 meeting](#)

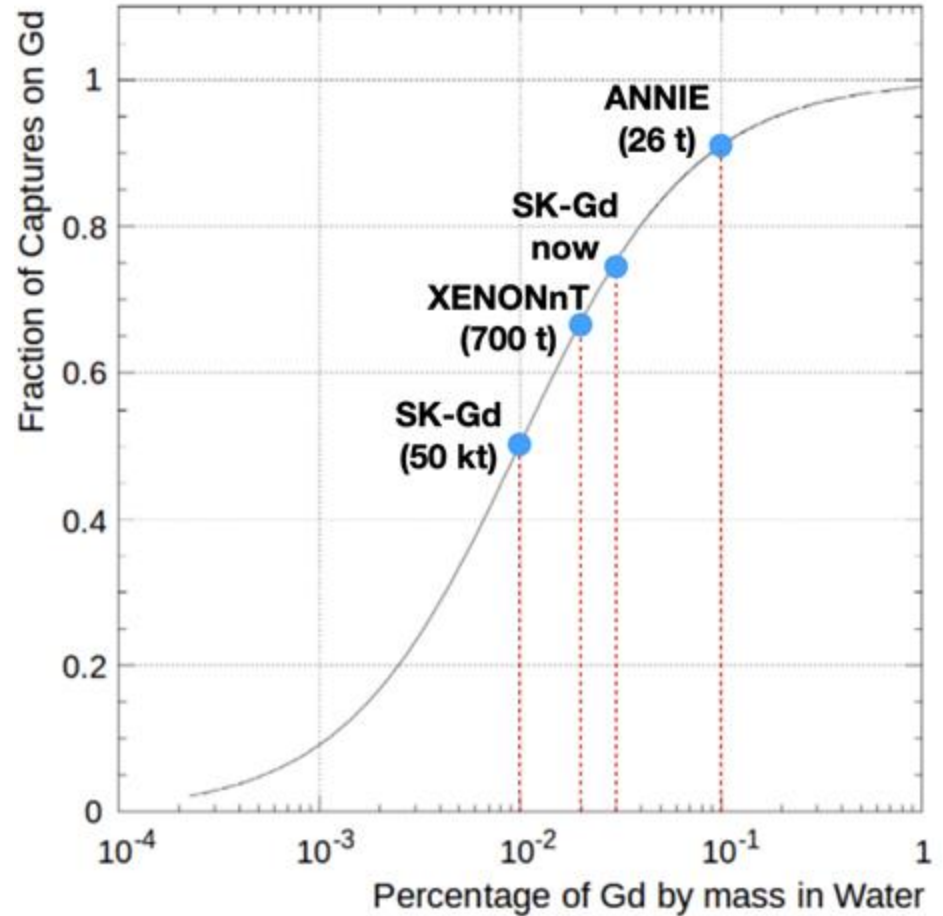
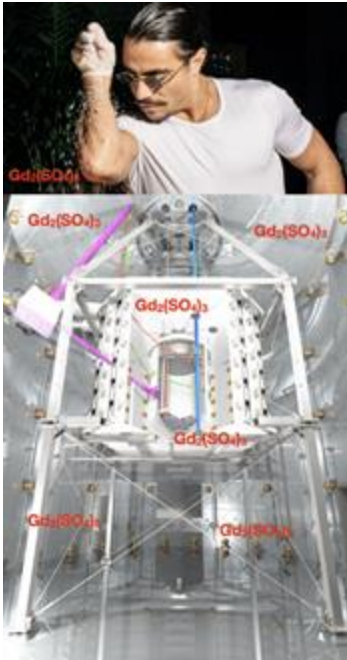


Main strategy:

