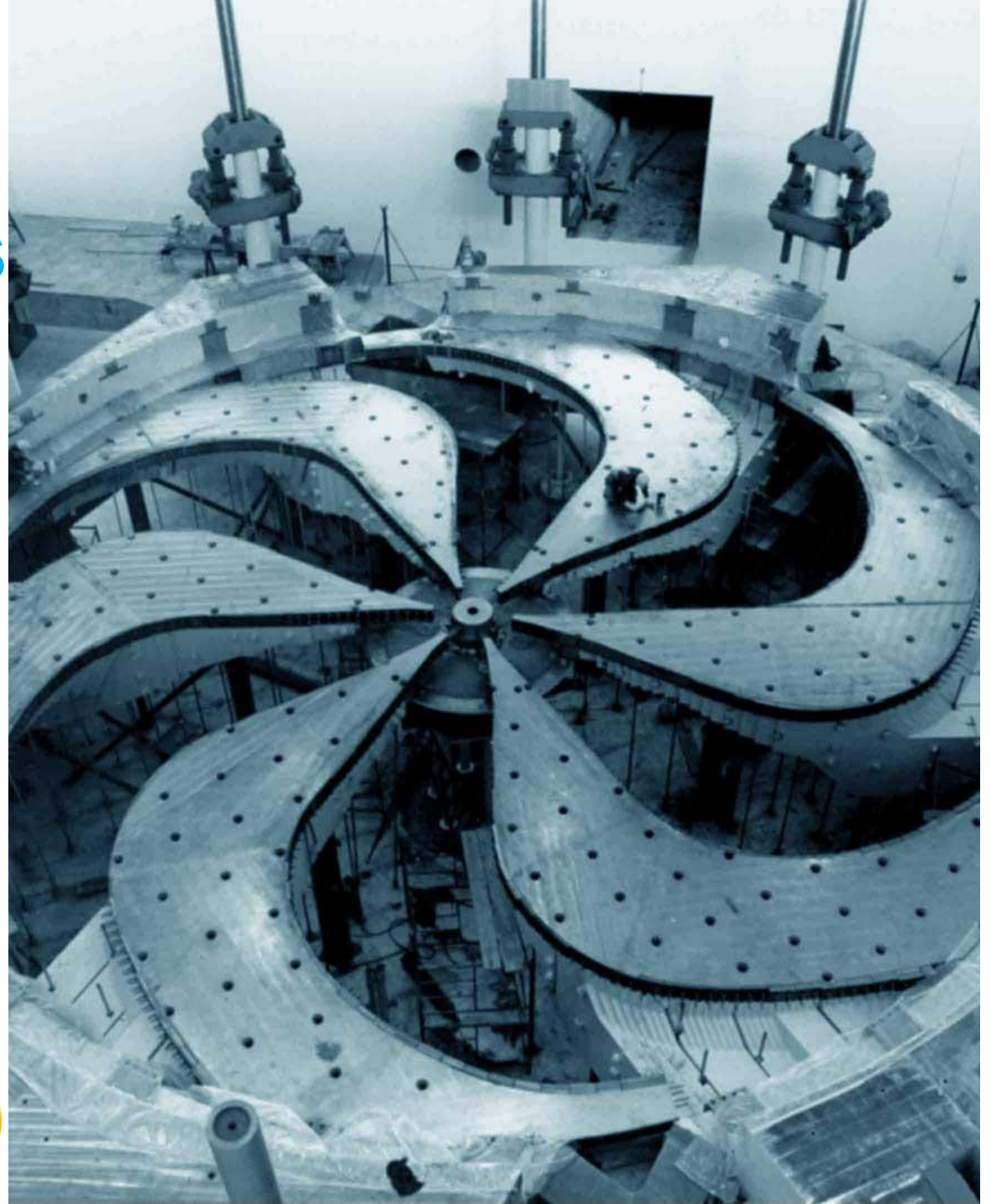


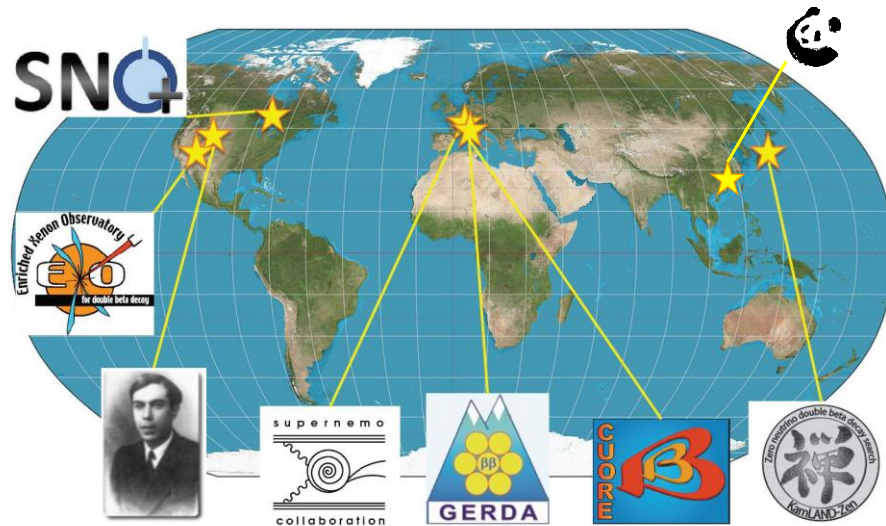
(Ab Initio) Theory Status for Neutrinoless Double Beta Decay NMEs

Jason D. Holt
TRIUMF, McGill University
Next-Gen Xe; Montreal QC
Nov 12, 2025

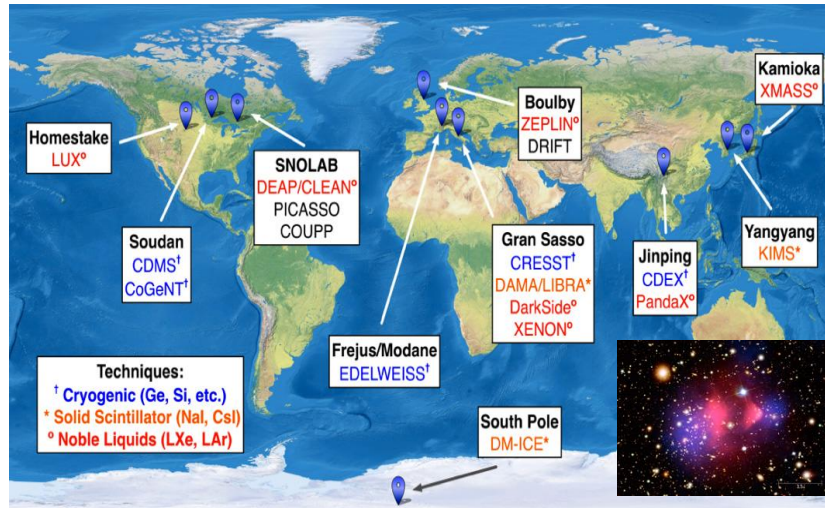


Worldwide search for BSM physics w/ neutrinos, dark matter, and radioactive molecules!

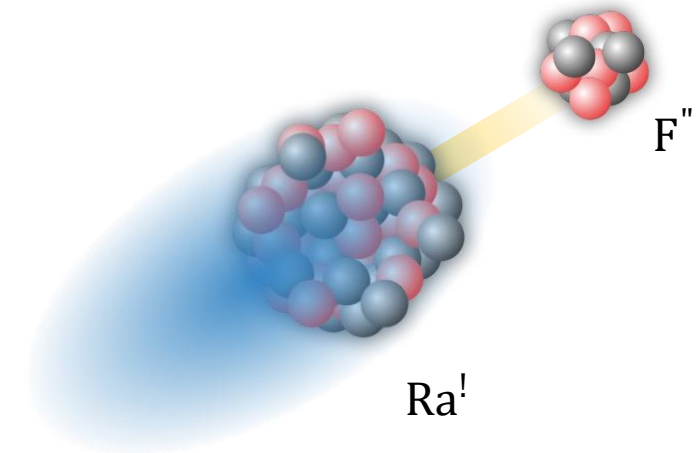
$0\nu\beta\beta$ Decay



Dark Matter Direct Detection



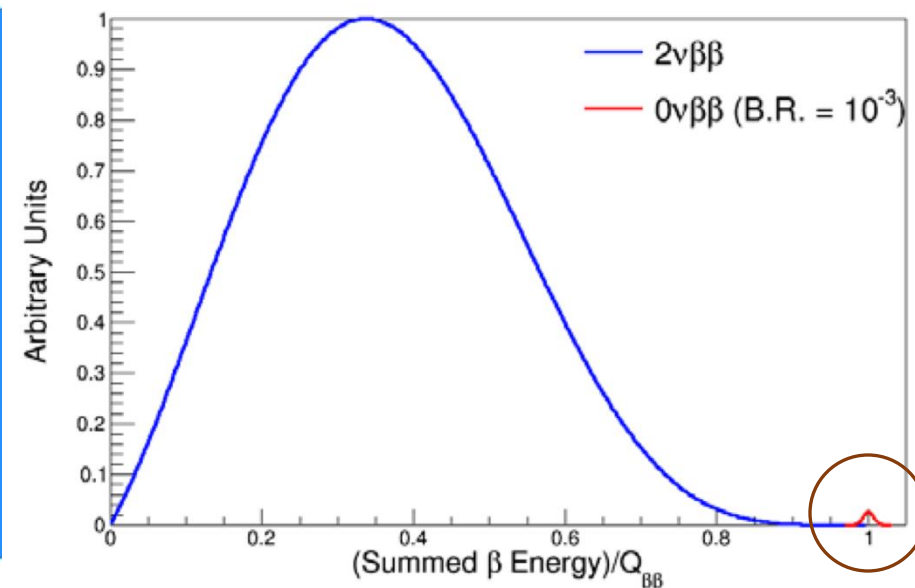
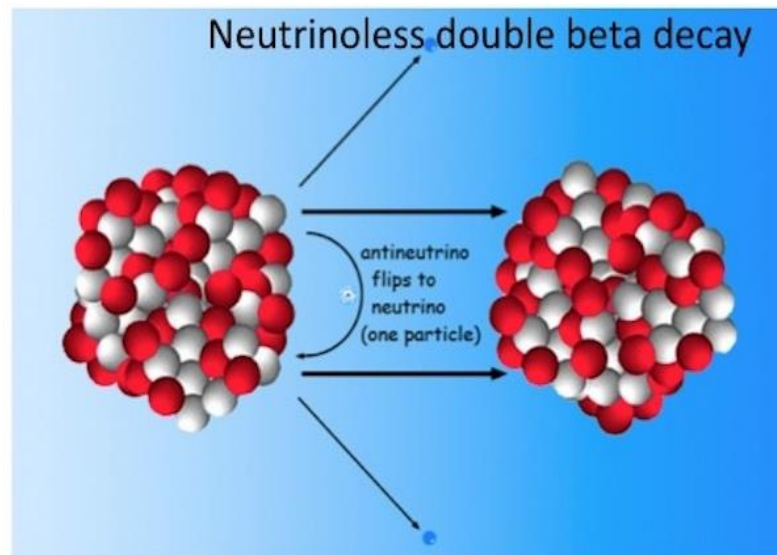
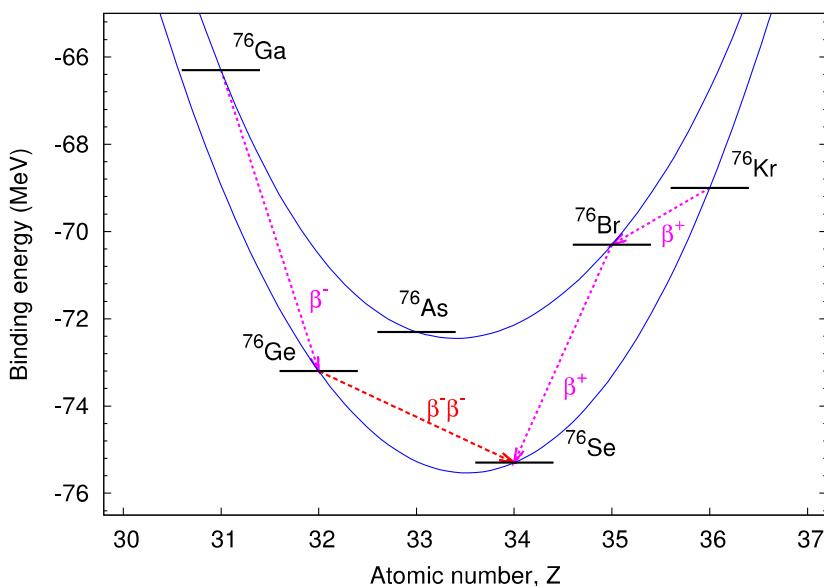
Symmetry Violation



Billions invested worldwide – theory of the nucleus is essential

Strategic planning for discovery (**motivation**) + **interpretation**

Neutrino own antiparticle $\longleftrightarrow$ $0\nu\beta\beta$ decay



Tremendous impact on BSM physics

Lepton-number violation

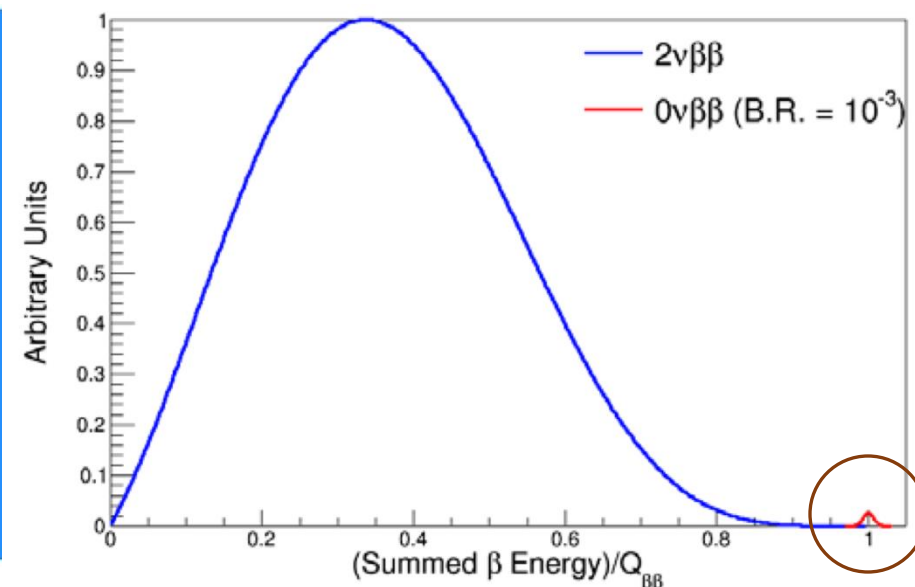
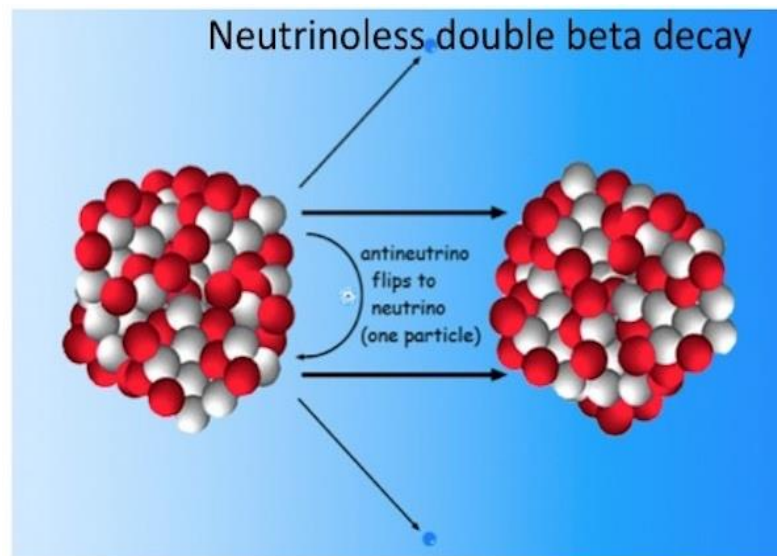
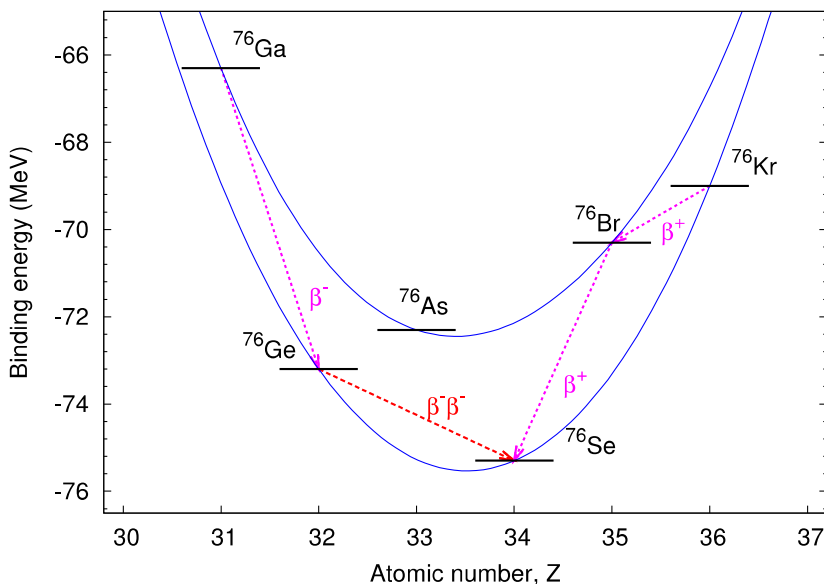
Majorana character of neutrino

Matter/antimatter asymmetry

Absolute neutrino mass scale

$$\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1} = G^{0\nu} |M^{0\nu}|^2 \boxed{\langle m_{\beta\beta} \rangle^2} \langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei} m_i \right|$$

Neutrino own antiparticle $\longleftrightarrow$ $0\nu\beta\beta$ decay



Tremendous impact on BSM physics

Lepton-number violation

Majorana character of neutrino

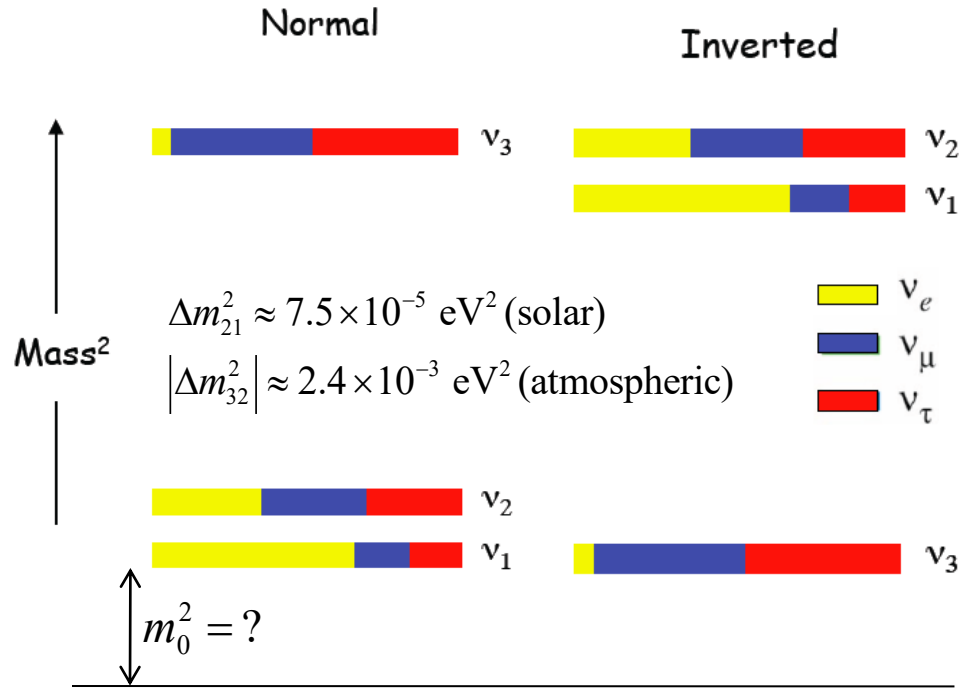
Matter/antimatter asymmetry

Absolute neutrino mass scale

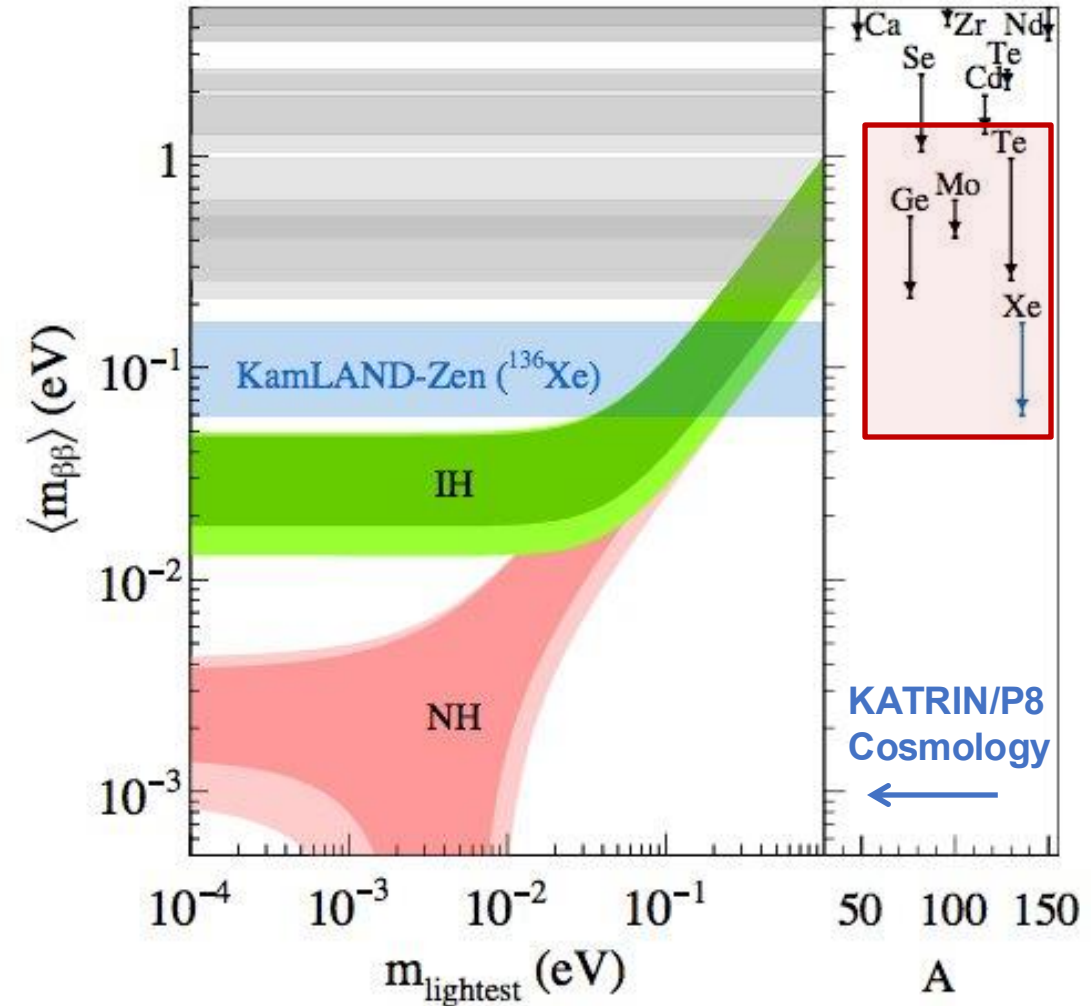
$$\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1} = G^{0\nu} \boxed{M^{0\nu}}^2 \boxed{\langle m_{\beta\beta} \rangle^2} \langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei} m_i \right|$$

NME not observable: must be calculated

Next-generation searches probe **Inverted Hierarchy**

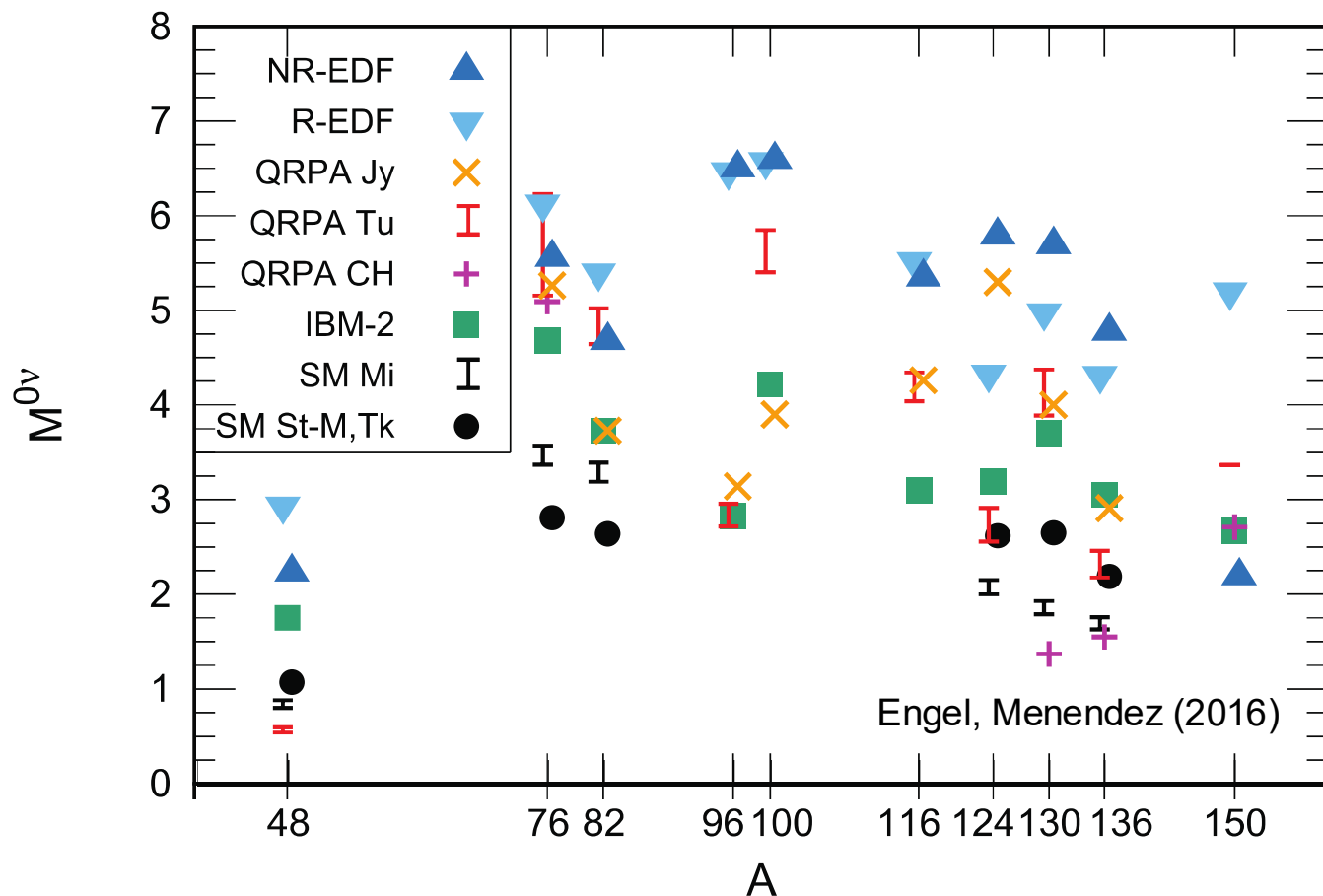


$$\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1} = G^{0\nu} \boxed{M^{0\nu}}^2 \langle m_{\beta\beta} \rangle^2 \quad \langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei} m_i \right|$$



Spread from **Nuclear Matrix Element**; bands **do not** represent rigorous uncertainties

All calculations to date from **extrapolated** phenomenological models; large spread in results

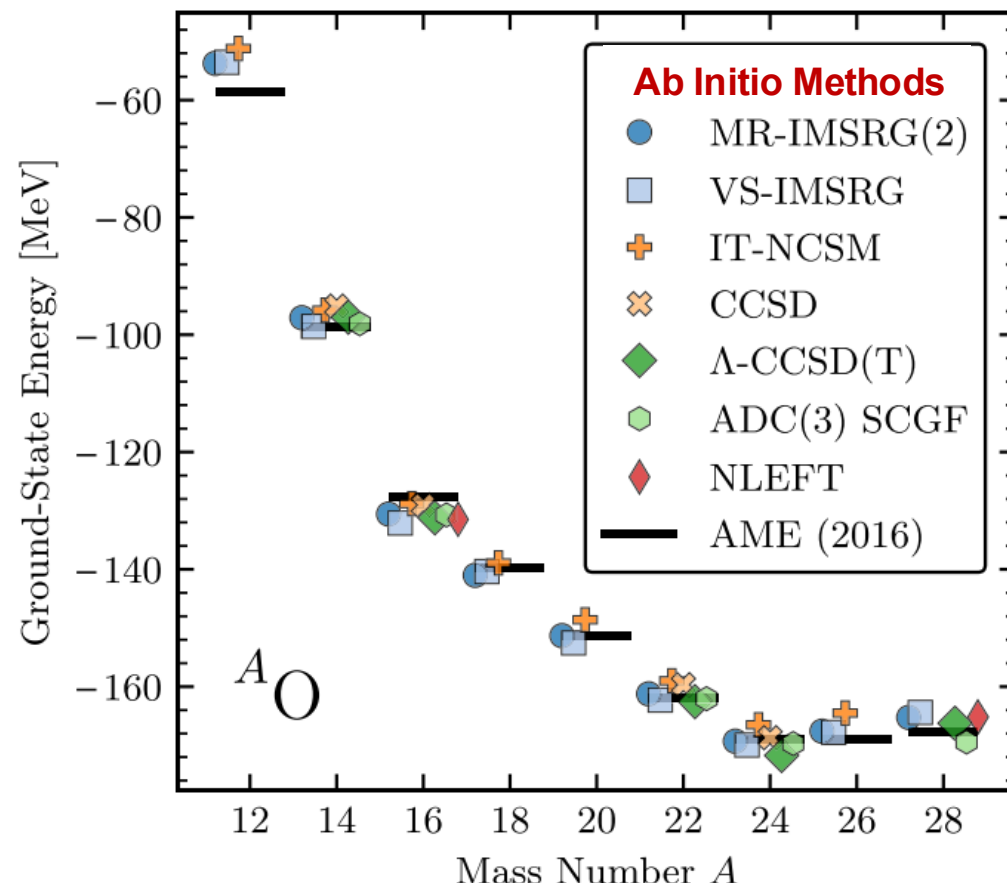
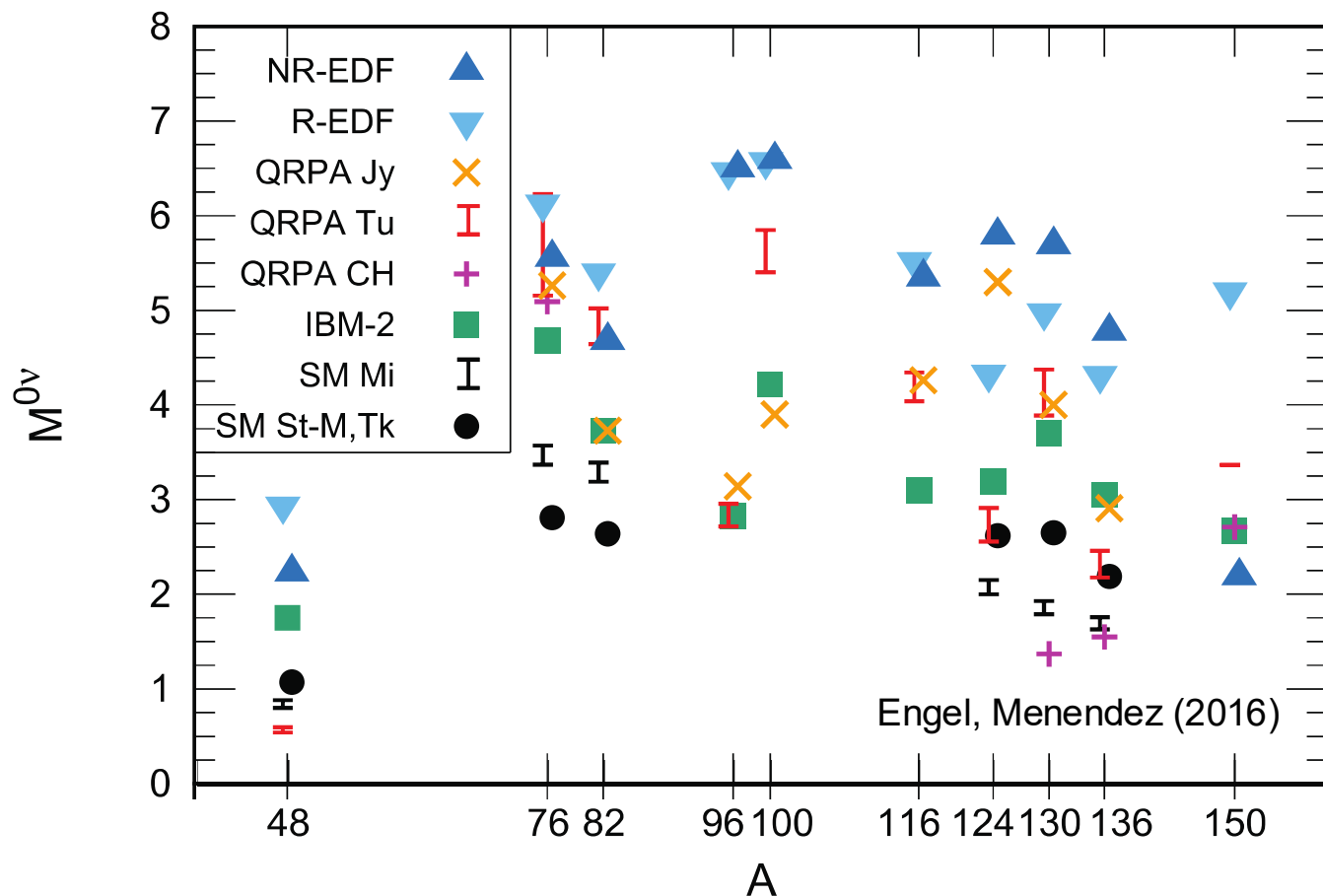


All models missing essential physics: correlations, single-particle levels, two-body currents

Impossible to assign rigorous uncertainties... but, hey, could be worse



All calculations to date from extrapolated phenomenological models; large spread in results

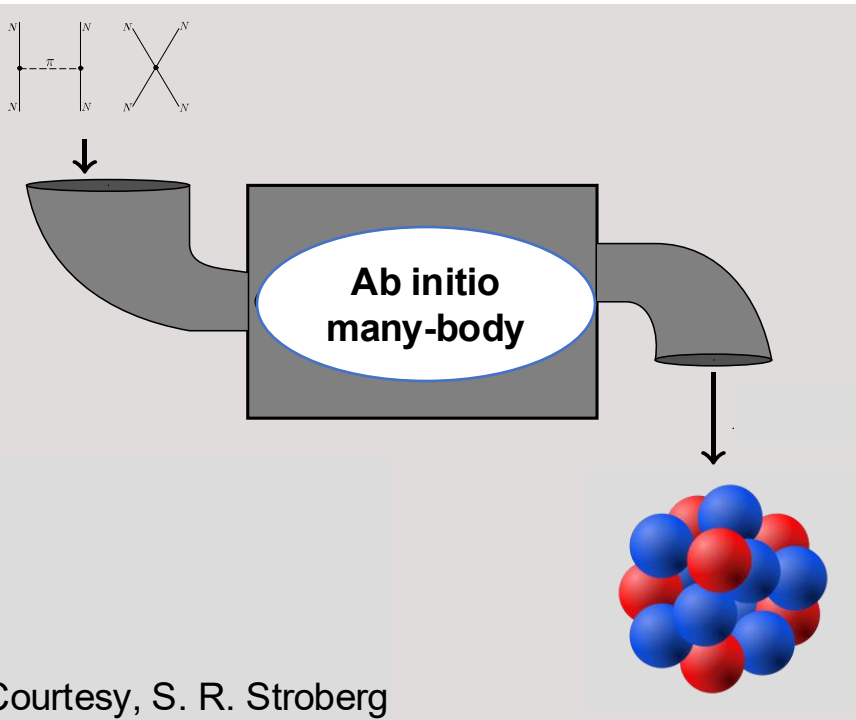


New approach to NME calculations: **ab initio theory** consistent results when extrapolated

Aim of modern nuclear theory: develop unified *first-principles* picture of structure and reactions

(Approximately) solve nonrelativistic Schrödinger equation

$$H\psi_n = E_n\psi_n$$

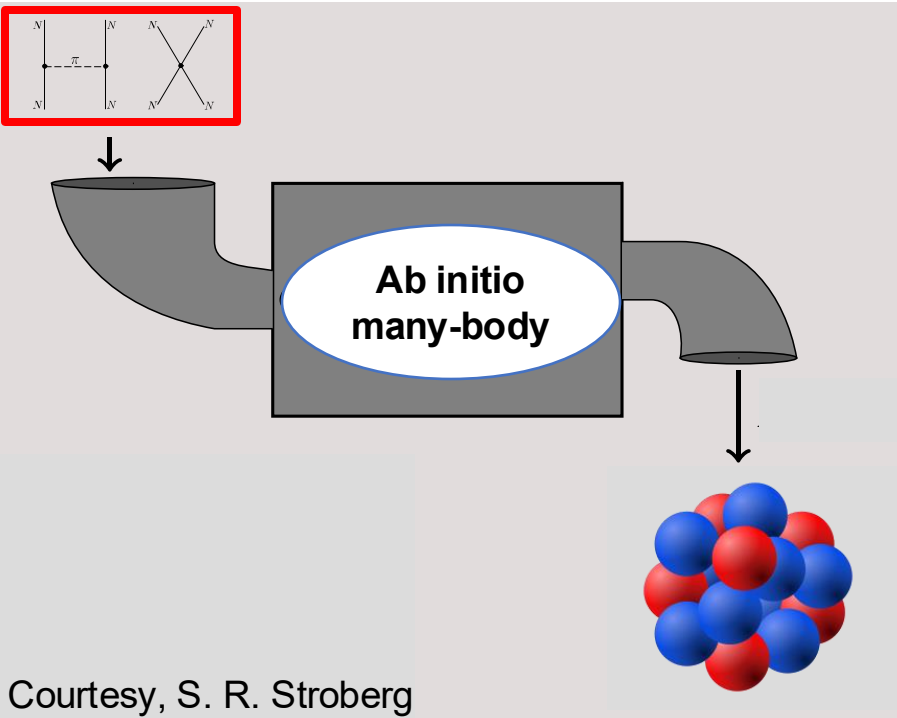
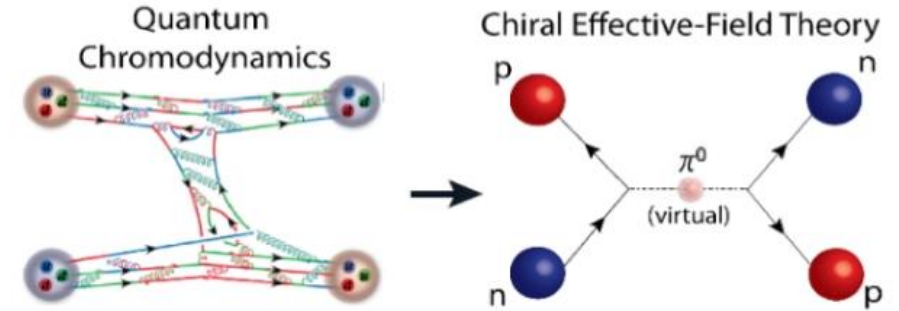


Courtesy, S. R. Stroberg

Chiral Effective Field Theory: low-energy expansion of QCD, nucleons + pions

High-energy information encoded in contact terms

$$\boxed{H}\psi_n = E_n\psi_n$$



Courtesy, S. R. Stroberg

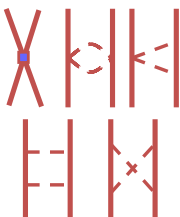
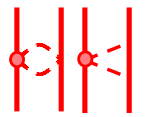
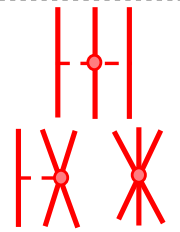
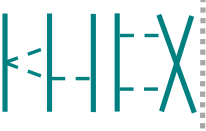
Chiral Effective Field Theory

Consistent treatment of
 - 2N, 3N, 4N, ... forces
 - Electroweak physics

Quantifiable uncertainties

Interactions

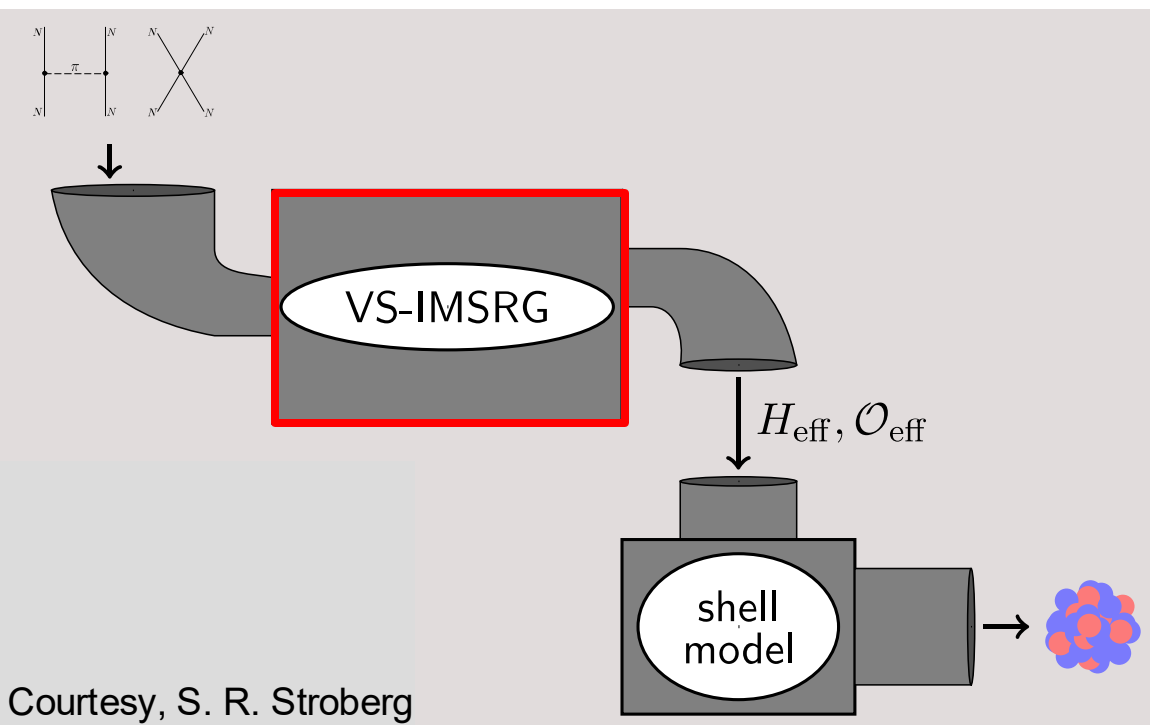
1.8/2.0, N2LO_{GO}, N3LO_{LNL}
 34 non-implausible