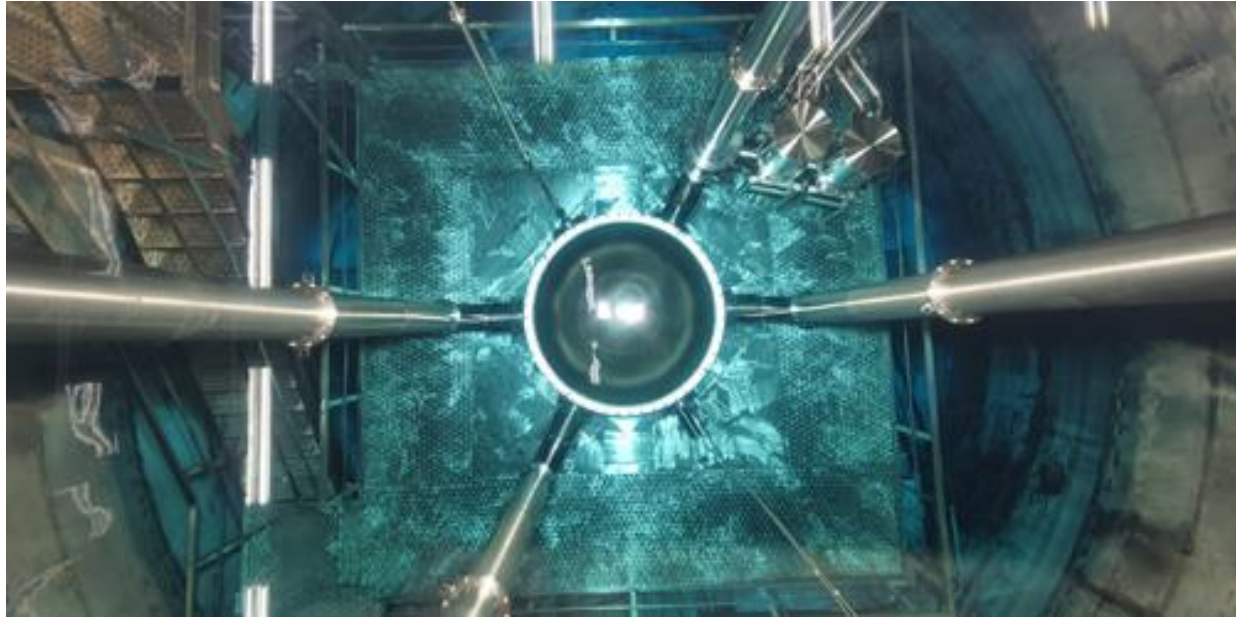


# Some designs - XENONnT

- Goal to increase Gd concentration
- Plans for recovery at end of run



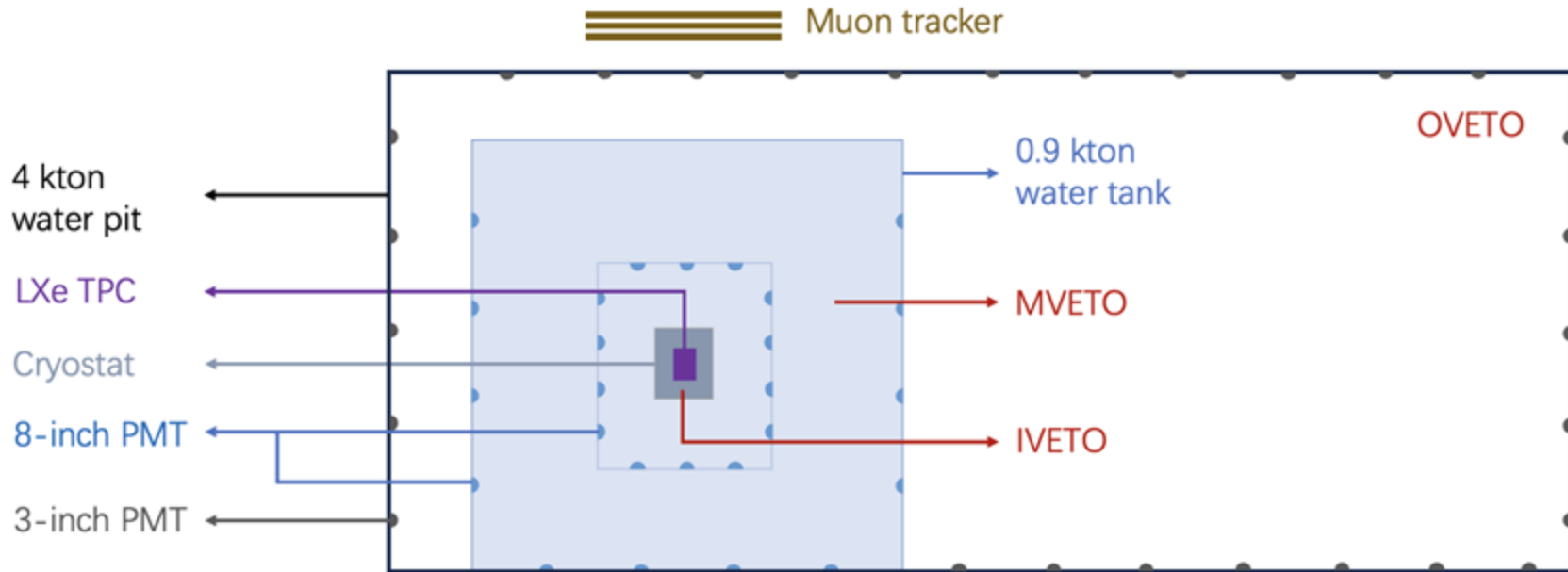
# Some designs - PandaX



- In PandaX-4T, just a very big tank - 10m x 13m with 900 tonnes of water

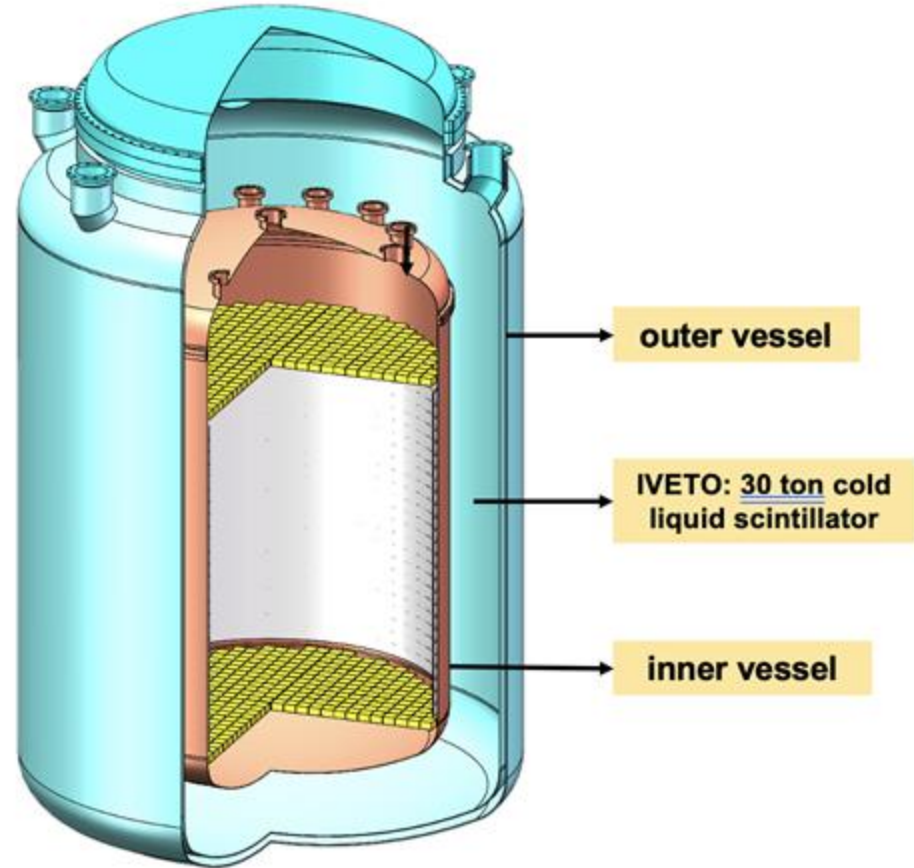
# Some designs - PandaX-xT

- MVETO: 0.9 kton water Cherenkov detector (currently running)
- OVETO: 4 kton water Cherenkov detector (under development)
- IVETO: cold liquid scintillator (R&D in progress)



# Some designs - PandaX-xT

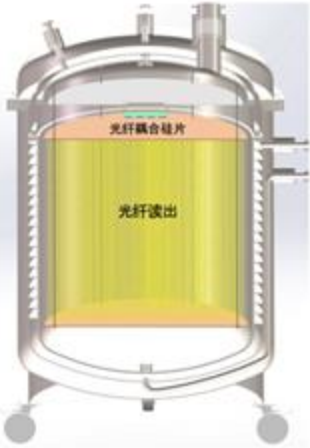
- Cold liquid scintillator between a steel outer vessel and a copper shell
  - a. Similar to nEXO HFE design in terms of cryo and pressure
- Fibers + SiPM for light detection in the cold LS
- Under R&D



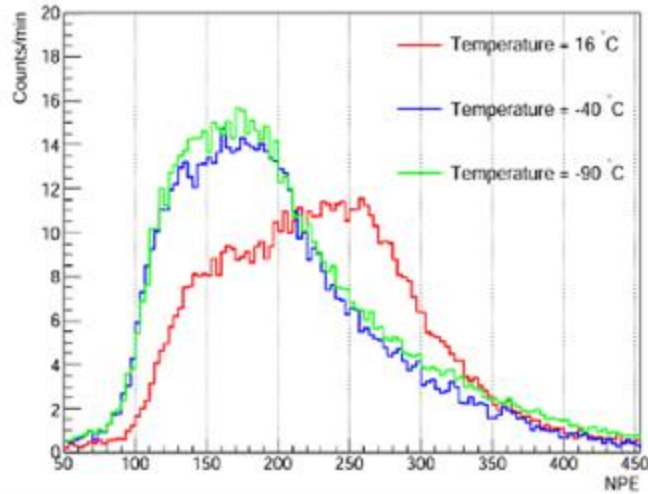


# Some designs - PandaX-xT

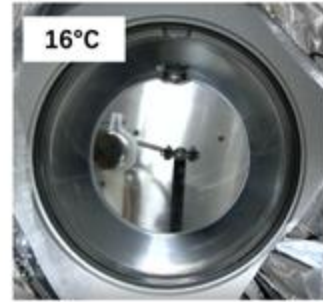
- Light yield drops cold
  - Working on 1 m<sup>3</sup> prototype
- See J. Huang talk at PandaX Open meeting for more - [link](#)



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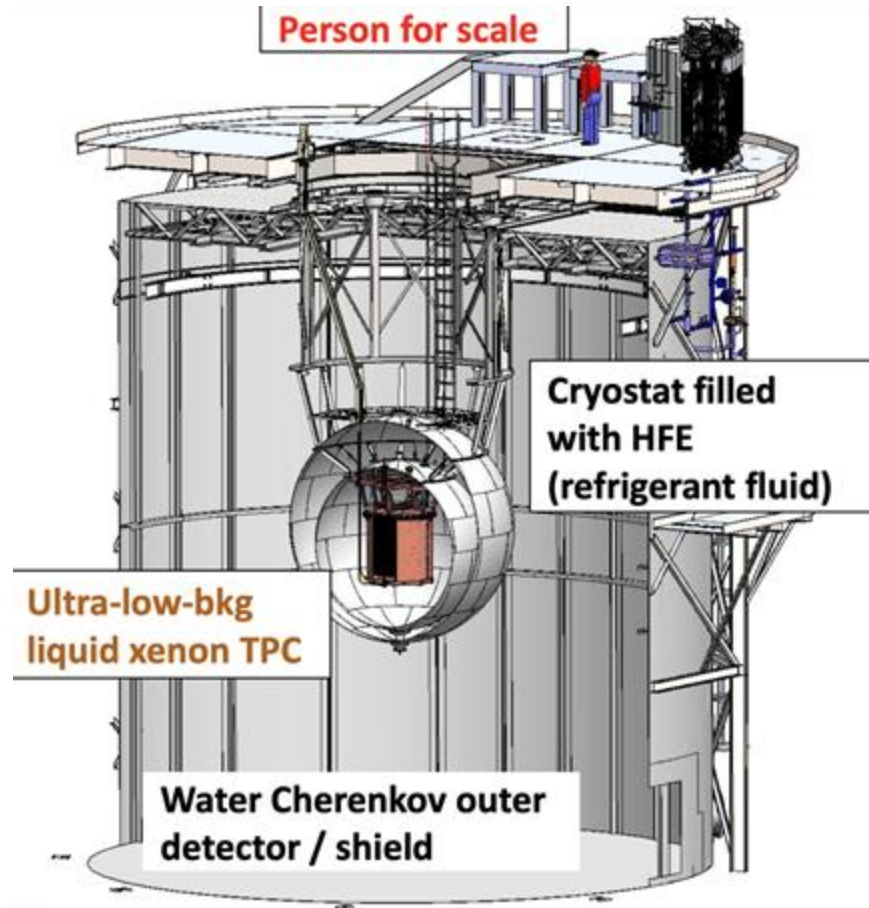


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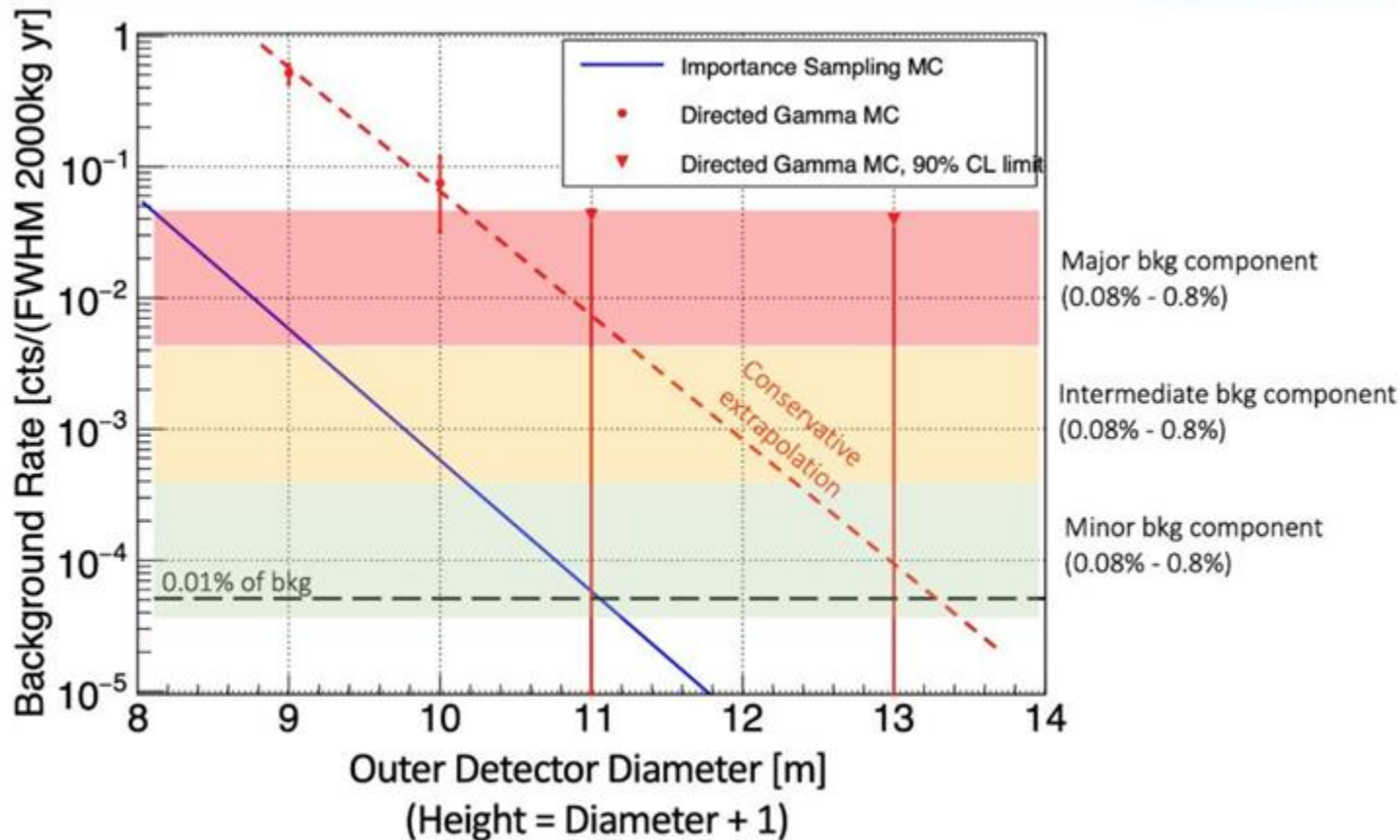


# Some designs - nEXO

- 12x12.8 m water tank
- Passive shielding (from gammas)
  - Reduces cavern gammas to subdominant contribution
- Active for muons
- Tyvek on outside