	NN force	NNN force	NNNN force
Q^0 LO			
Q^2 NLO			
Q^3 N ² LO			
Q^4 N ³ LO			

Polynomially scaling methods tailored for doubly magic, semi-magic chains

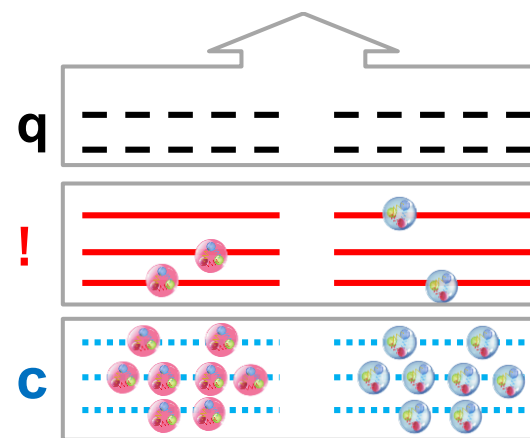
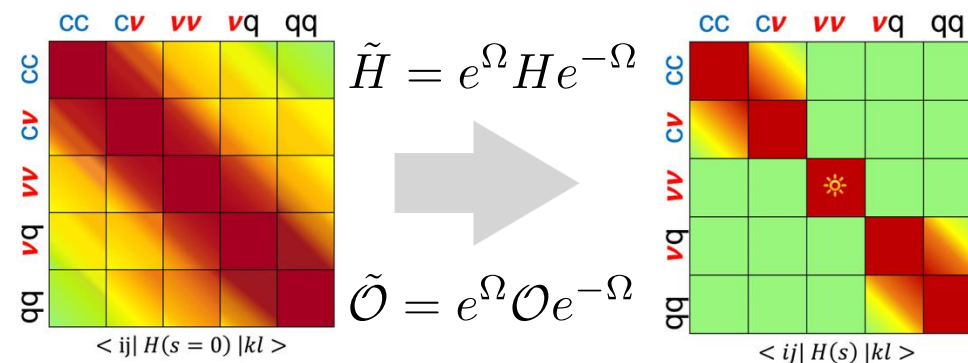
Valence-space IMSRG: Novel approach for all open-shell nuclei

$$H\psi_n = E_n\psi_n$$



Courtesy, S. R. Stroberg

Ab initio → scope of nuclear shell model



Valence-space
Hamiltonian

Decoupled core



Ragnar Stroberg

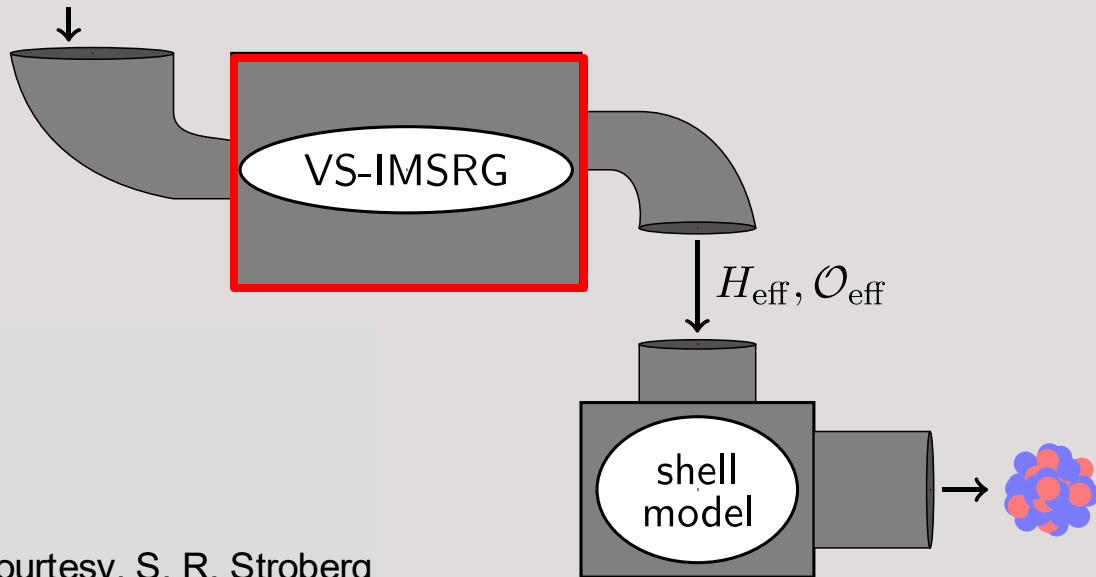
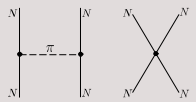
Nucleus-Dependent Valence-Space Approach to Nuclear Structure

S. R. Stroberg, A. Calci, H. Hergert, J. D. Holt, S. K. Bogner, R. Roth, and A. Schwenk
Phys. Rev. Lett. **118**, 032502 – Published 20 January 2017

Aim of modern nuclear theory: develop unified *first-principles* picture of structure and reactions

(Approximately) solve nonrelativistic Schrödinger equation

$$H\psi_n = E_n\psi_n$$



Courtesy, S. R. Stroberg

Methods Exact up to Truncations

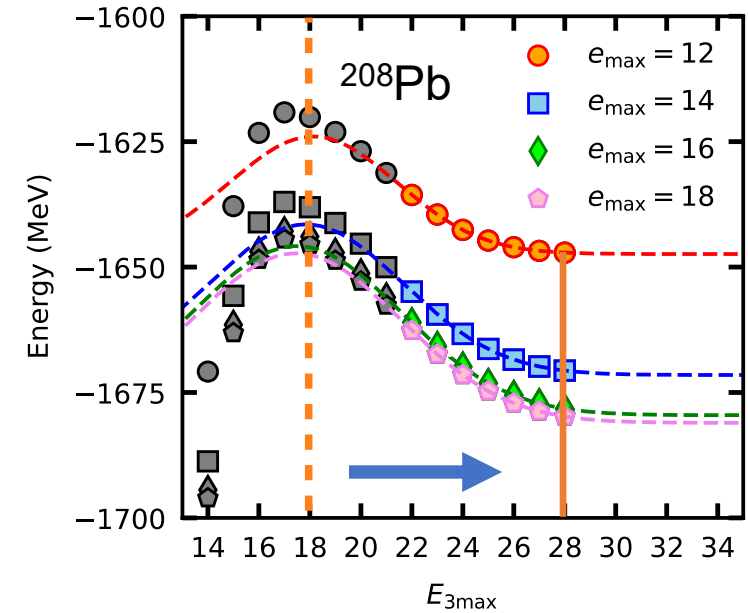
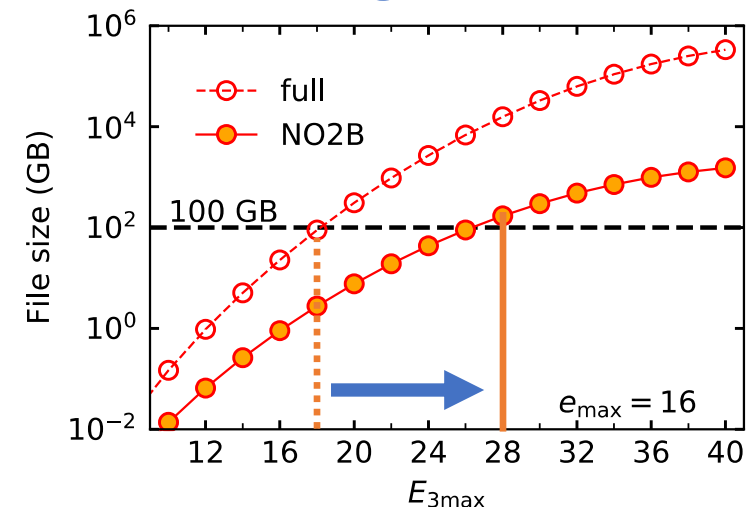
- ✓ Single-particle basis $e_{\max} = 2n + l$
- ✓ Storage limits of 3N forces $e_1 + e_2 + e_3 \leq E_{3\max}$
- 👤♂ Many-body operators: e.g., CCSD(T), IMSRG(2)

Limited by typical memory/node: $e_1 + e_2 + e_3 \leq E_{3\max} = 18$

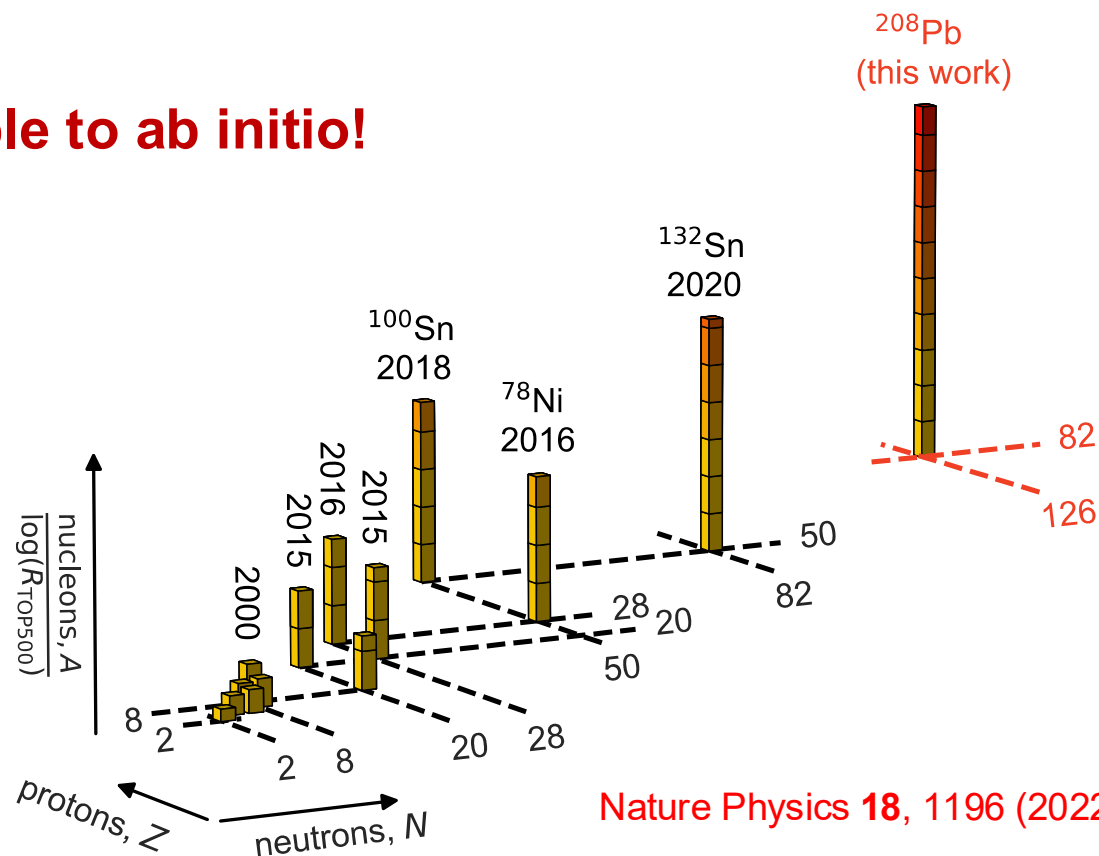
Clever storage reduces needs: TB $\rightarrow$ GB

^{208}Pb now straightforward:
Solving the Schrödinger equation for 200+ nucleons!

(Super)heavy accessible to ab initio!

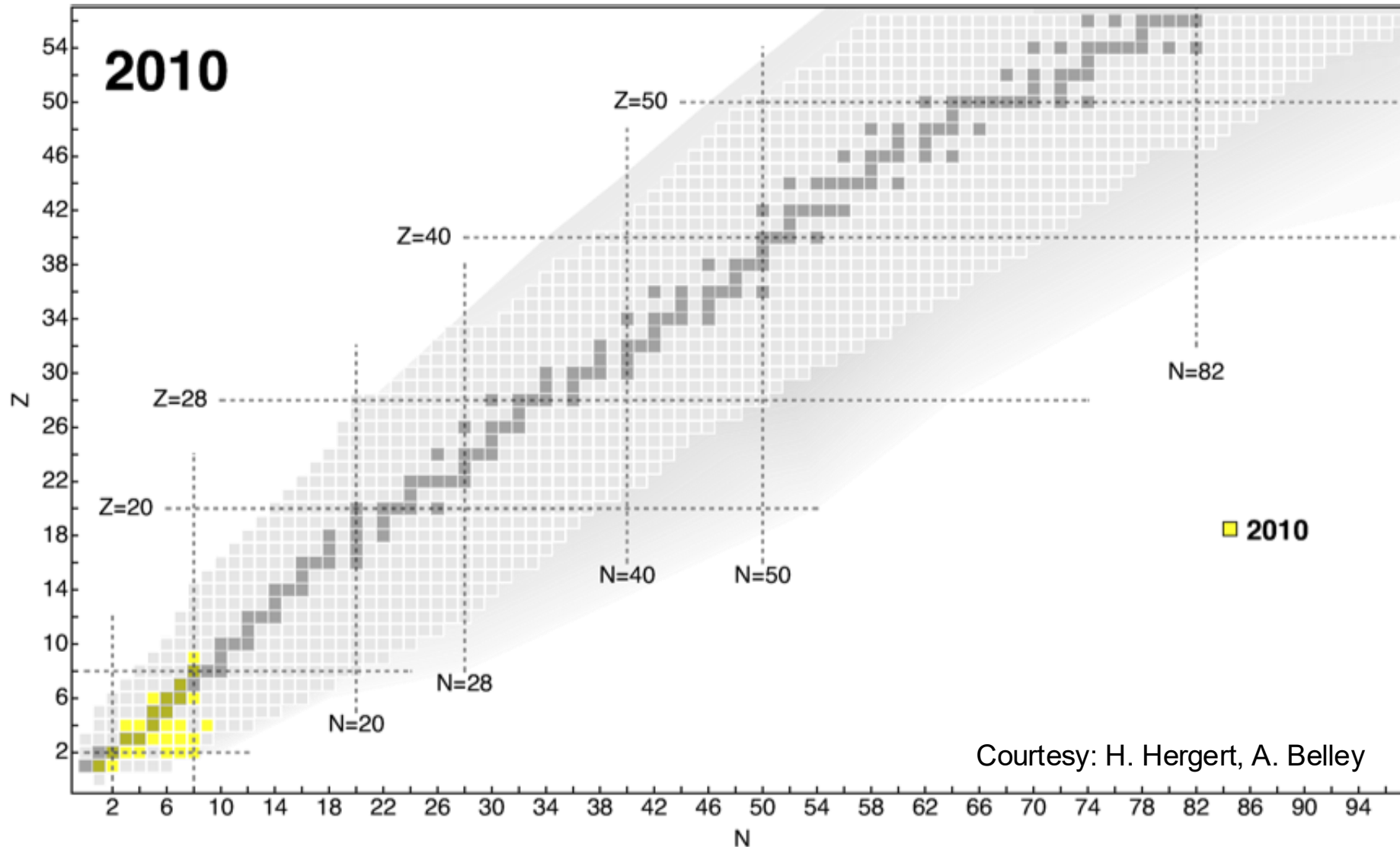


16 years: $^{16}\text{O} \rightarrow ^{56}\text{Ni}$
6 years: $^{56}\text{Ni} \rightarrow ^{208}\text{Pb}$



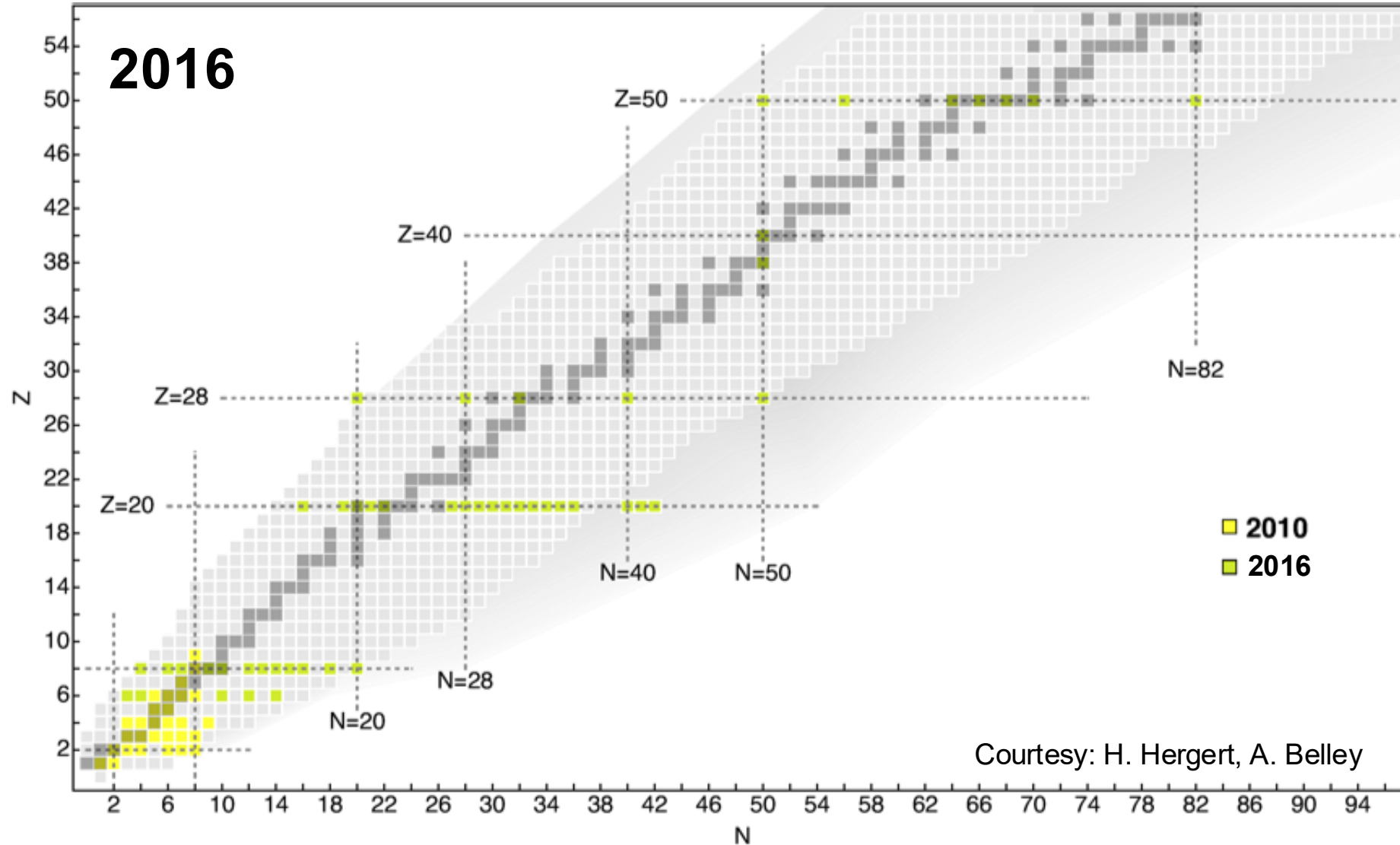
Nature Physics **18**, 1196 (2022)

2010 Quasi-exact methods for light nuclei (^{16}O); limited capability for 3N forces



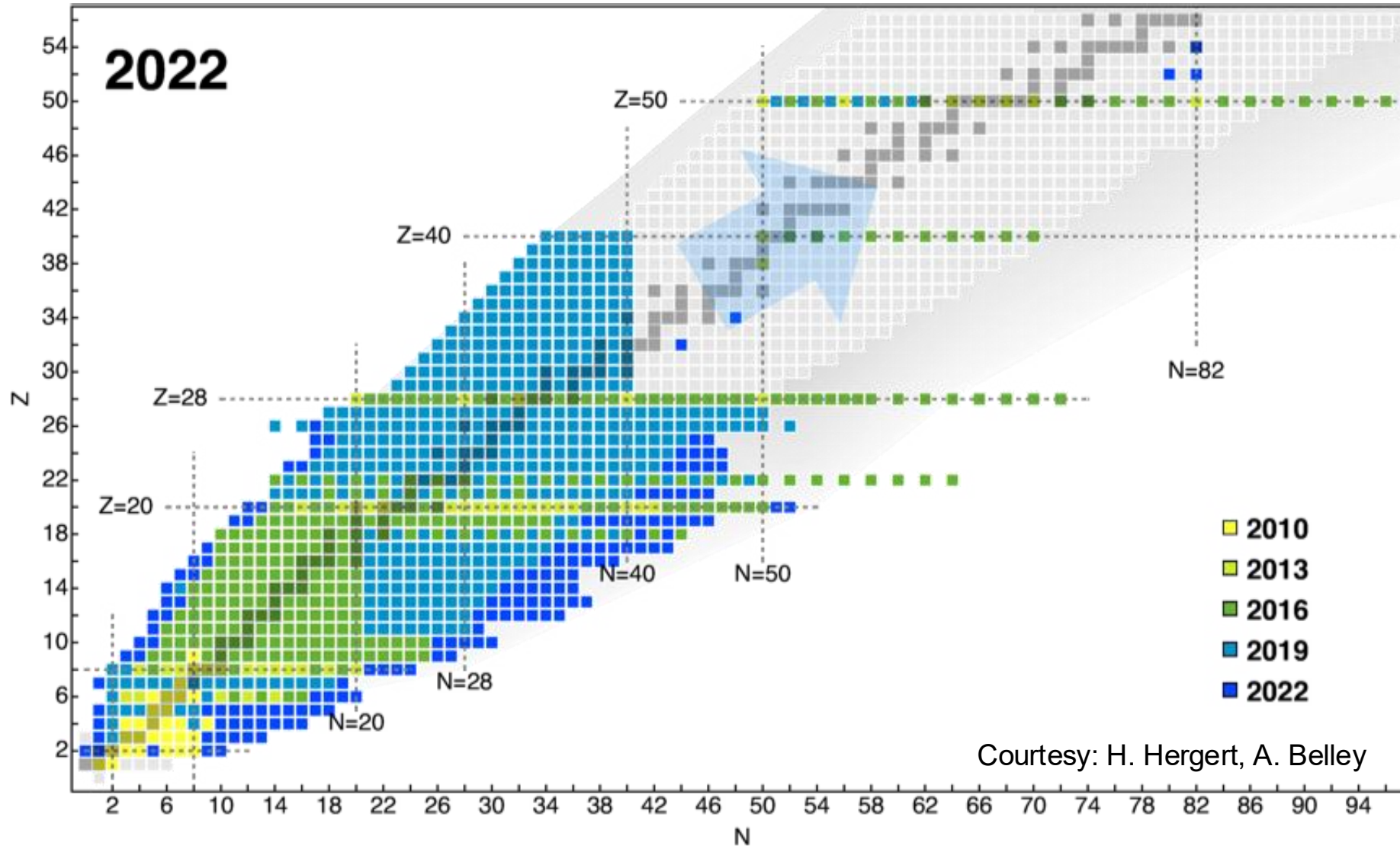
2016 Polynomially scaling methods (Coupled Cluster, IMSRG)

Progress to doubly magic $^{56/78}\text{Ni}$ + semi-magic chains

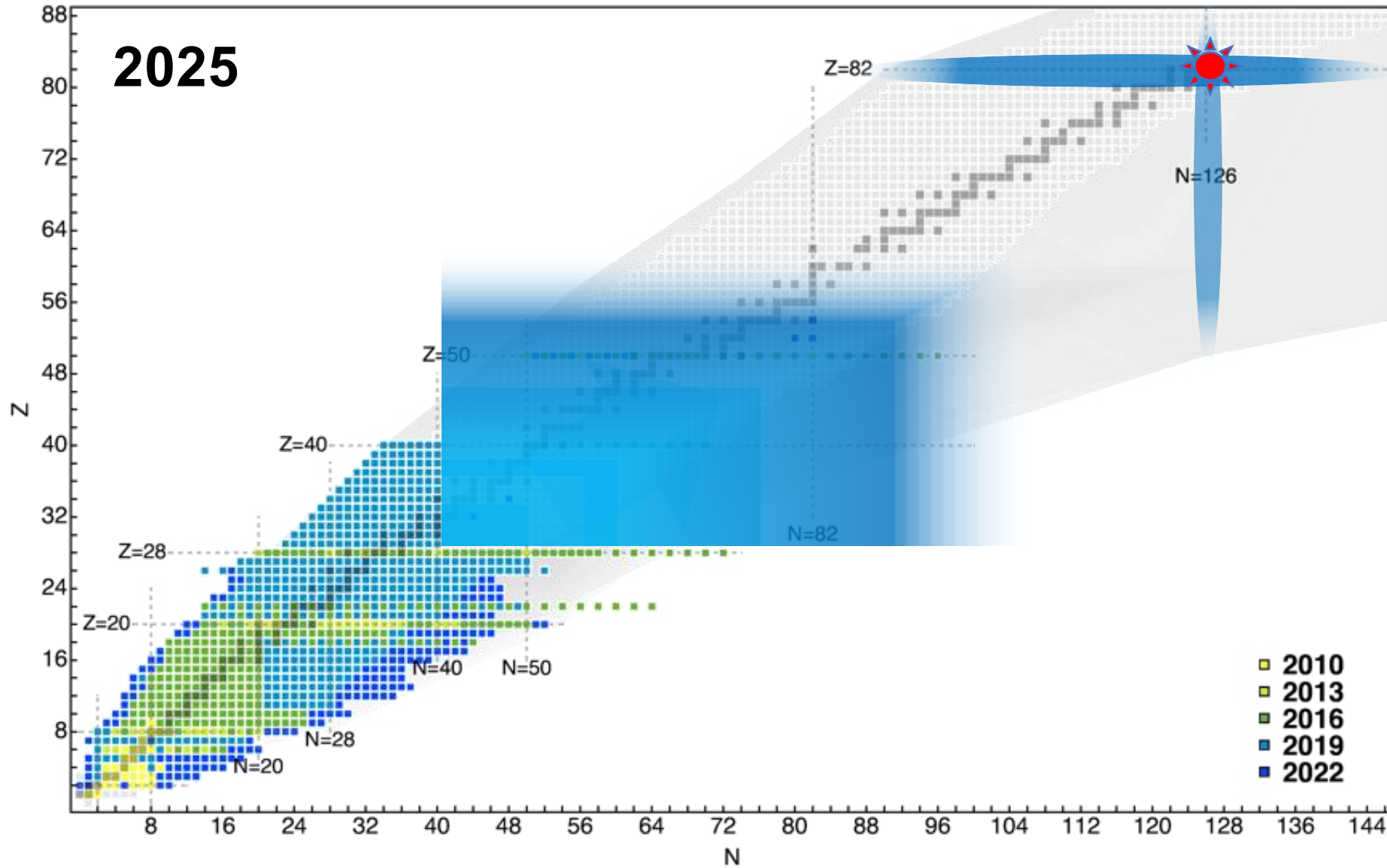


2022 Valence-space approach: essentially *all* properties of *all* nuclei

What can we do with this advance?

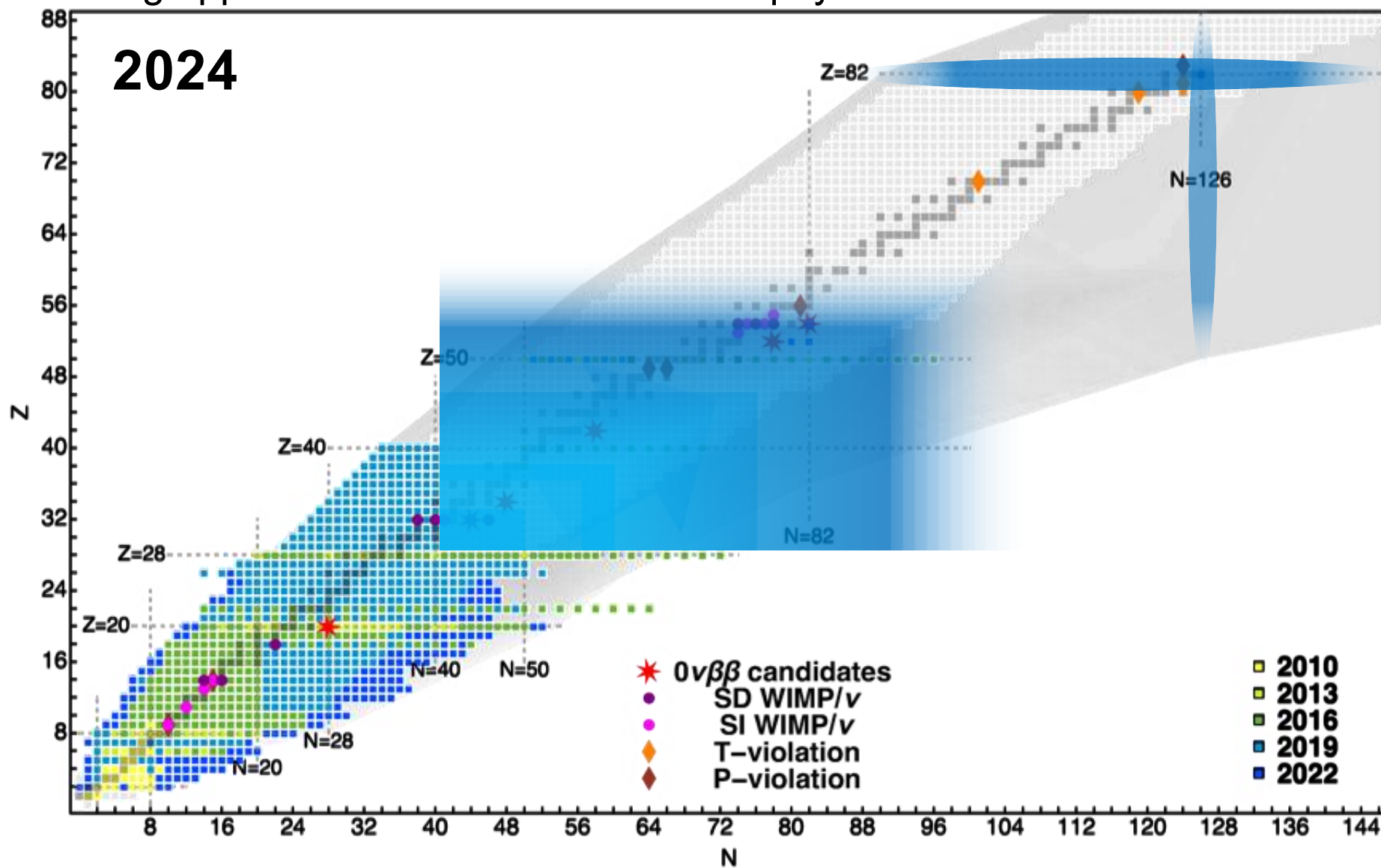


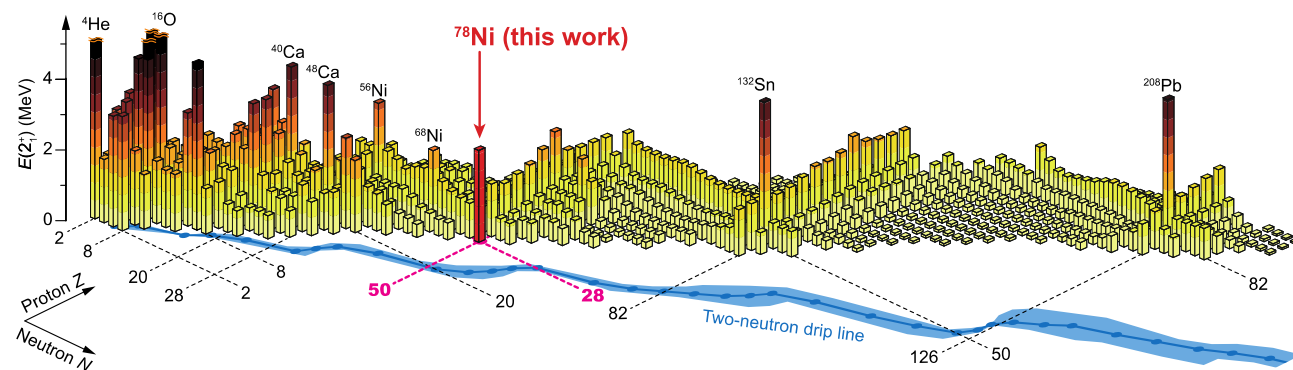
Rapid progress in ab initio reach: **global + heavy**



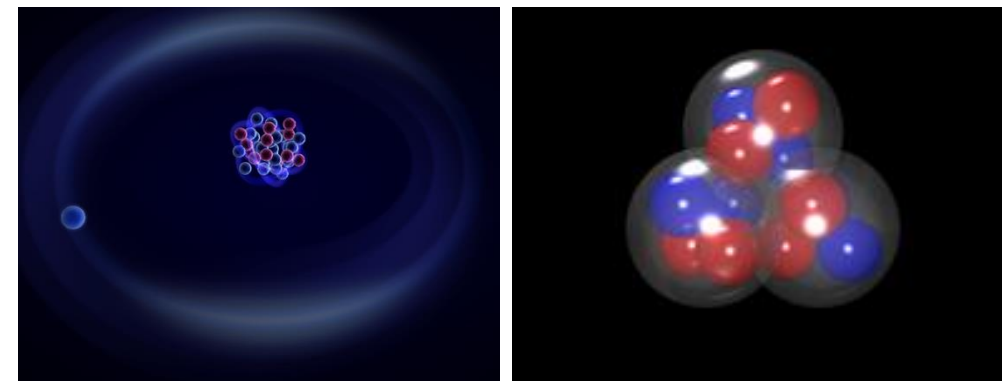
Rapid progress in ab initio reach: **global + heavy**

Exciting applications to searches for BSM physics!

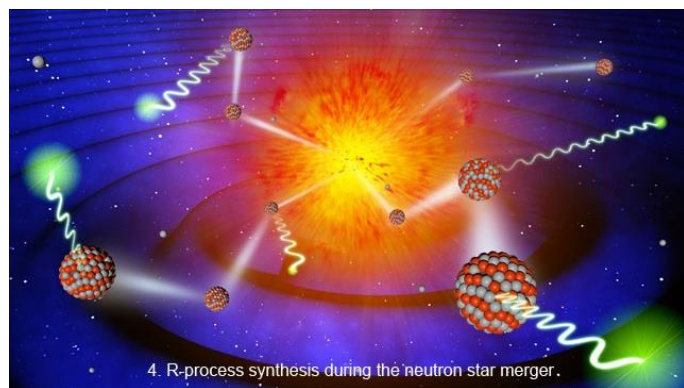




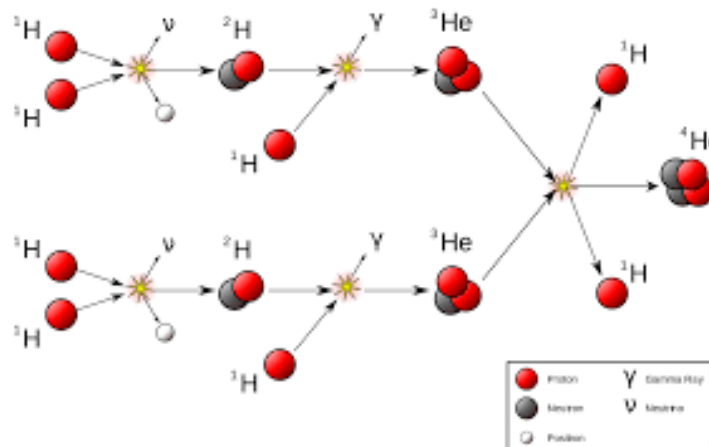
Limits of existence + formation/evolution of magic numbers



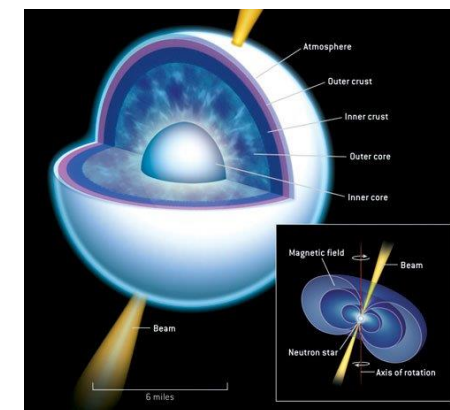
Nuclear skins/halos/clusters



Heavy Element Nucleosynthesis



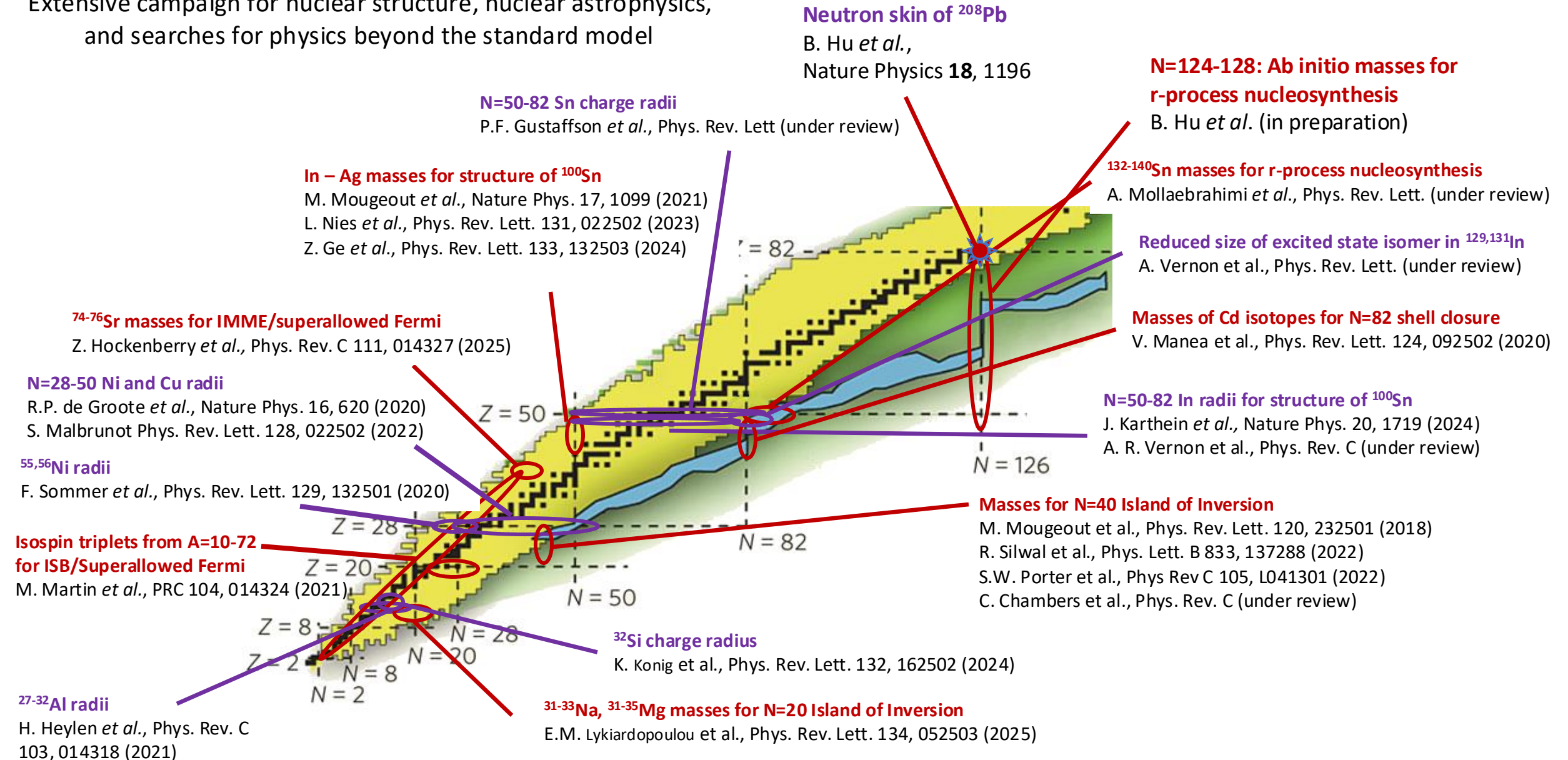
Continuum and nuclear reactions



Infinite matter/Neutron stars

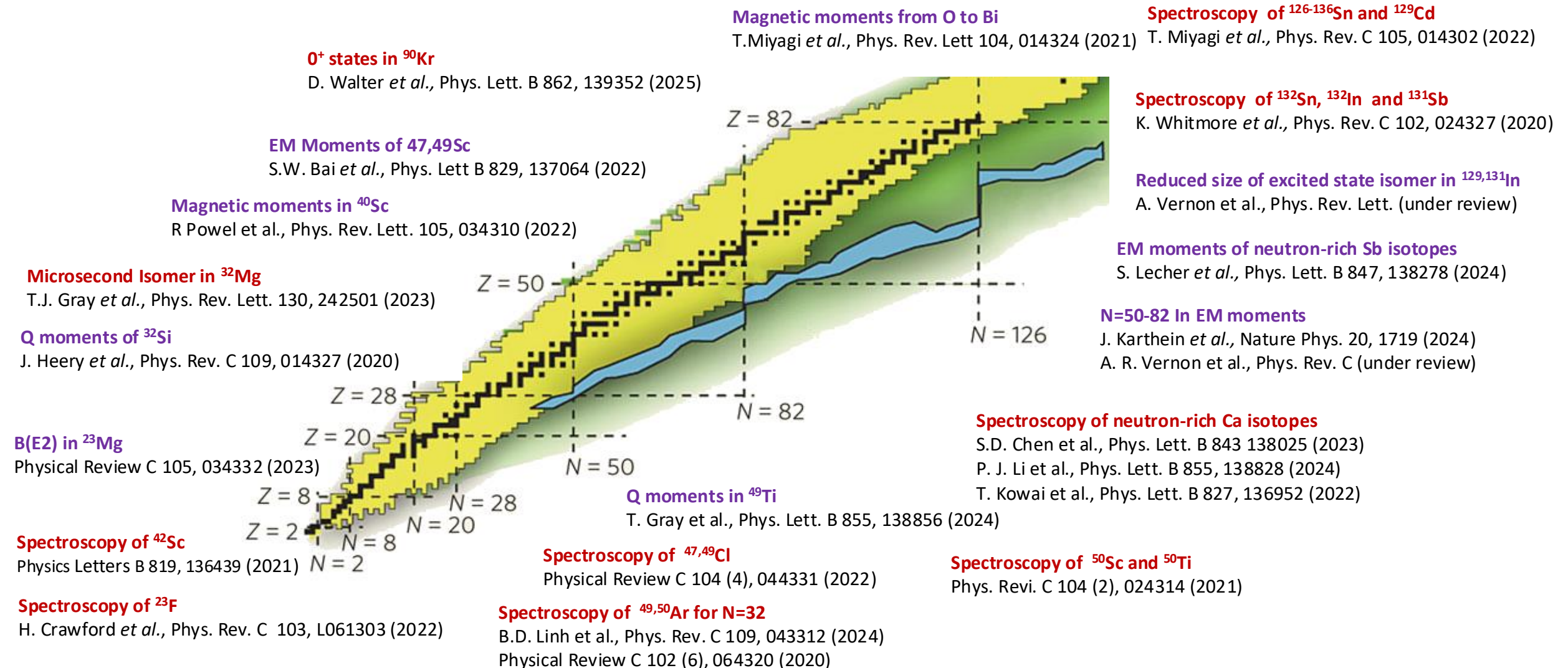
Ab Initio VS-IMSRG Campaign: Masses and Charge Radii