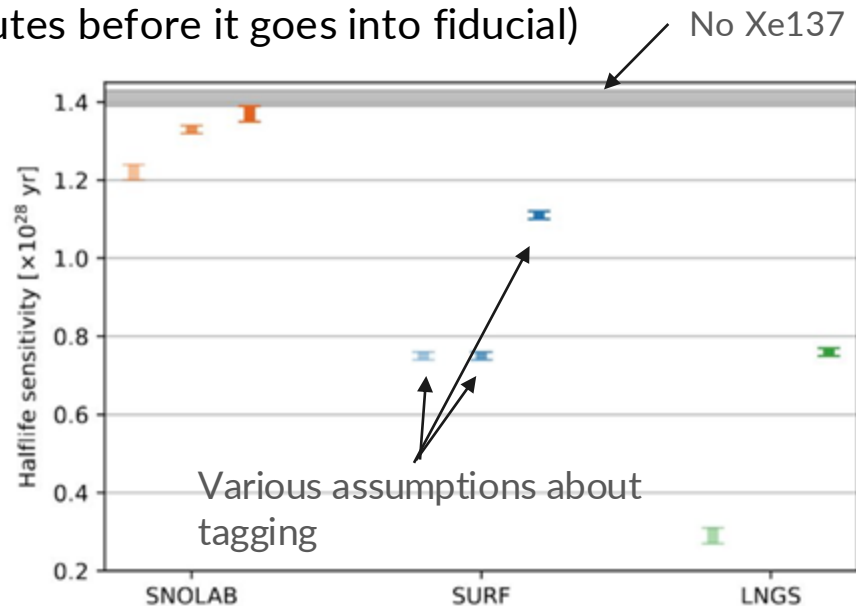
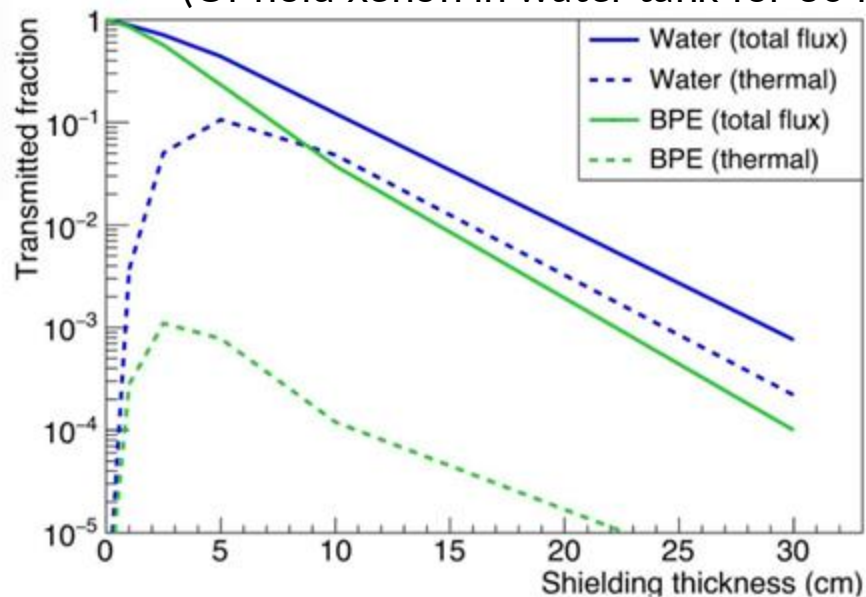
# Some designs - nEXO





# Some designs - nEXO

- Neutrons absorbed in water tank
- Neutron capture on  $^{136}\text{Xe}$  outside tank (site dependent)
  - Shielding required for external Xe in purification system
  - (Or hold xenon in water tank for 30 minutes before it goes into fiducial)



# Some designs - nEXO

- Also includes a skin as part of an optically open TPC design
  - Reduces gamma and radon backgrounds by tagging

(<https://arxiv.org/abs/2009.10231>)

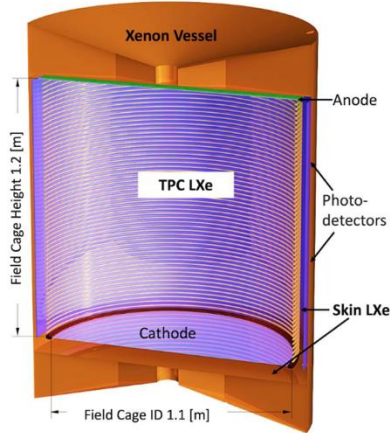


Figure 1: Cross-section of the proposed TPC for nEXO.

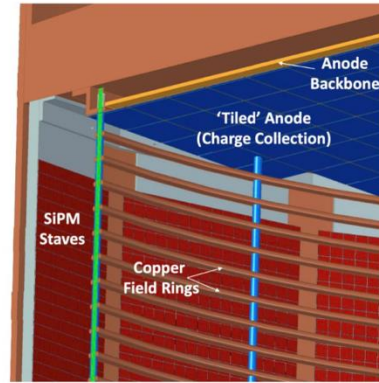
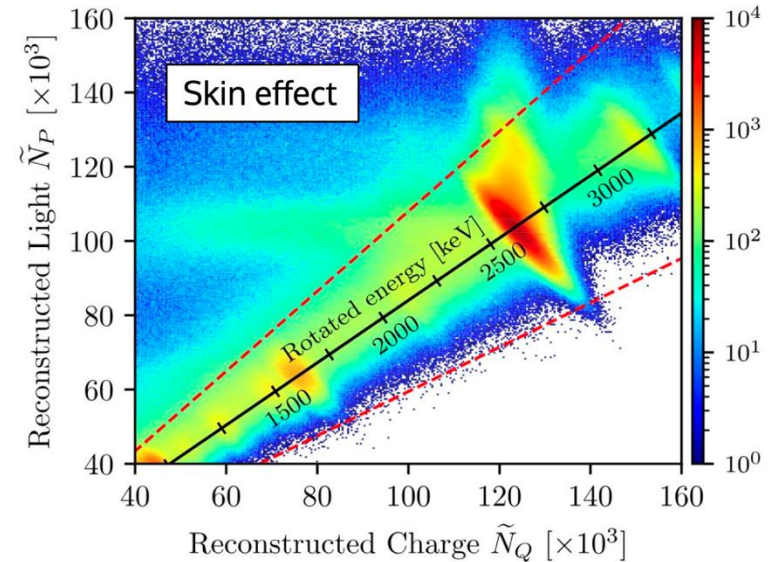
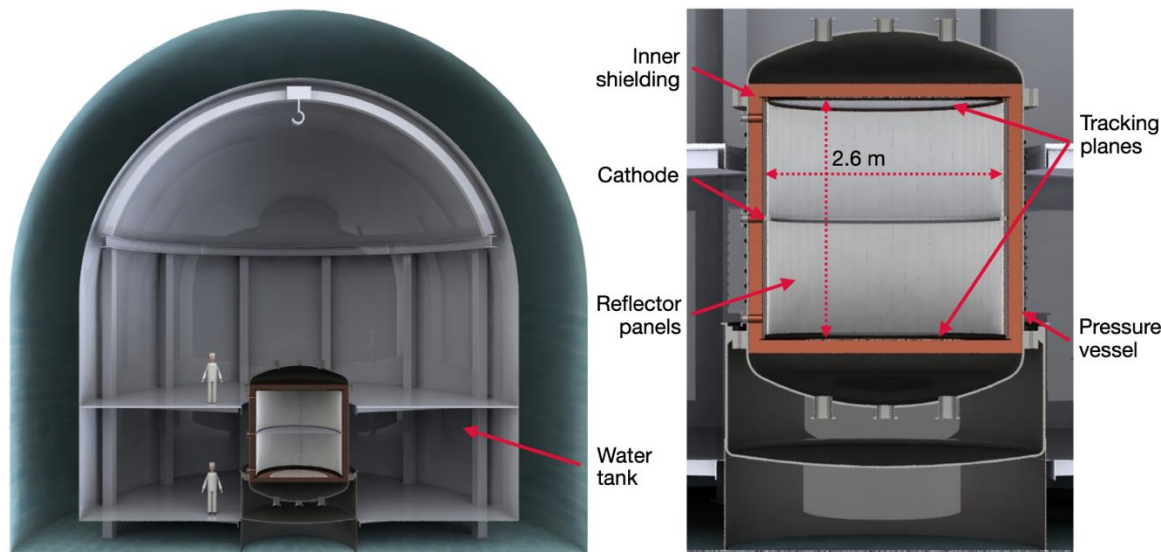