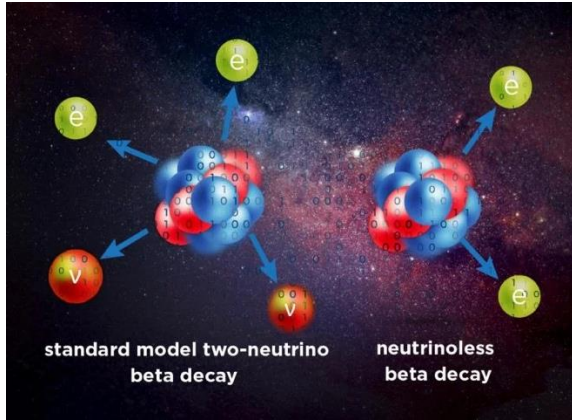
Extensive campaign for nuclear structure, nuclear astrophysics,
and searches for physics beyond the standard model



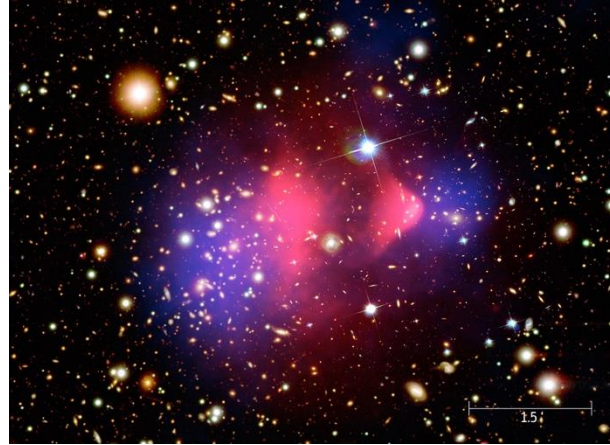
Ab Initio VS-IMSRG Campaign: Spectra and EM Moments

Extensive campaign for nuclear structure, nuclear astrophysics,
and searches for physics beyond the standard model

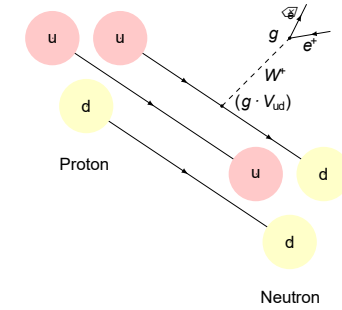




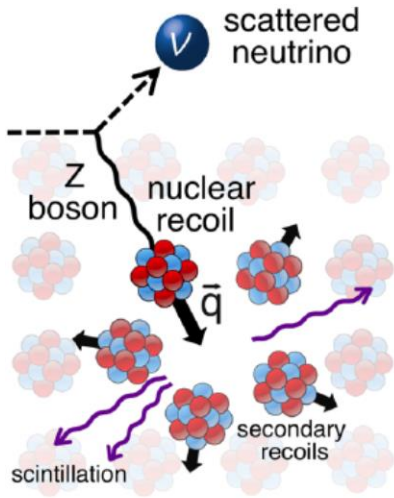
Neutrinoless double-beta decay



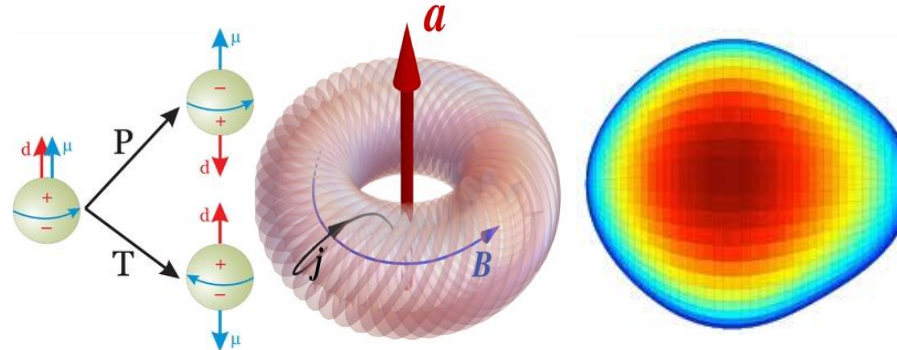
Dark matter direct detection



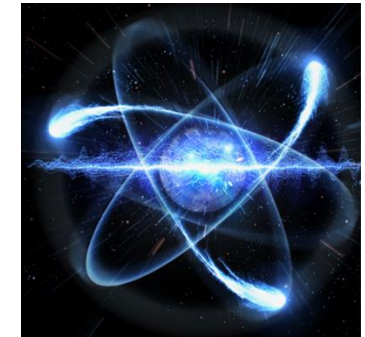
Superallowed Fermi transitions



Neutrino scattering



Symmetry-violating moments



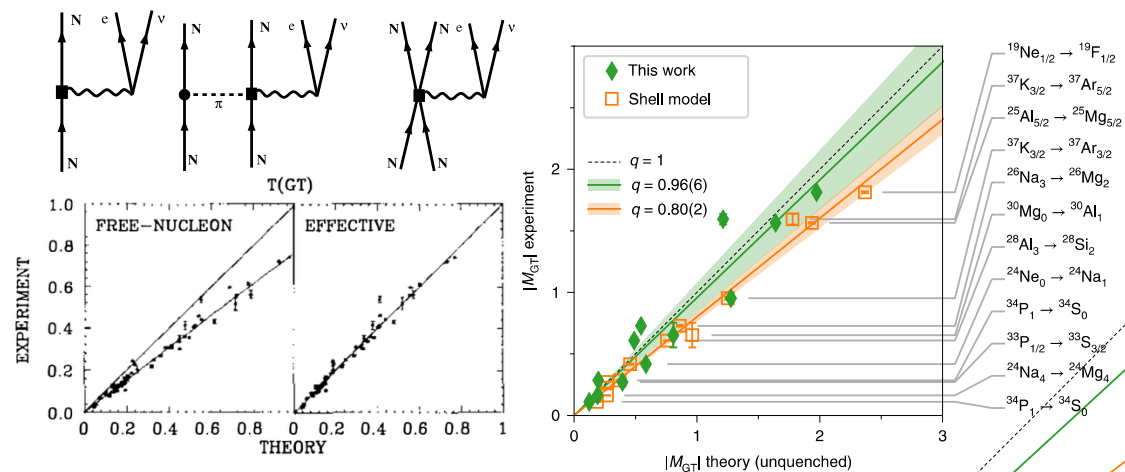
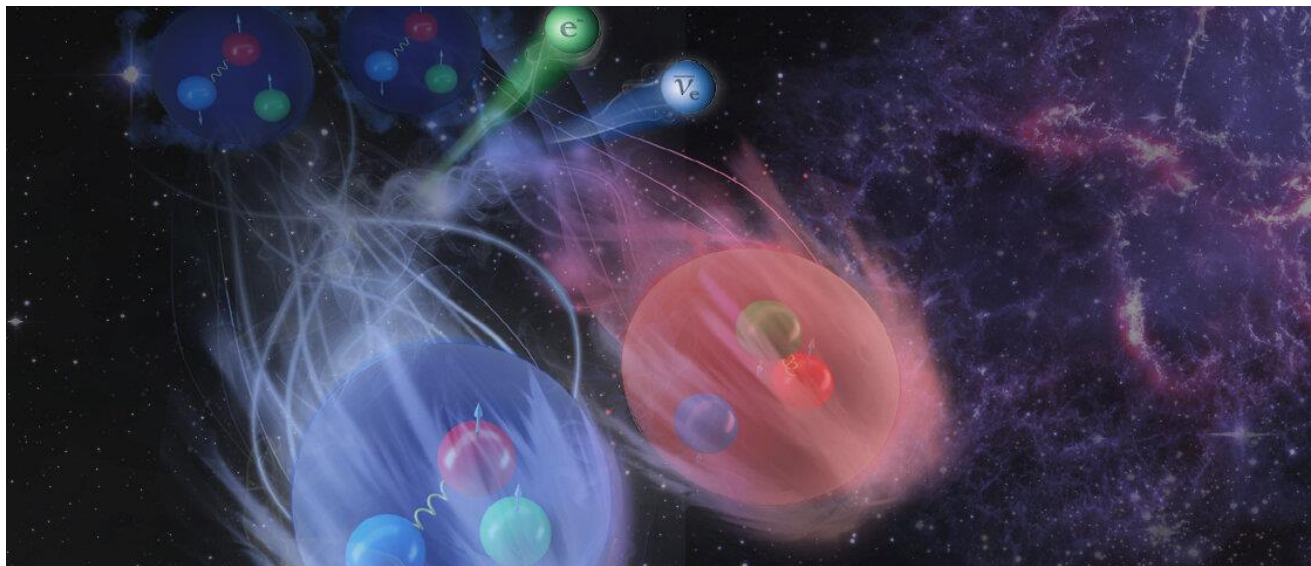
Atomic theory

Two-Body Currents for Gamow-Teller Transitions and g_A Quenching

Ab initio survey of **GT matrix elements**

- Many-body correlations
- Consistent NN+3N forces + 2BC

Broad experimental agreement with $g_A=1.25$



LETTERS

<https://doi.org/10.1038/s41567-019-0450-7>

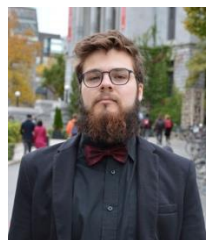
nature
physics

Discrepancy between experimental and theoretical β -decay rates resolved from first principles

P. Gysbers^{1,2}, G. Hagen^{3,4*}, J. D. Holt¹, G. R. Jansen^{3,5}, T. D. Morris^{3,4,6}, P. Navrátil¹, T. Papenbrock^{3,4}, S. Quaglioni⁷, A. Schwenk^{8,9,10}, S. R. Stroberg^{1,11,12} and K. A. Wendt⁷

$0\nu\beta\beta$ Decay for All Major Players: ^{76}Ge

^{100}Mo , ^{130}Te , ^{136}Xe



A. Belley



L. Jokiniemi



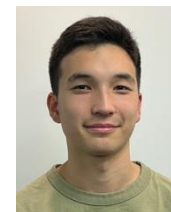
I. Ginnett



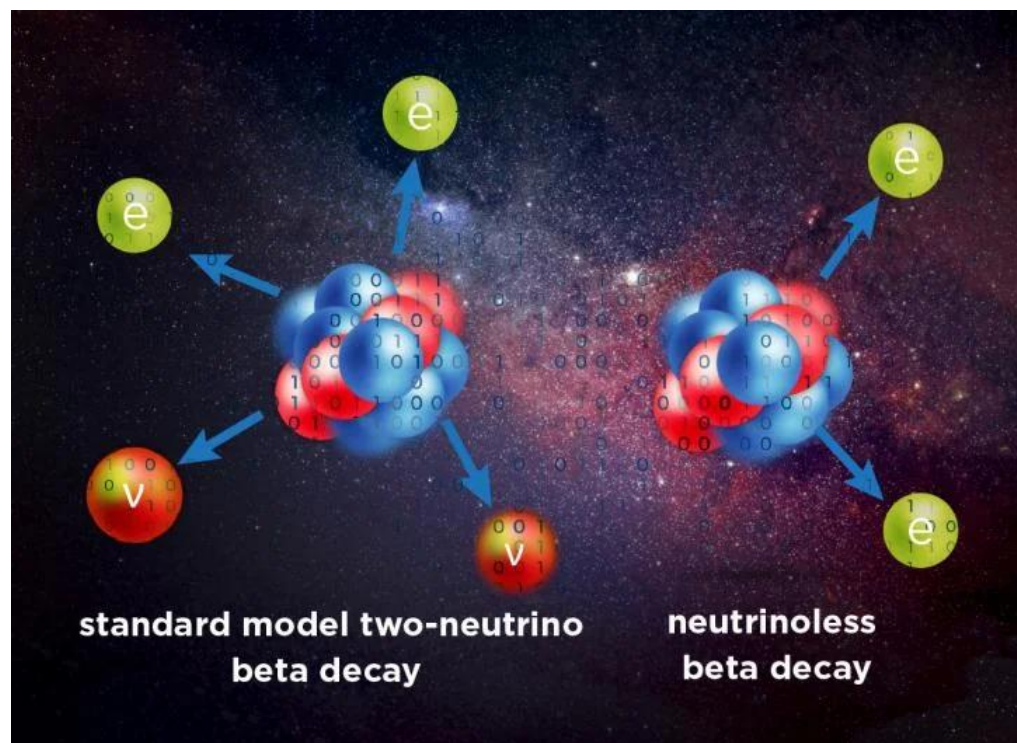
J. Pitcher



A. Todd



T. Shickele



Ab Initio Neutrinoless Double-Beta Decay Matrix Elements for ^{48}Ca , ^{76}Ge , and ^{82}Se

A. Belley, C. G. Payne, S. R. Stroberg, T. Miyagi, and J. D. Holt
Phys. Rev. Lett. **126**, 042502 – Published 29 January 2021

Ab initio uncertainty quantification of neutrinoless double-beta decay in Ge

Phys. Rev. Lett.

A. Belley, J. M. Yao, B. Bally, J. Pitcher, J. Engel, H. Hergert, J. D. Holt, T. Miyagi, T. R. Rodríguez, A. M. Romero, S. R. Stroberg, and X. Zhang

Physical Sciences - Article

Ab initio calculations of neutrinoless $\beta\beta$ decay refine neutrino mass limits

Status: Under Review

Jason Holt, Antoine Belley, Takayuki Miyagi, Steven Stroberg

nature portfolio

Proper renormalization requires short-range contact term at leading order

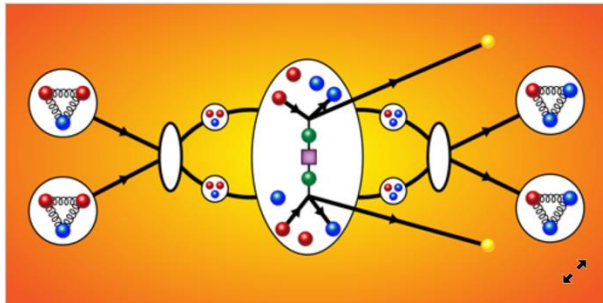
Physics

SYNOPSIS

A Missing Piece in the Neutrinoless Beta-Decay Puzzle

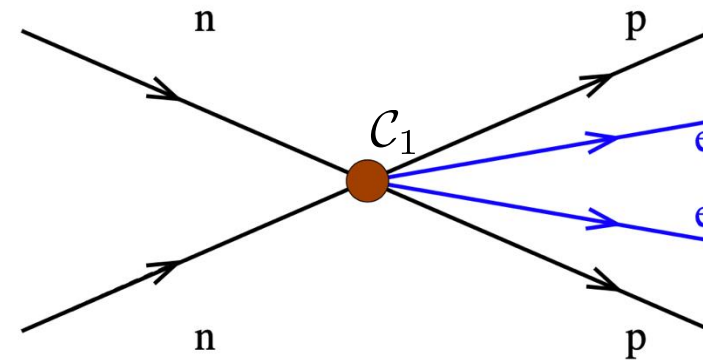
May 16, 2018 • Physics 11, s58

The inclusion of short-range interactions in models of neutrinoless double-beta decay could impact the interpretation of experimental searches for the elusive decay.



J. de Vries/Nikhef; adapted by APS/Alan Stonebraker

V. Cirigliano et al. PRL (2018), PRL (2022)



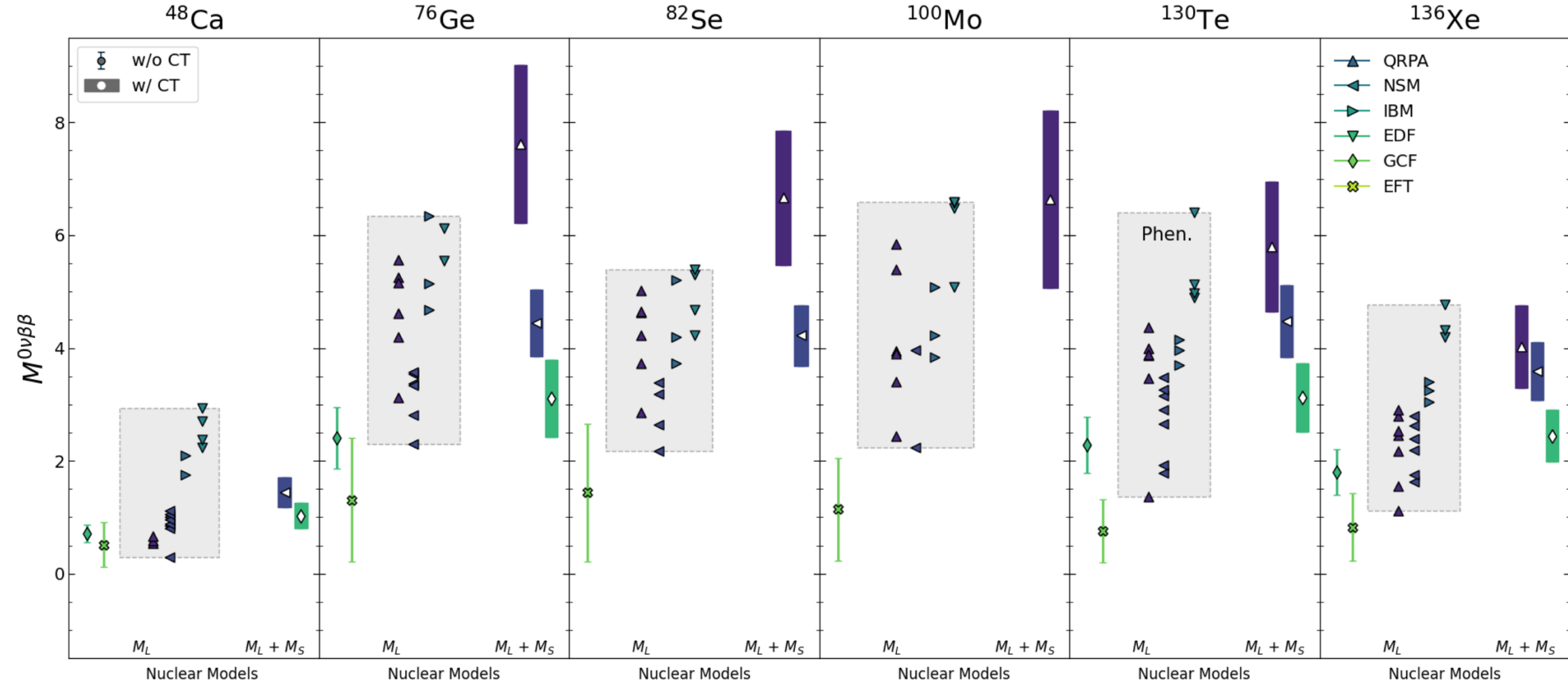
New physics inside blob:
High-energy ν exchange

New paradigm for $0\nu\beta\beta$ decay: include long- and short-range terms

$$M^{0\nu} \rightarrow M_L + M_S = M_{GT} + \frac{M_F}{g_A^2} + M_T + M_{CT}$$

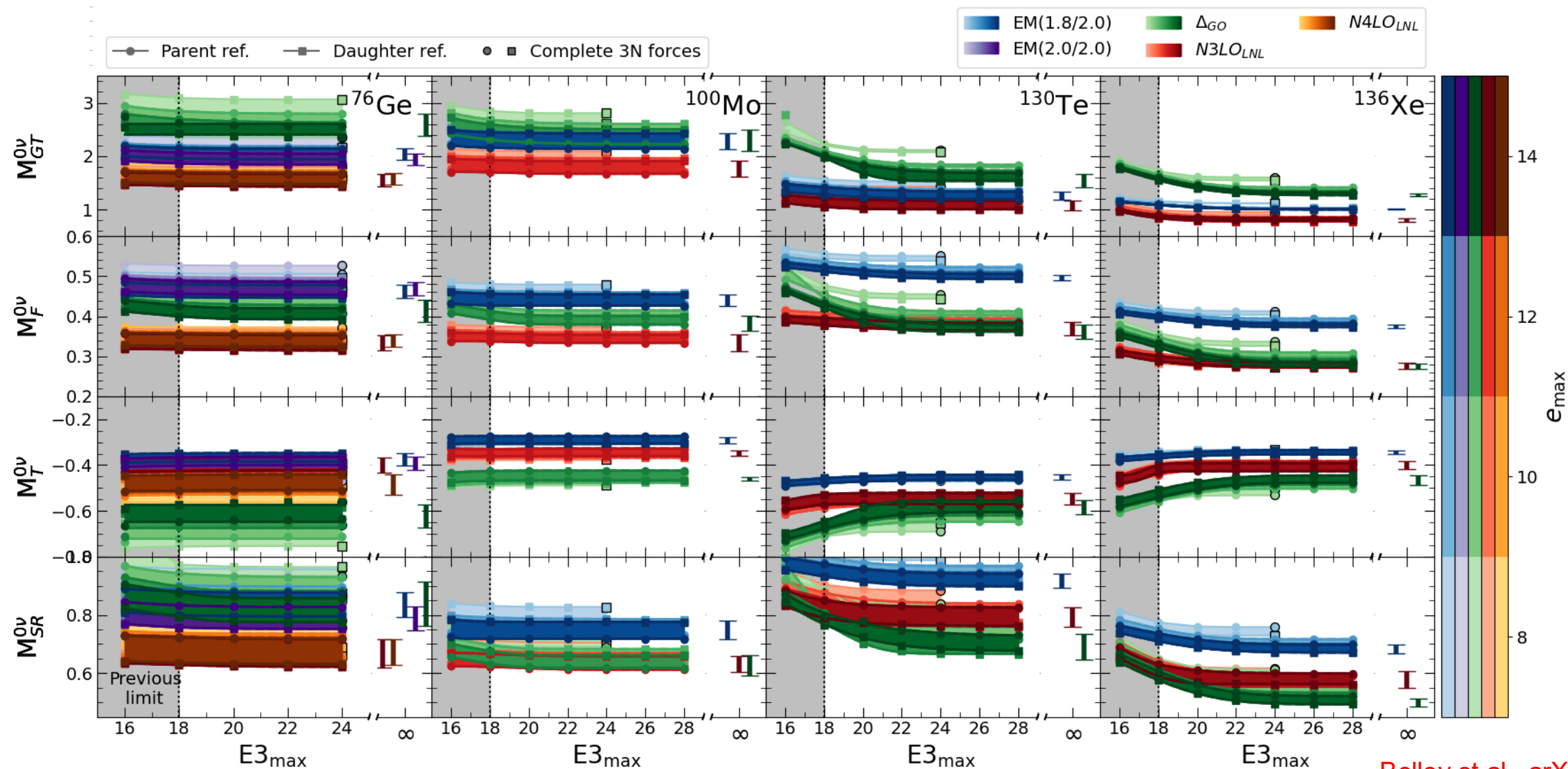
Estimated to ~30% - no consistent way to assess with phenomenology

Compiled values updated w/ phenomenology + short-range contact



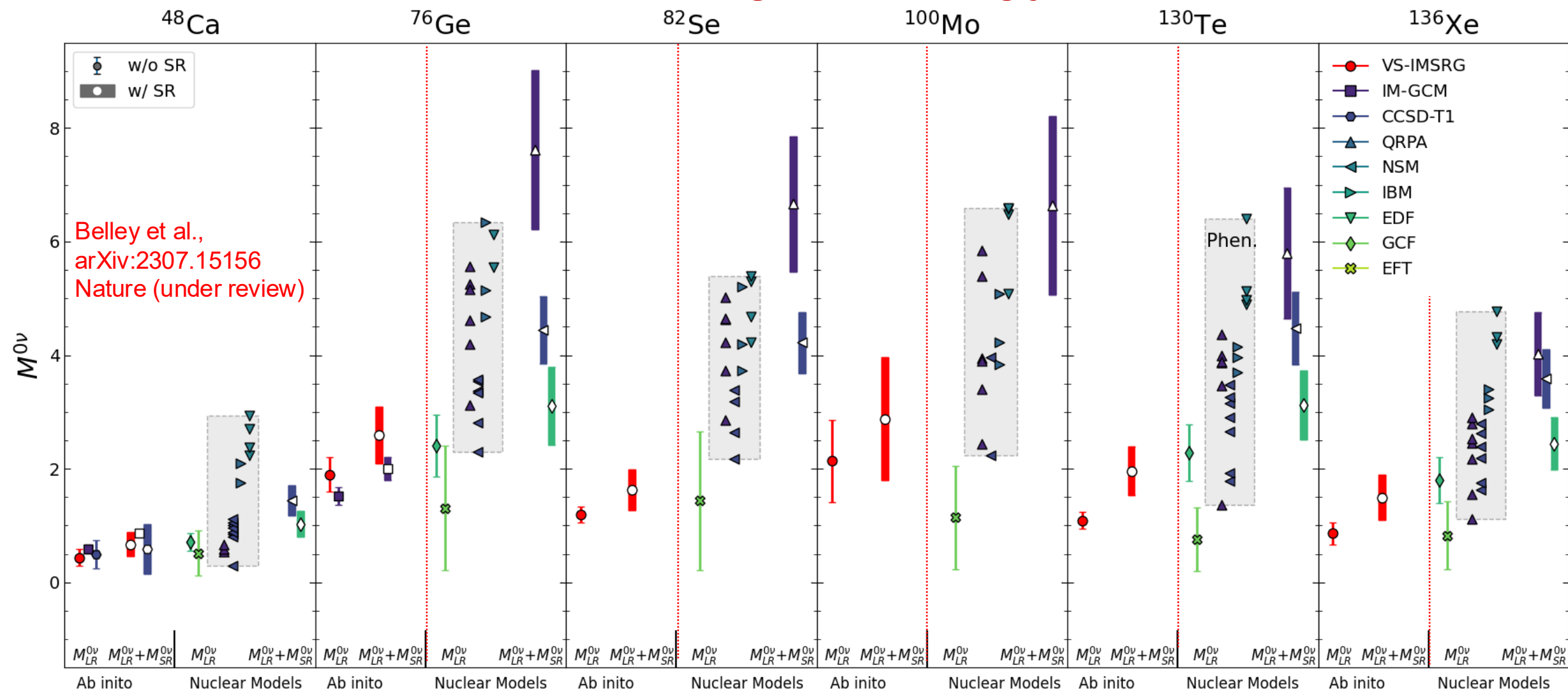
New(er) results increase factors up to >10 uncertainty 😬 can we do better?

Converged NMEs for major players in global searches: ^{76}Ge , ^{100}Mo , ^{130}Te , ^{136}Xe



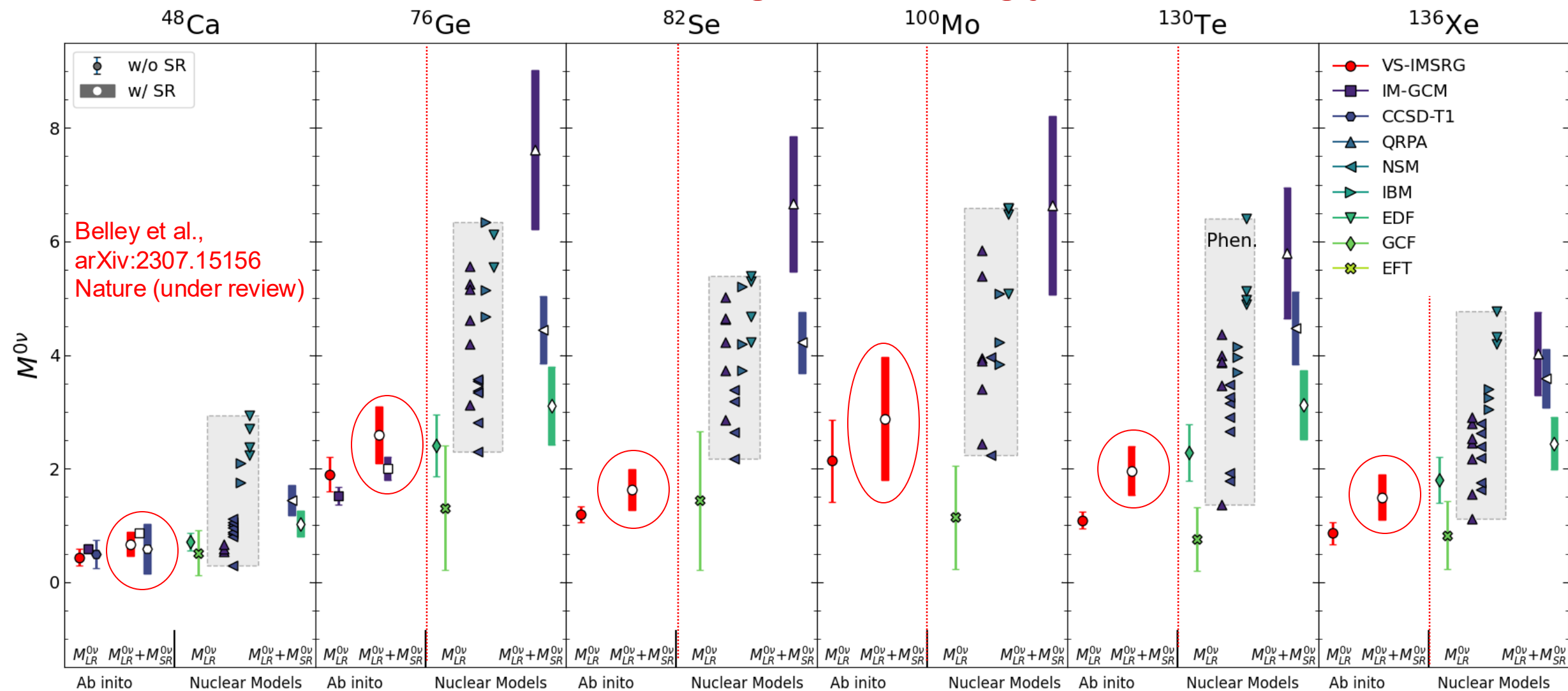
Converged NMEs for major players in global searches: ^{76}Ge , ^{100}Mo , ^{130}Te , ^{136}Xe

Ab initio results: differences from models; **large NMEs strongly disfavored**



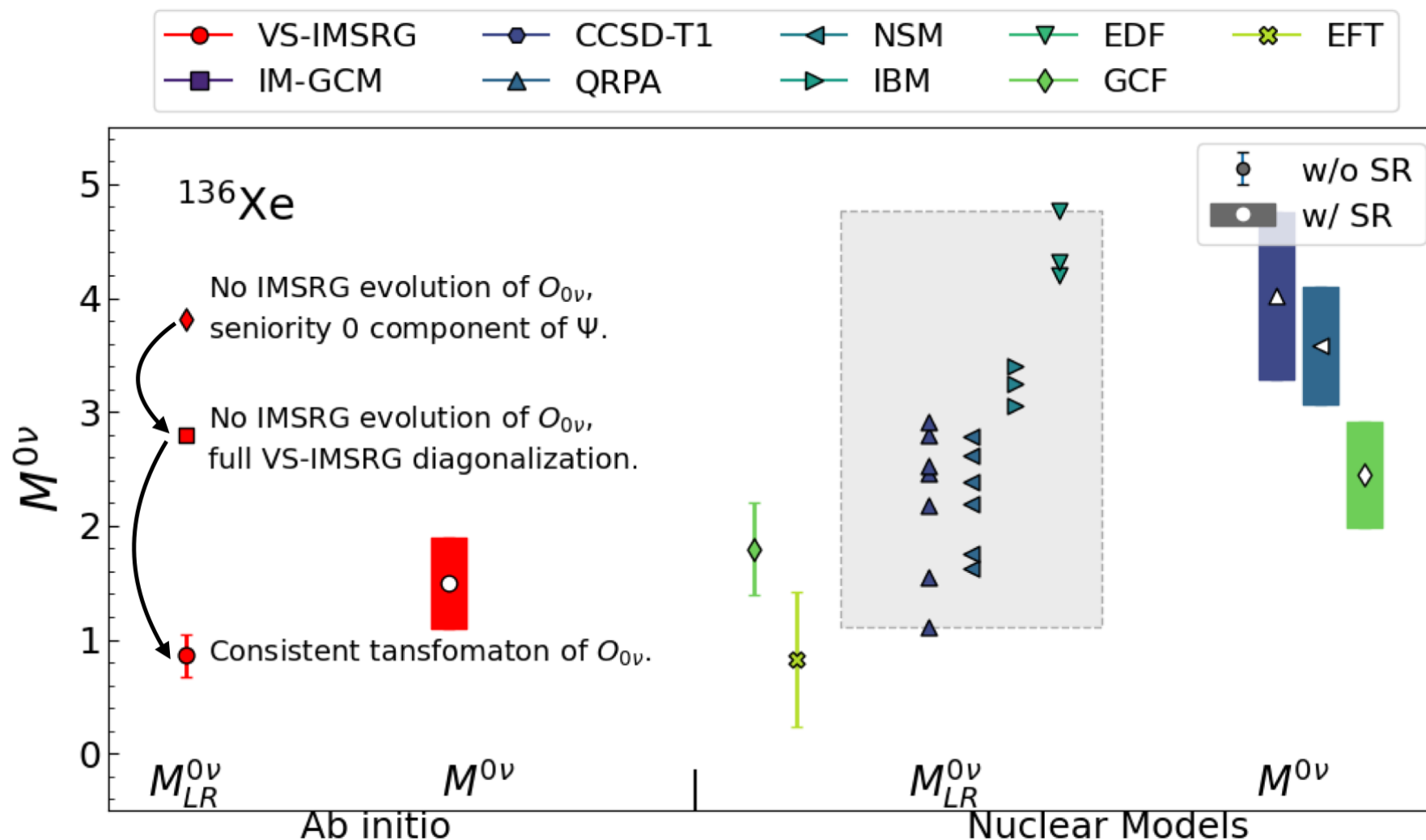
Converged NMEs for major players in global searches: ^{76}Ge , ^{100}Mo , ^{130}Te , ^{136}Xe

Ab initio results: differences from models; **large NMEs strongly disfavored**



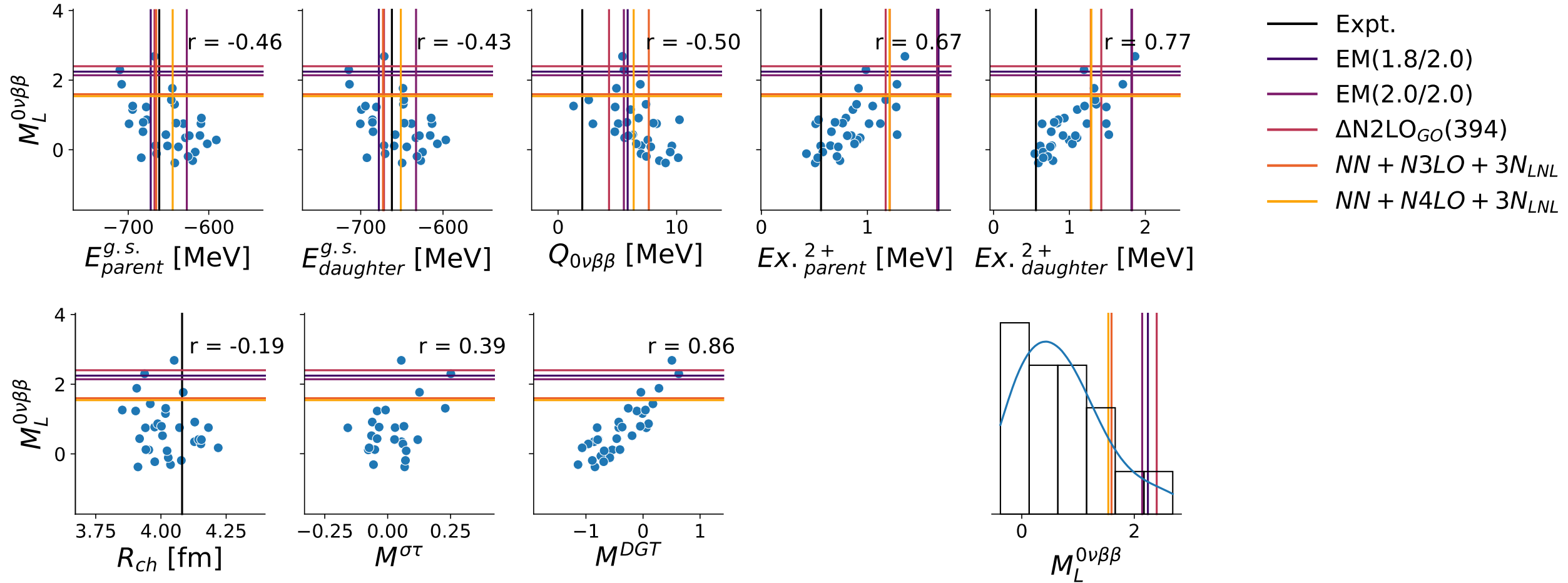
Correlations in wavefunction and operator drive small NMEs

Ab initio results: differences from models; **large NMEs strongly disfavored**



^{76}Ge : Explore correlations with other observables from 34 interactions

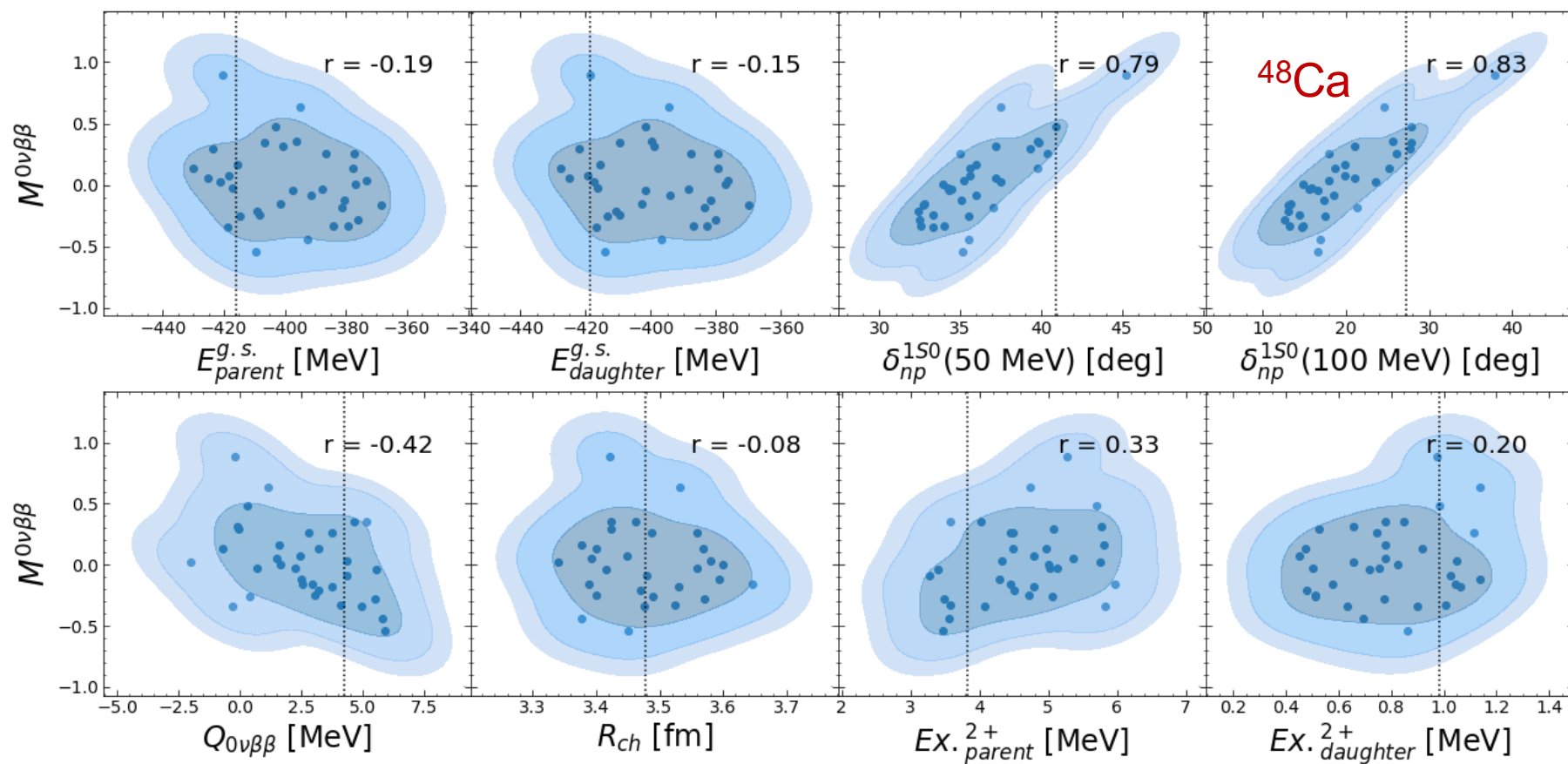
No obvious correlations, except DGT



Use machine learning for further insights?