Figure 2: Cutaway sketch of the anode section of the TPC. The configuration of the charge collection tiles is shown below a cutaway of the anode backbone. [4]



# Some designs - NEXT

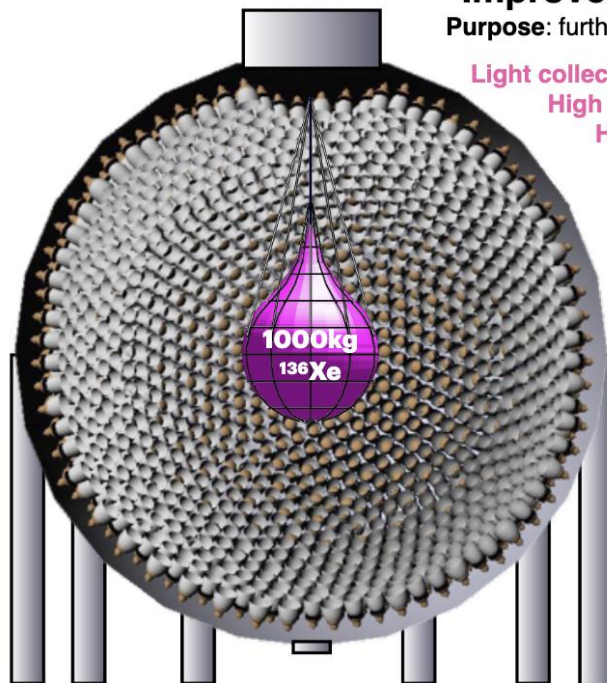
- Standard water tank with potential neutron absorption doping (like Gd)
- Inner copper shielding ([2005.06467](#))



**Figure 2.** Left: conceptual design of a tonne-scale NEXT detector installed inside a water tank. Right: detail of the internal structures of the detector. The active volume, 2.6 m in diameter and height, would hold a mass of  $^{136}\text{Xe}$  of approximately 1109 kg at 15 bar.

# Some designs – KAMLAND-Zen2

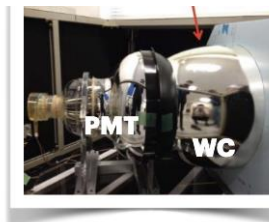
- From [Spencer Axani](#) yesterday
- Existing LS veto with addition of PEN balloon



## Improved energy resolution

**Purpose:** further separate  $2\nu\beta\beta$  from the  $0\nu\beta\beta$ .

Light collection with Winston Cones (x1.8)  
High light yield scintillator (x1.4)  
High QE 20" PMTs (x1.9)

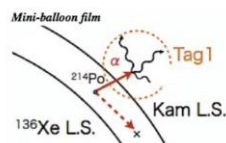


**4% → 2% energy resolution**

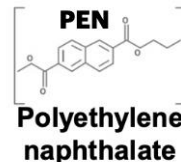
**x100 reduction in  $2\nu\beta\beta$  background rate.**

## Improved inner balloon

**Purpose:** reduce backgrounds originating from balloon.

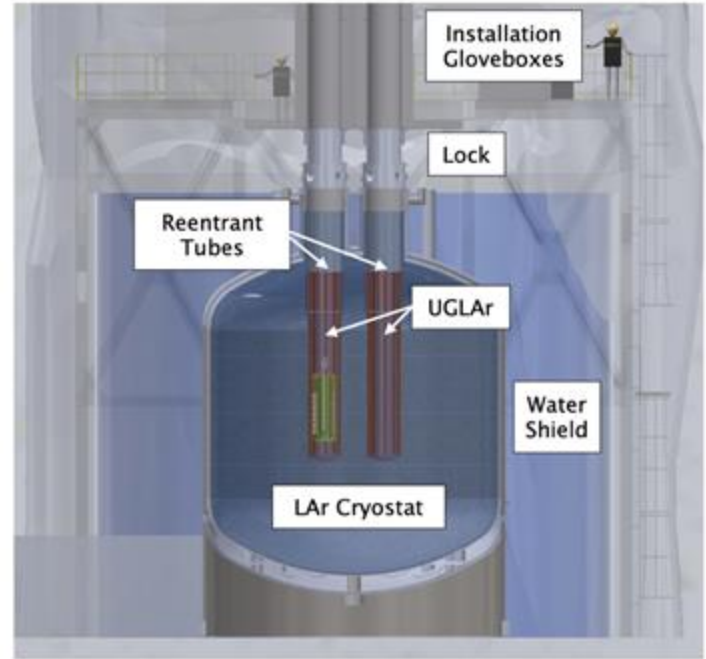


**Tag  $^{214}\text{Bi}$  decays.**



# Some designs - GERDA/LEGEND

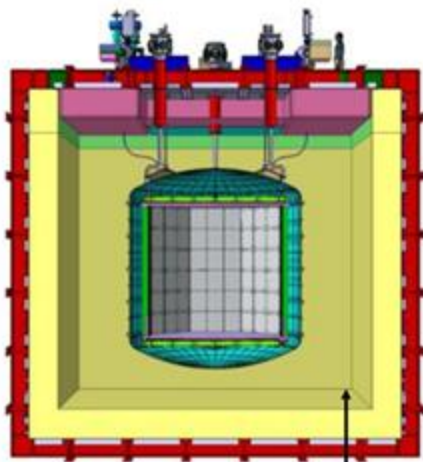
- Liquid argon active veto, also providing cooling to the Ge detectors
- 7 m diameter cryostat, 332 tonnes of LAr
- Inner reentrant tubes with 0.95 m diameter to hold the detectors with underground LAr
  - $^{42}\text{Ar}$  decays to  $^{42}\text{K}$  which famously can attach to surfaces and beta decay at 3525 keV
  - UAr eliminates need for nylon shrouds used in GERDA
- Light detection through WLS polystyrene fibers coupled to SiPMs (coated with TPB)





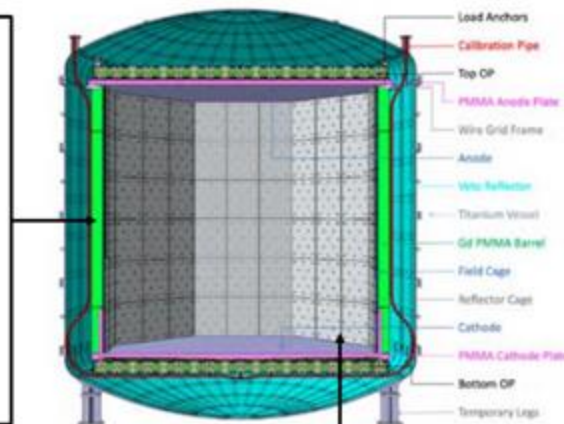
# Some designs - DarkSide-20k

- Outer veto for muons of atmospheric argon
- Neutron inner veto (skin) of UAr



- Cosmogenic (Outer) Veto**
- 650 t of Atmospheric Argon (AAr)
  - Membrane “ProtoDUNE-like” cryostat  $8 \times 8 \times 8 \text{ m}^3$
  - Sparsely instrumented with SiPMs.