^{76}Ge : Explore correlations with other observables from 34 interactions

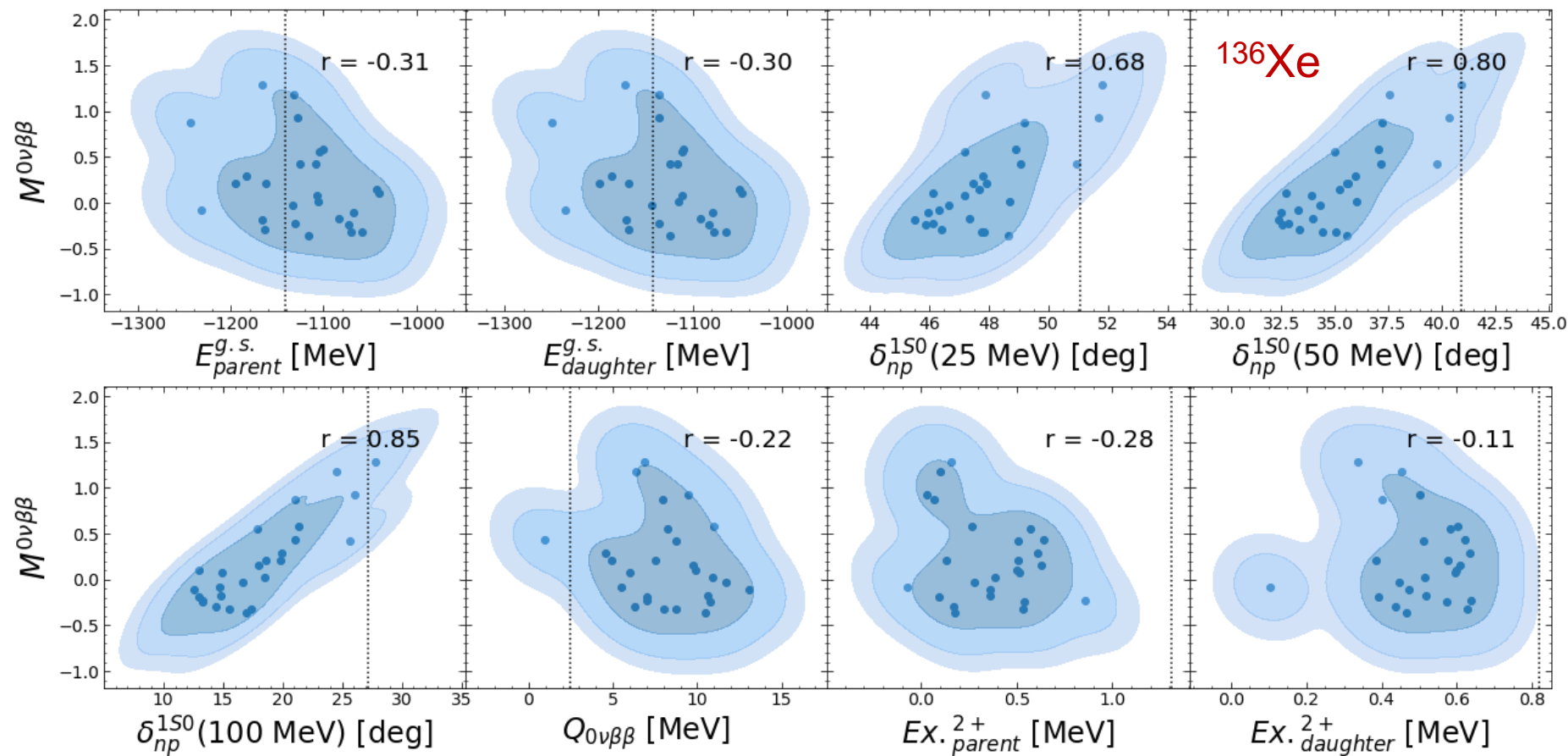
No ab initio correlations, except DGT



Use machine learning for further insights?

^{76}Ge : Explore correlations with other observables from 34 interactions

No ab initio correlations, except DGT



Use machine learning for further insights?

Explore dependence on chiral EFT LECs: requires many samples (as in ^{208}Pb)

Use gaussian processes as an emulator

Multi-Fidelity Gaussian Process: connects few (complicated) high-fidelity data points (e.g., full IMSRG) w/ many low-fidelity data points (HF, low e_{max} , etc)

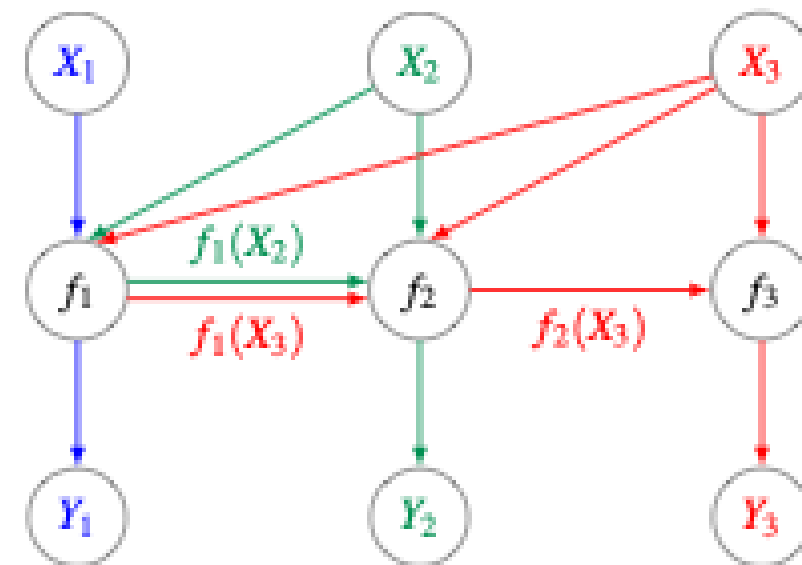
Difference function fit with Gaussian process: predict HF from LF

Deep Gaussian Process: Neural network links multiple GP

Include outputs of previous fidelity as new HF point:
Improves modeling of difference between LF and HF

Adapted for multi output:

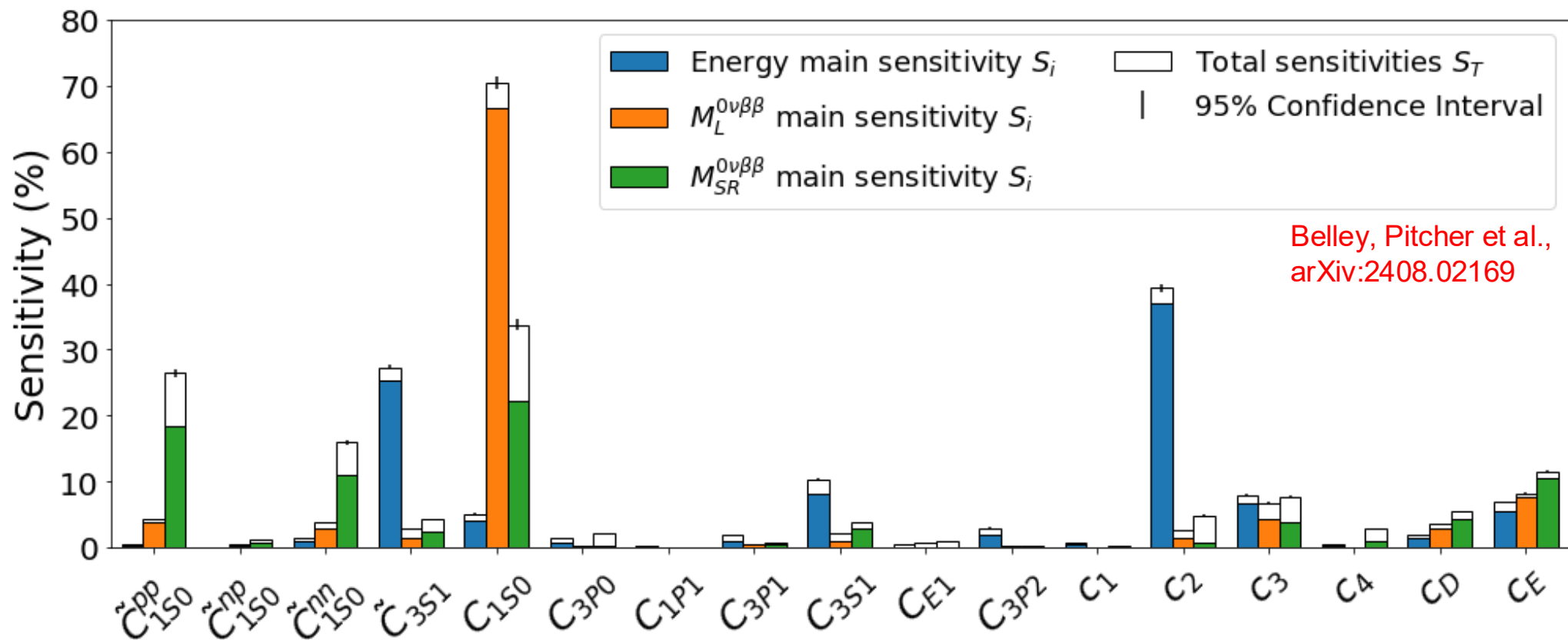
Multi-Output Multi-Fidelity Deep Gaussian Process (MM-DGP)



Belley, Pitcher et al., in preparation

Newly developed MM-DGP data-driven emulator for VS-IMSRG

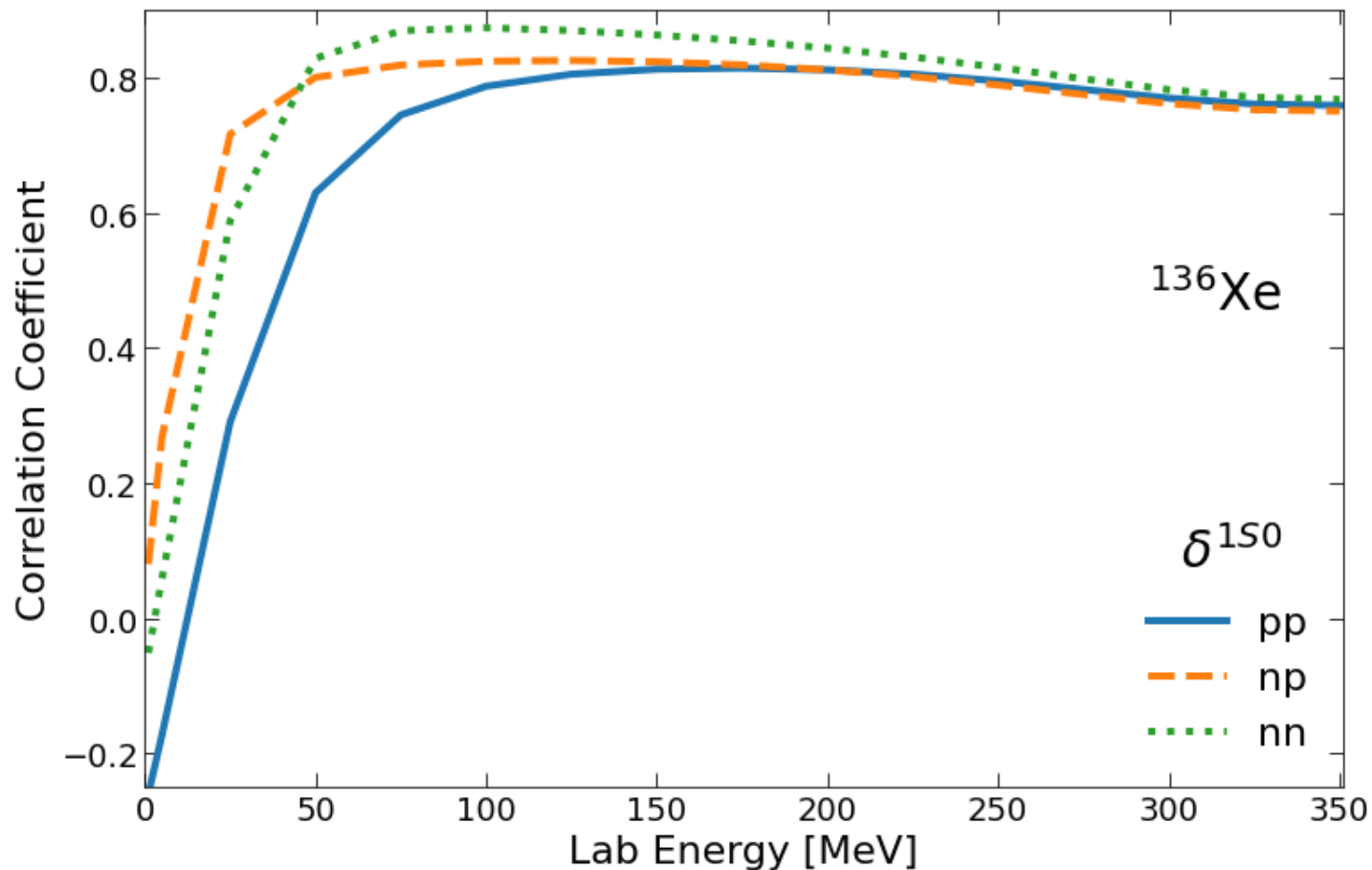
Sensitivity analysis: Connect long-range $0\nu\beta\beta$ decay to terms in nuclear forces



Highly sensitive to single constant "C150" – related to 1S_0 phase shift

Clear correlation with 1S_0 phase-shift

Strong correlation emerges for energies $> 50\text{MeV}$

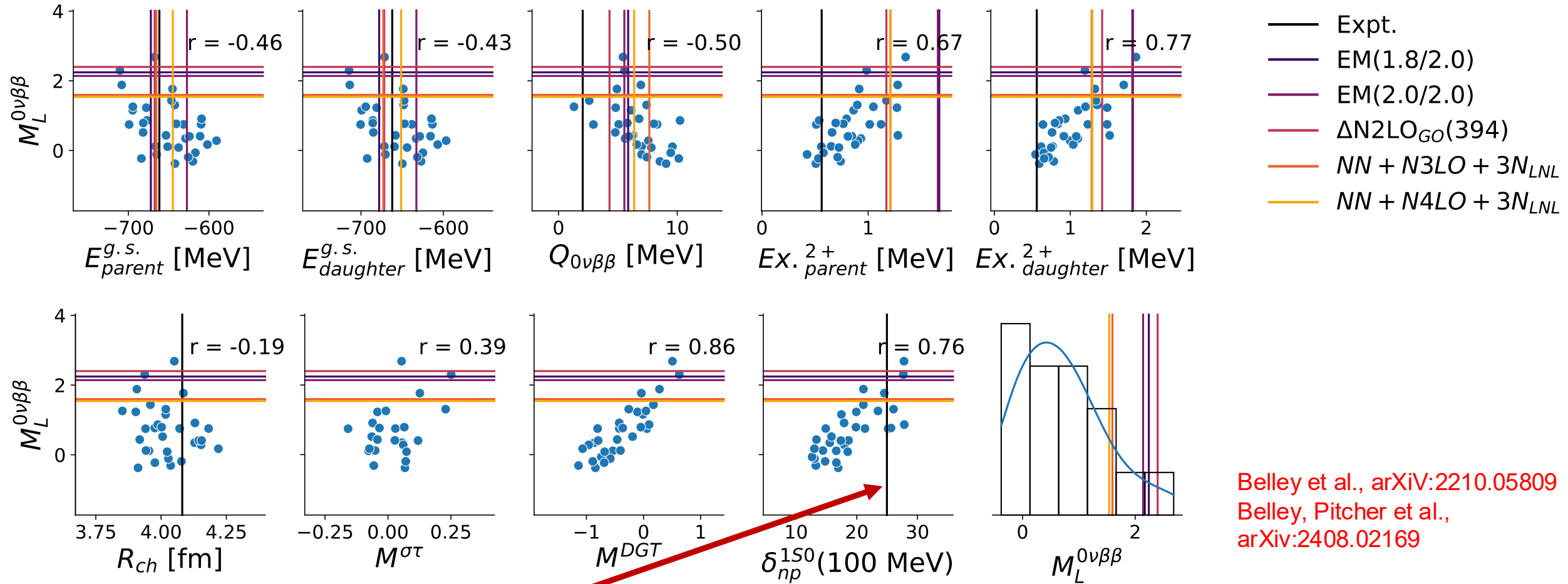


Belley, Pitcher et al.,
arXiv:2408.02169

Constrained by part of interaction for the two nucleons when “close” to each other

Explore correlations with other observables from systematic analysis (34 interactions)

Few clear correlations, except DGT in ^{76}Ge



Belley et al., arXiv:2210.05809
 Belley, Pitcher et al.,
 arXiv:2408.02169

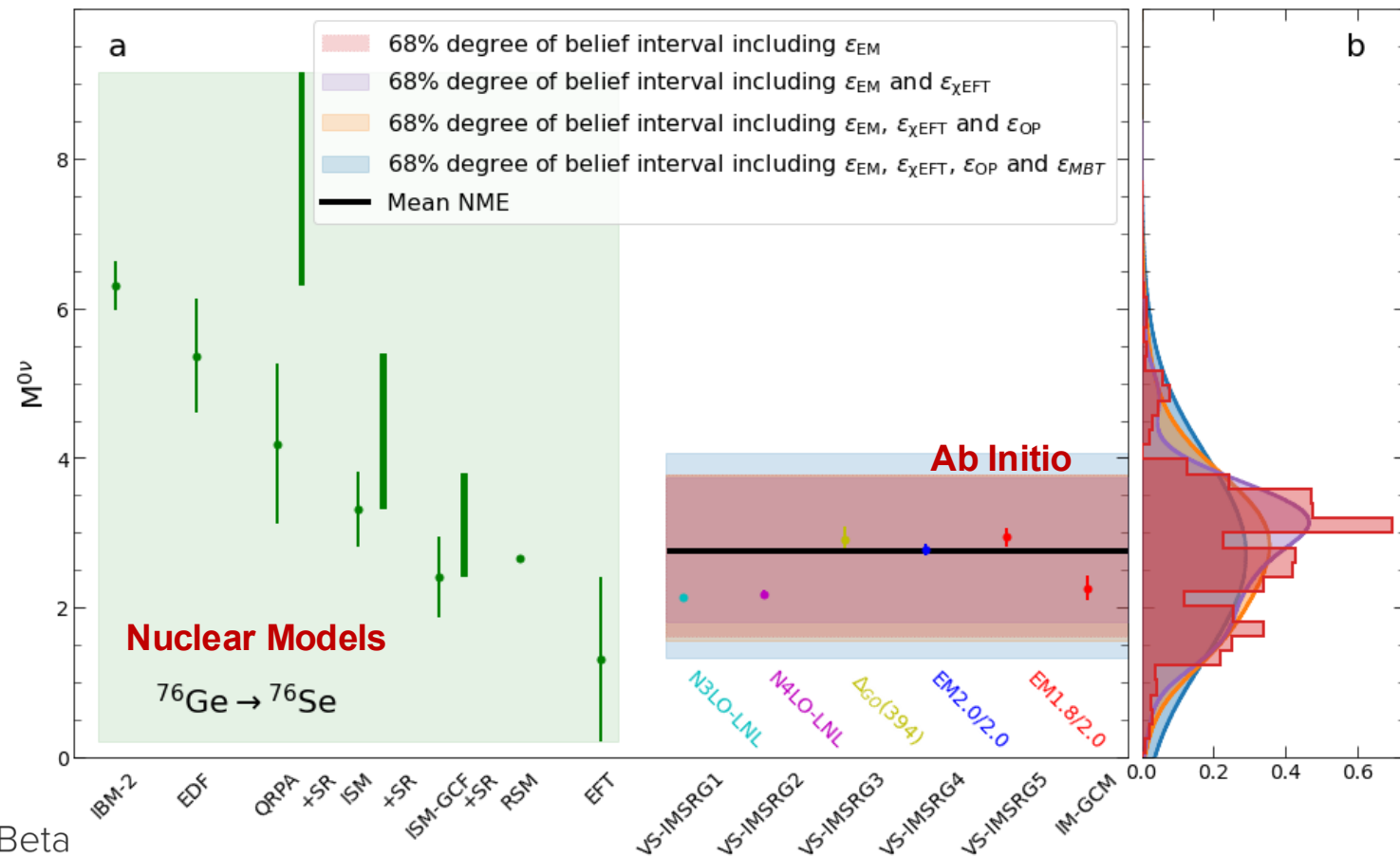
Consider *all* physics contributing to theoretical error in $0\nu\beta\beta$ decay

$$y = y_{\text{MM-DDGP}} + \epsilon_{\text{emulator}} + \epsilon_{\text{EFT}} + \epsilon_{\text{many-body}} + \epsilon_{\text{operator}}$$

Nuclear many-body problem treated
w/ newly developed ML emulators

Construct PPD for NME in ^{76}Ge

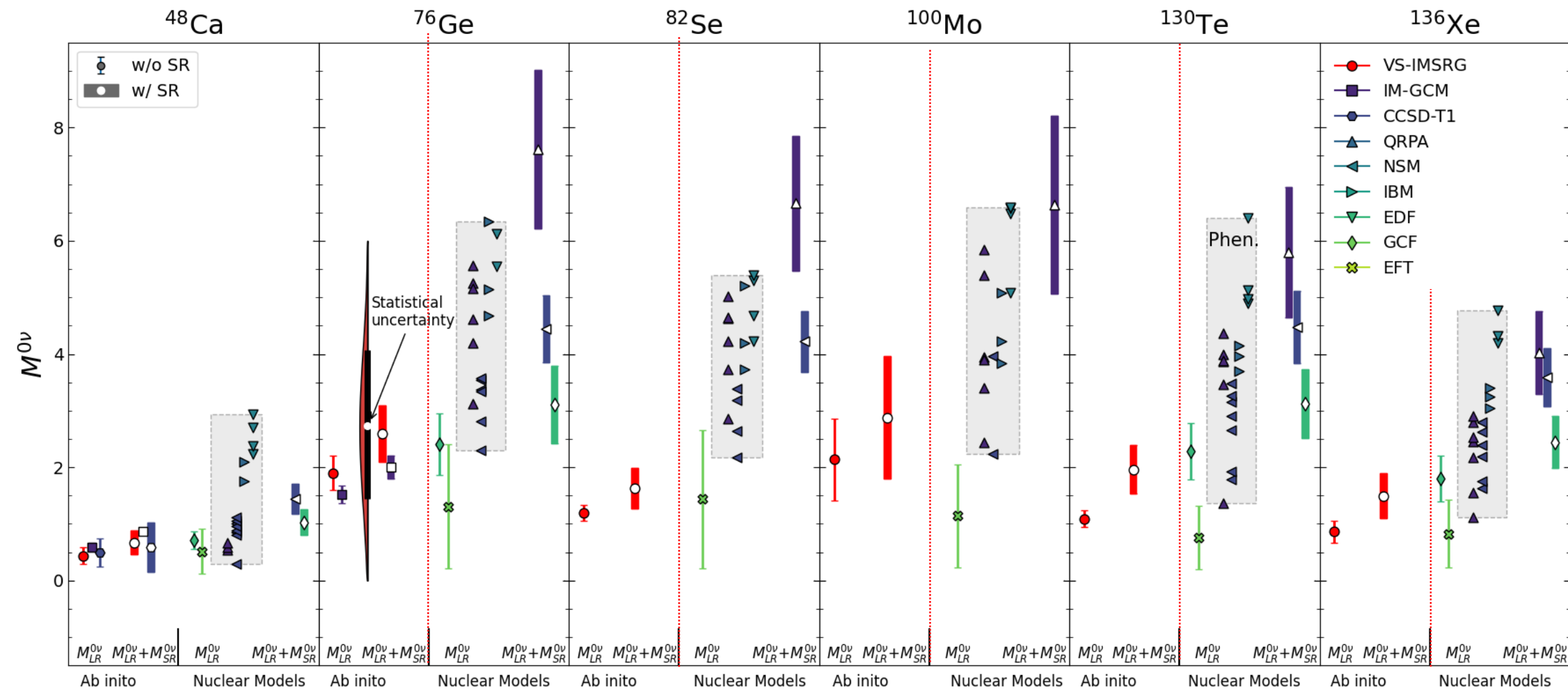
**Dramatically reduced uncertainty
compared to nuclear models**