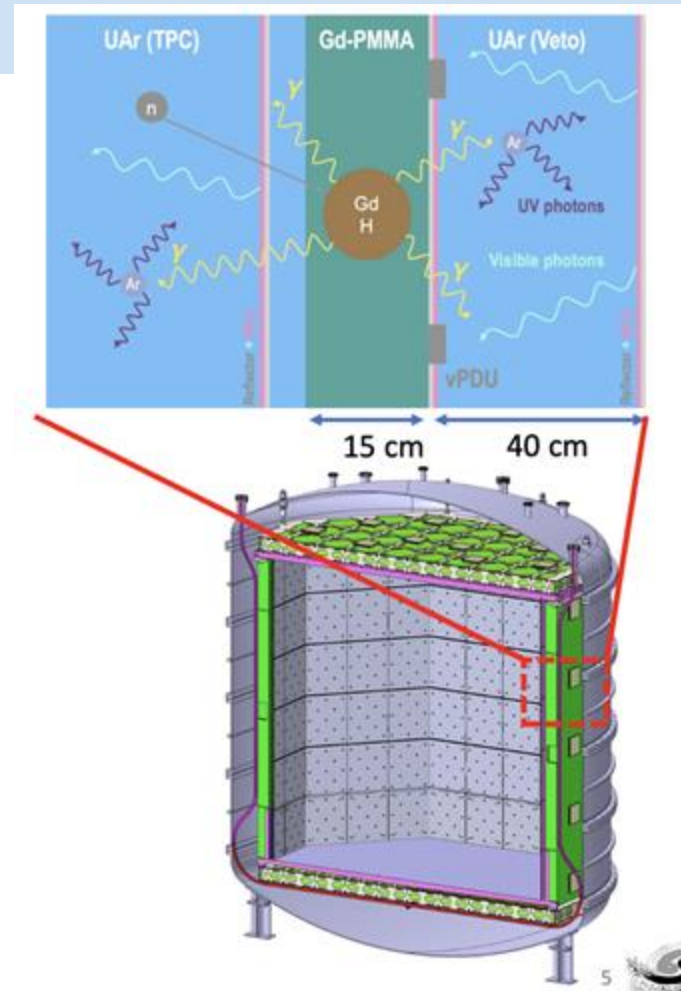
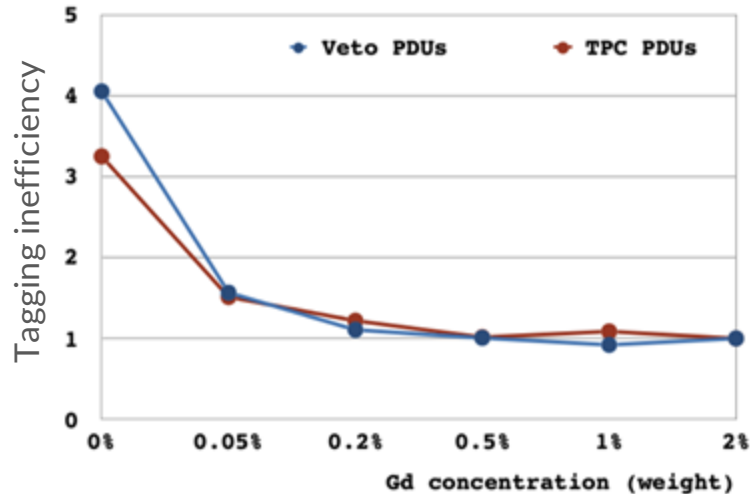
- Neutron (inner) Veto**
- 32 t of UAr.
  - Enclosed in Stainless Steel vessel and HDPE neutron shield.
  - Equipped with SiPMs covering  $5 \text{ m}^2$  from UK groups.
  - Light yield: 2 PE/keV.



- Dual-phase TPC**
- 50 t of UAr (20 t fiducial).
  - Two optical planes covering  $21 \text{ m}^2$ .
  - S1 and S2 yield is 10 PE/keV and 20 PE/e<sup>-</sup> respectively.



# Some designs - DarkSide-20k

- Outer veto for muons of atmospheric argon
- Neutron inner veto (skin) of UAr
  - Key is 15 cm Gd-loaded PMMA (acrylic)
  - Capture gammas interact in the argon volume



# Some comments - Requirements



		
<b>Passive shield</b>	Cost effective Labor saving (short term)	Limited capability No input to bkg model Extremely costly to add post facto
<b>Simple veto</b>	Known behaviour Simple modeling Balance between cost and complexity / capability	Balance between cost and complexity / capability
<b>More Most capable outer detector (MCOD)</b>	PID, fiducial., coincidence logic Save by spending	Enables physics Risk mitigation Future proofing

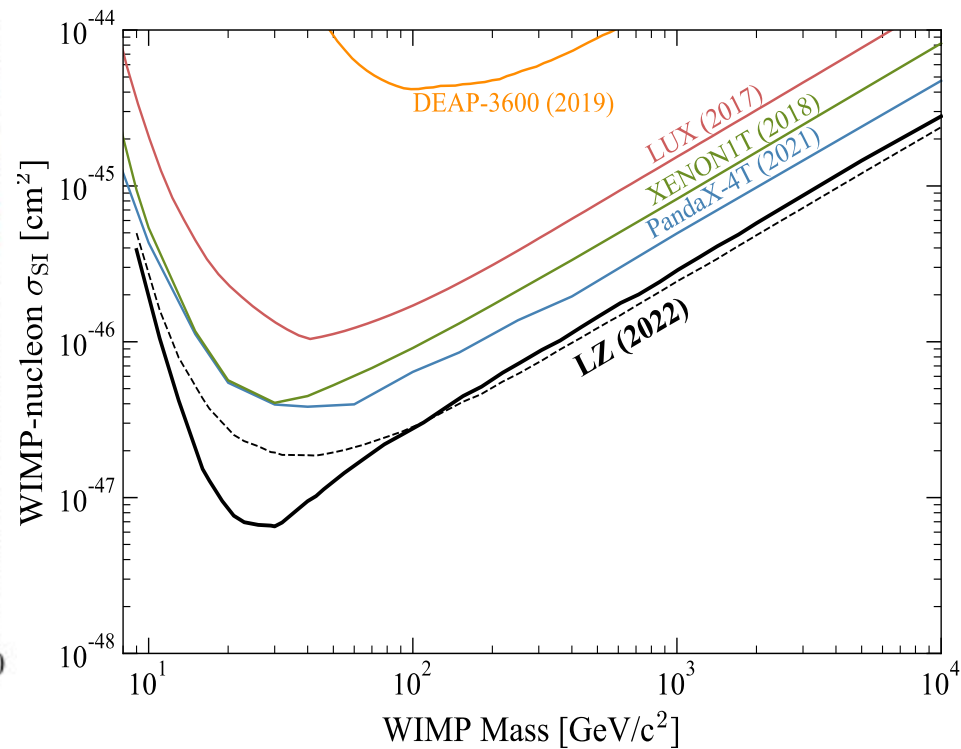
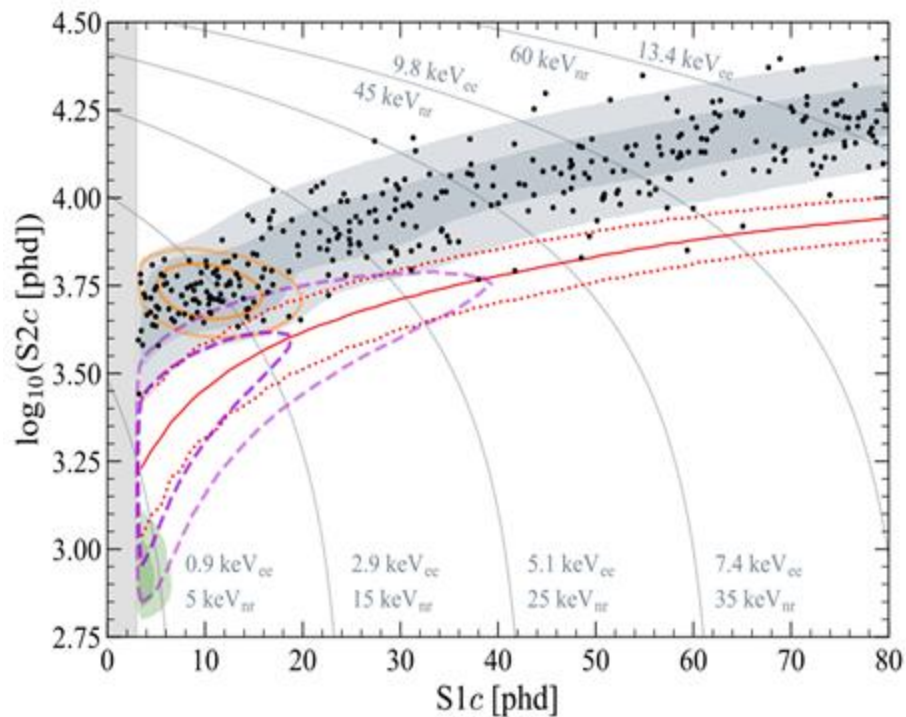


# Some comments - Requirements

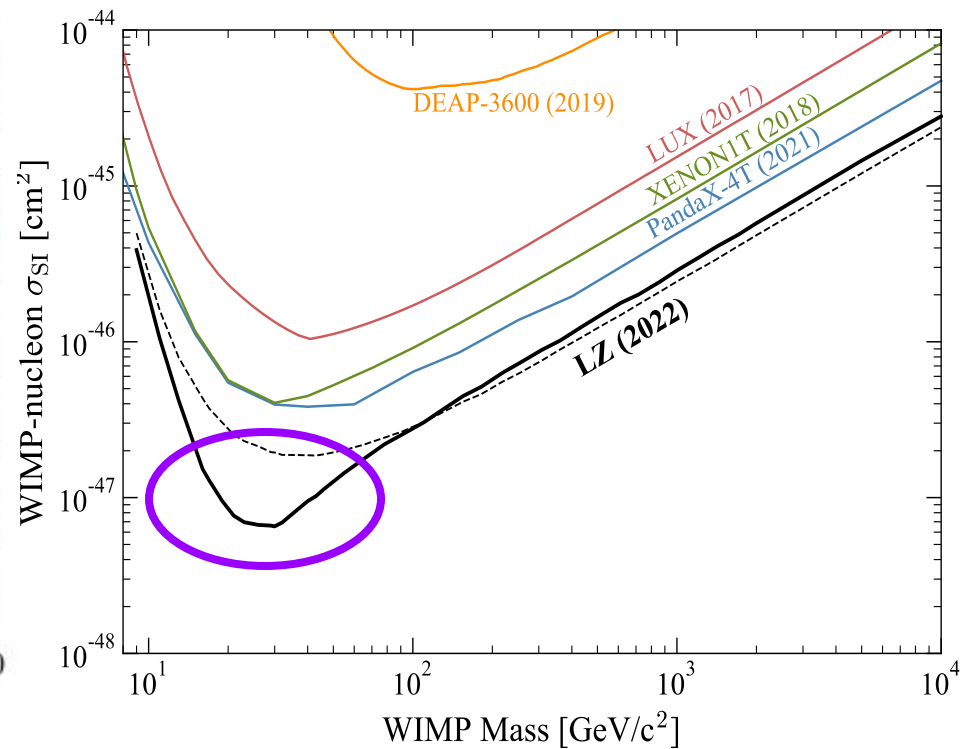
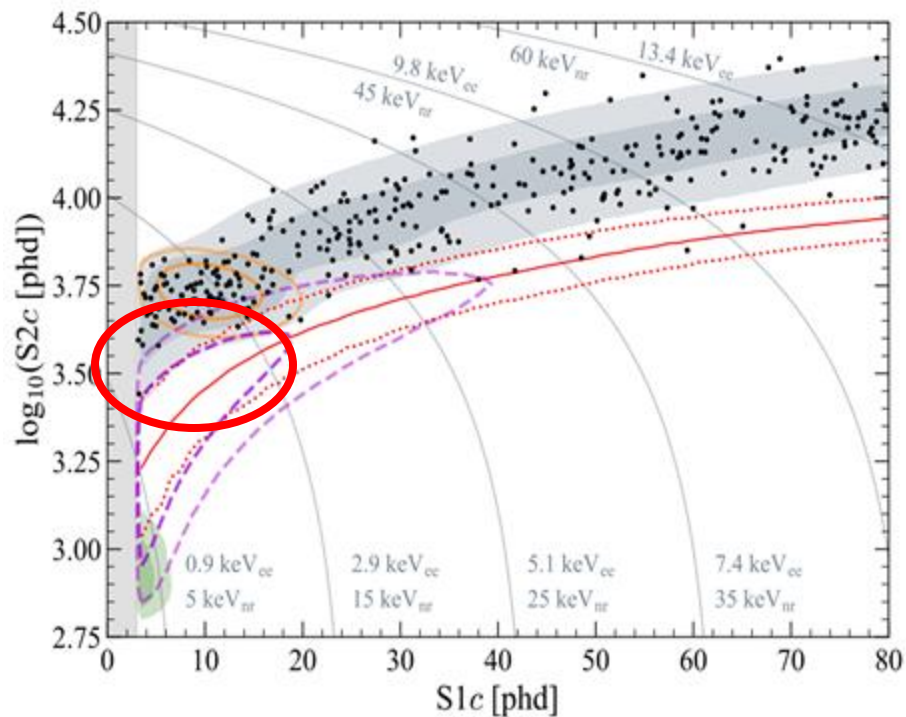


- Is the active veto *required*?
  - In LZ, OD cost was ~9% of the total project cost
  - Extra complication to build and fill
  - How does that scale with increased size?
- In large liquid xenon detector, for double beta decay, the skin is potentially the most important veto, knocking down internal  $^{208}\text{Tl}$  by order of magnitude
- Dark matter drives the neutron and high performance (low threshold) requirements
  - For such a large detector, it's not so obvious (to me) what that requirement needs to be (how much rejection is provided by MS in the TPC itself?)
- *Back to vibes - hard to put a cost on having that extra handle when the difference between a signal and a fluctuation is a couple of counts*

# Case study – LZ WS2022

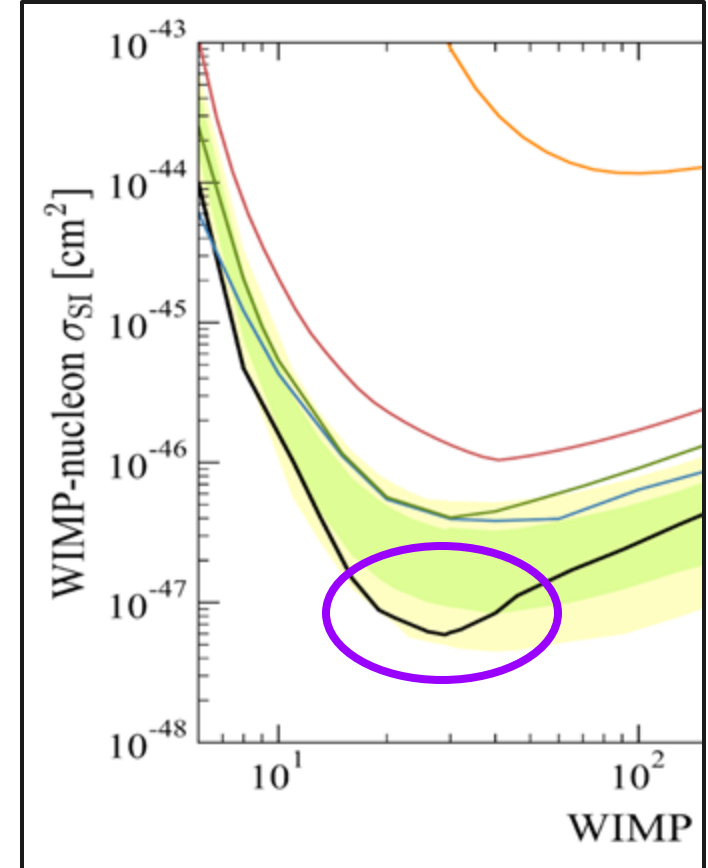
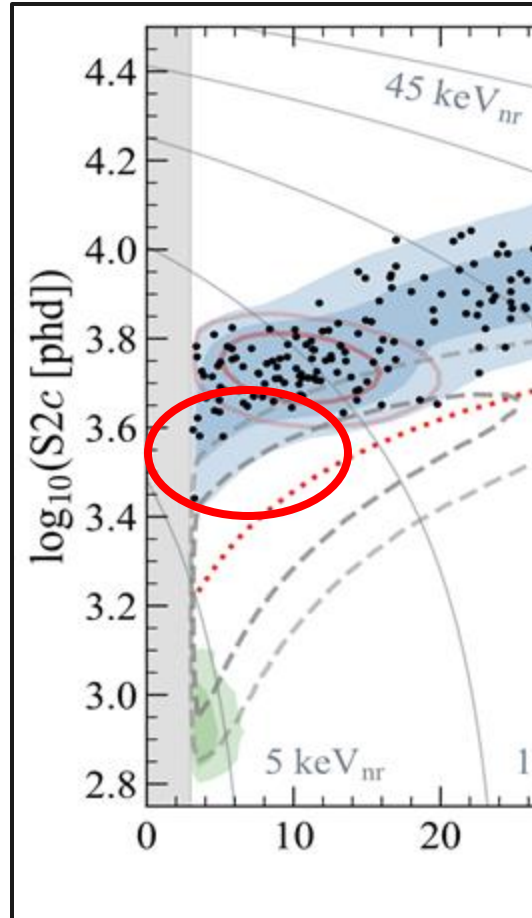


# Case study – LZ WS2022



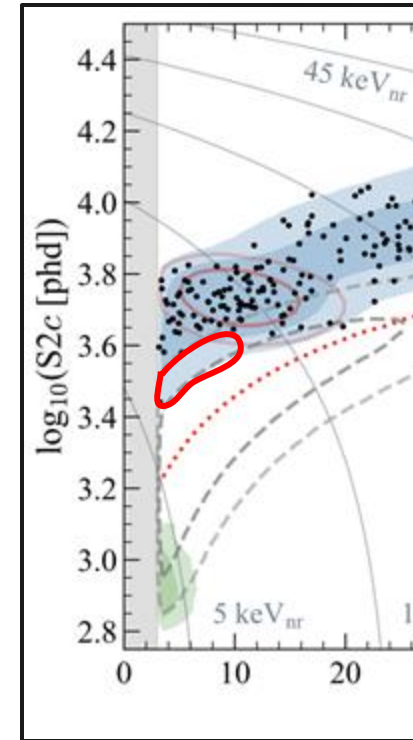
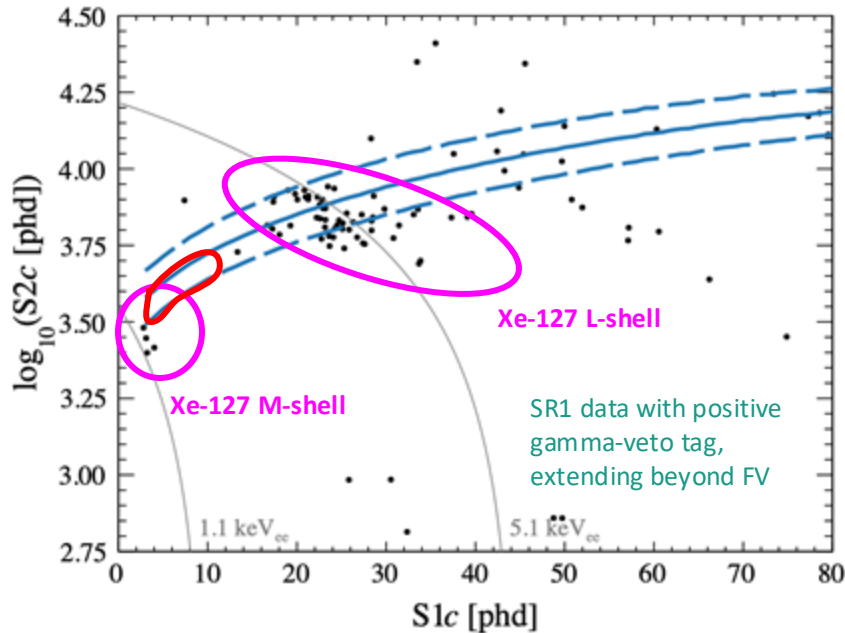
# Case study

- LZ WS2022 had a downward fluctuation in the observed upper limit (purple ellipse)
- Attributed to deficiency of events under the Ar-37 population (red ellipse).
- Alternative is problem with signal efficiency that changed over the run



# Case study

- The skin detector can tag gammas from Xe-127 electron capture decays
- The Xe-127 M-shell population is exactly in the deficient region
- Served as a cross check - 4 events are distributed throughout run consistent with the expected number of events.



# Conclusion

- Surprising variety of design choices for what started as "shielding/water tank"
- Each detector/collaboration needs to think about its own requirements carefully
- Critical component of a complete background model