Ab Initio Uncertainty Quantification of Neutrinoless Double-Beta Decay in ^{76}Ge

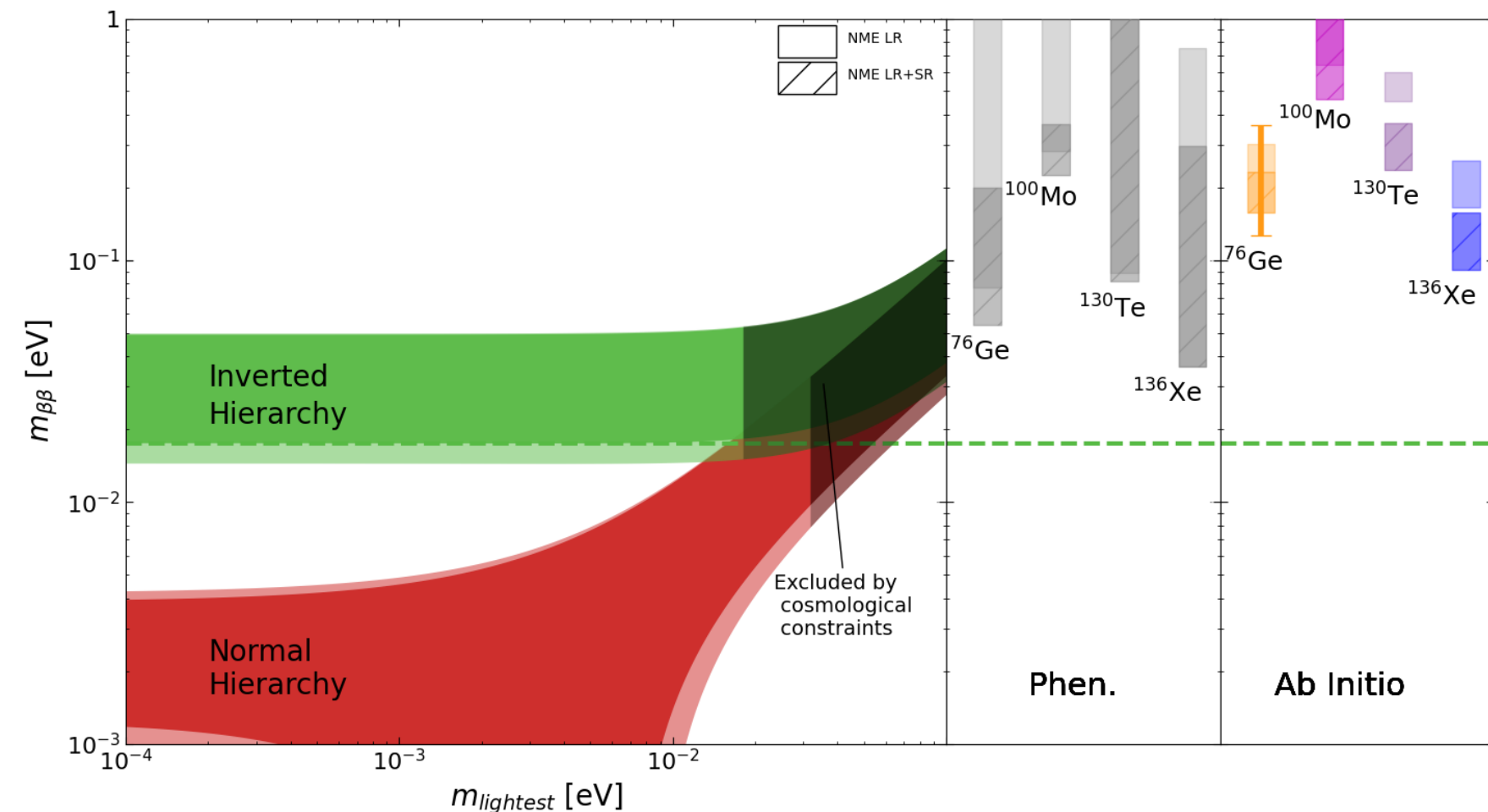
Converged NMEs for major players in global searches: ^{76}Ge , ^{100}Mo , ^{130}Te , ^{136}Xe

Independent PPD agrees with previous spread



Impact for current searches: large matrix elements disfavored, lowers expected rates

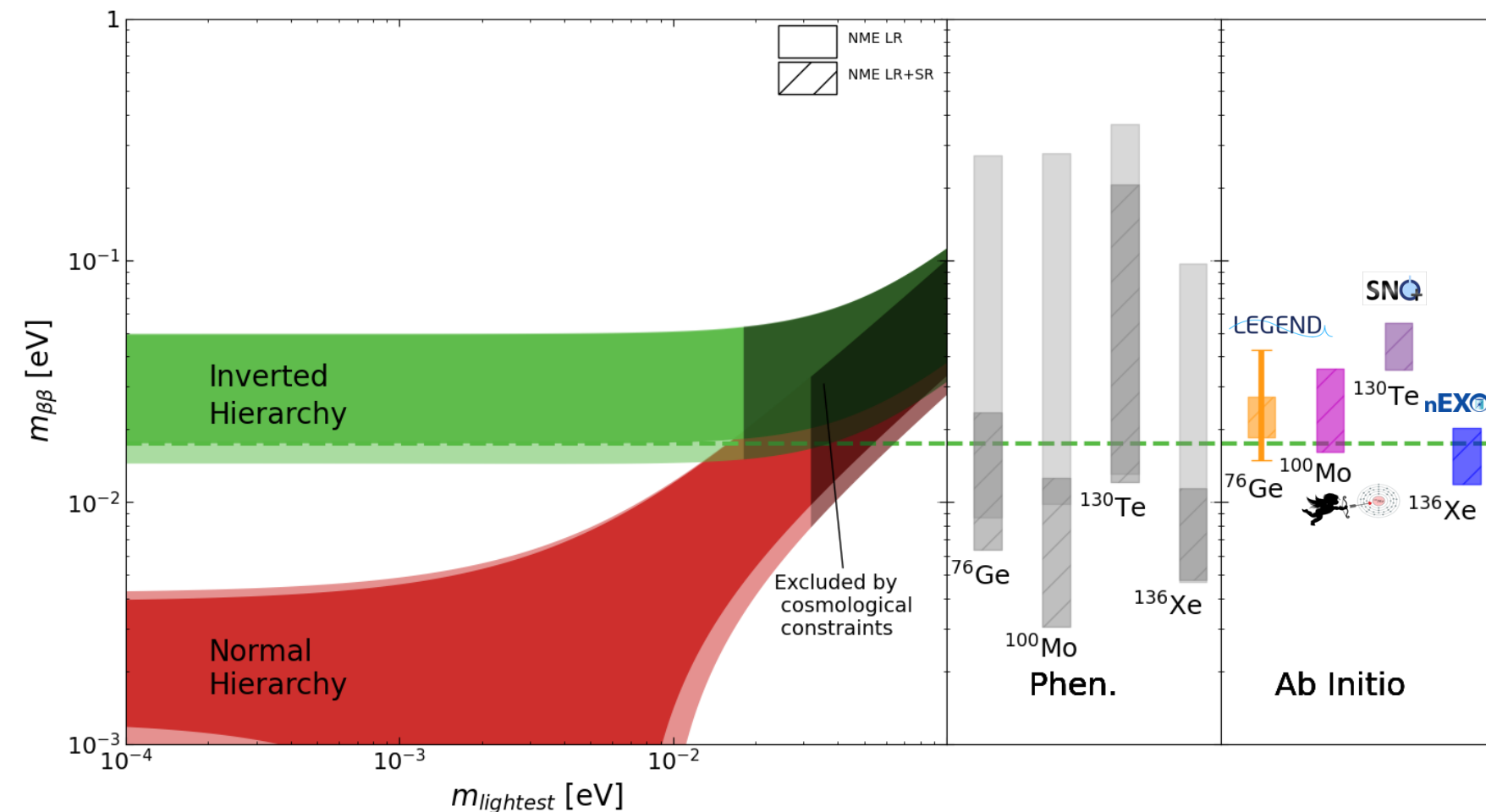
Current experimental reach – improved with effects of contact term



Highly unlikely inverted Hierarchy has been probed

Impact for next-generation searches: sensitivities from LEGEND, SNO+, nEXO, CUPID

Effect of short-range term improves experimental reach... **¹⁰⁰Mo now a major player**

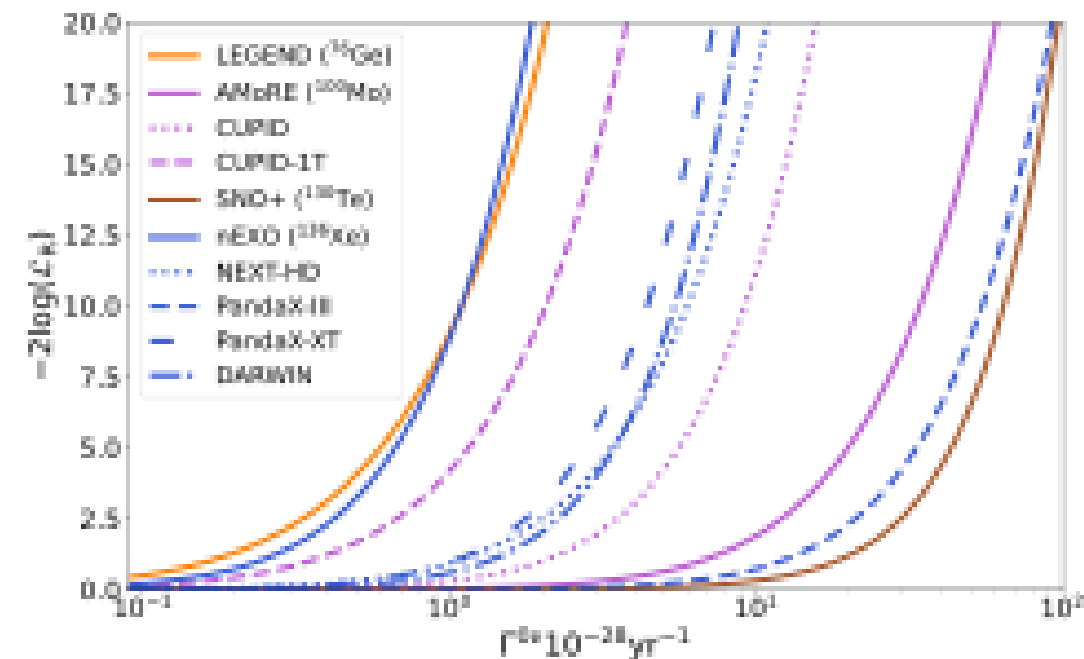
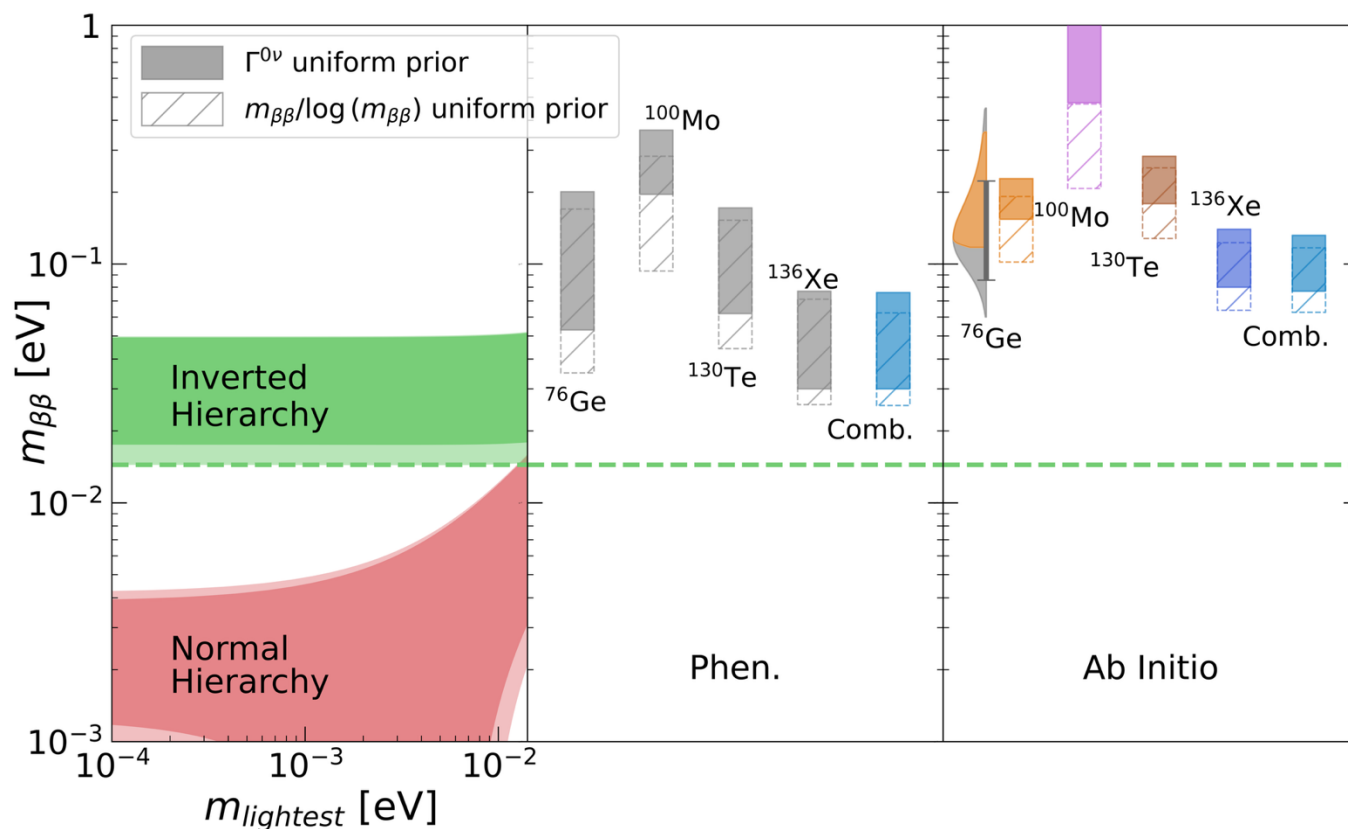


Uncertainty reduced by **over one order of magnitude!**

Obtain likelihood functions from experiments – convert decay rates to $m_{\beta\beta}$

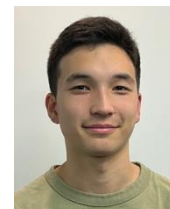
$$m_{\beta\beta} = \frac{m_e}{g_A^2 |M^{0\nu}|} \sqrt{\frac{\Gamma^{0\nu}}{G^{0\nu} \log 2}}$$

Likelihoods allow combined limit via Bayes' Theorem



Analysis based on S. Biller PRD (2021)

Combined ab initio limits:
IH fully probed with multiple isotopes

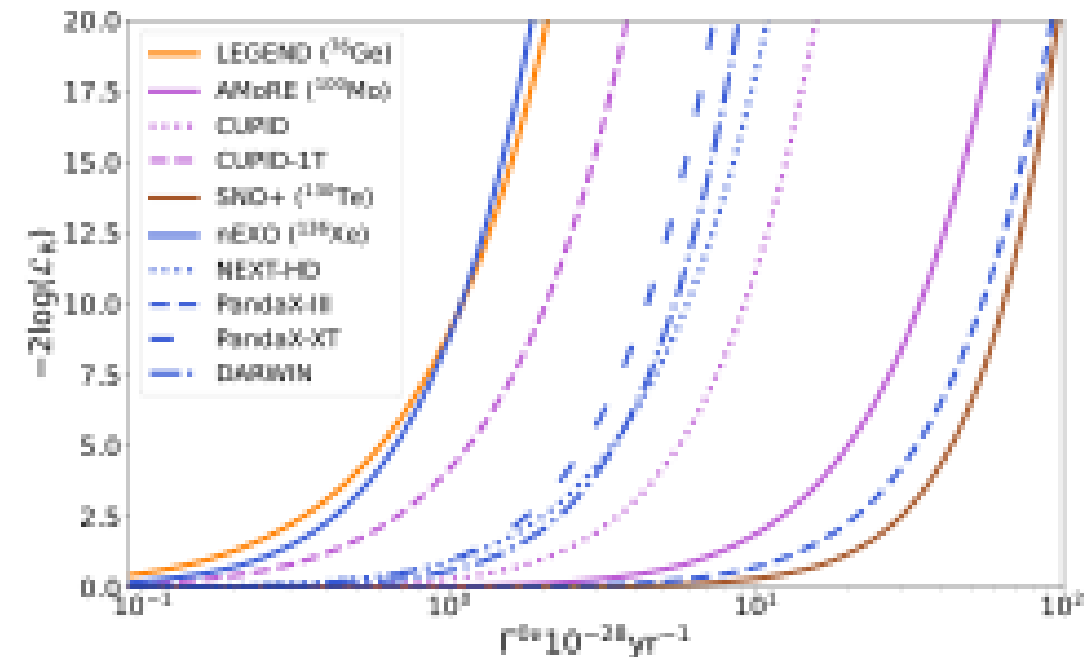
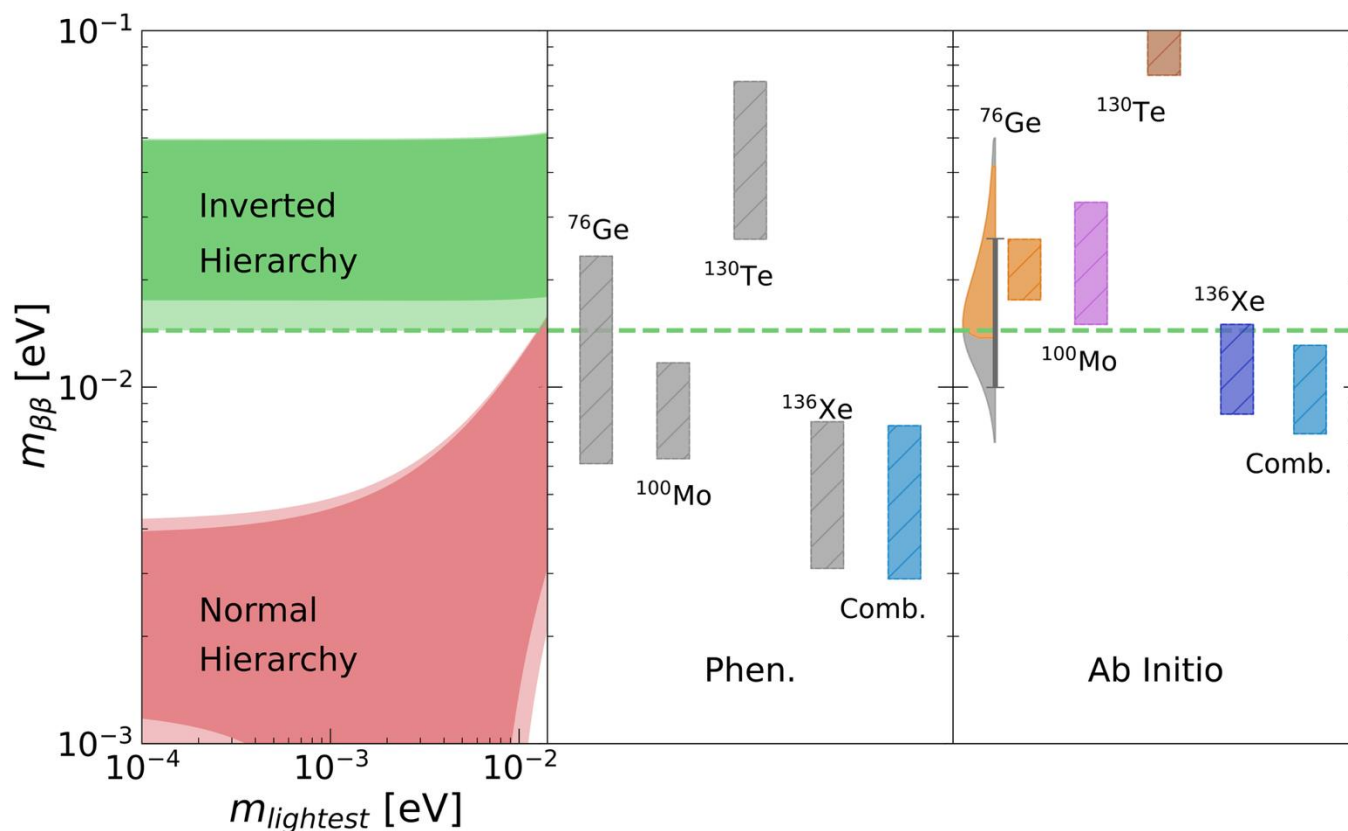


Shickele et al., in prep

Obtain likelihood functions from experiments – convert decay rates to $m_{\beta\beta}$

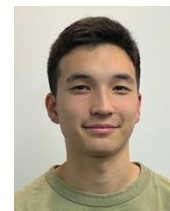
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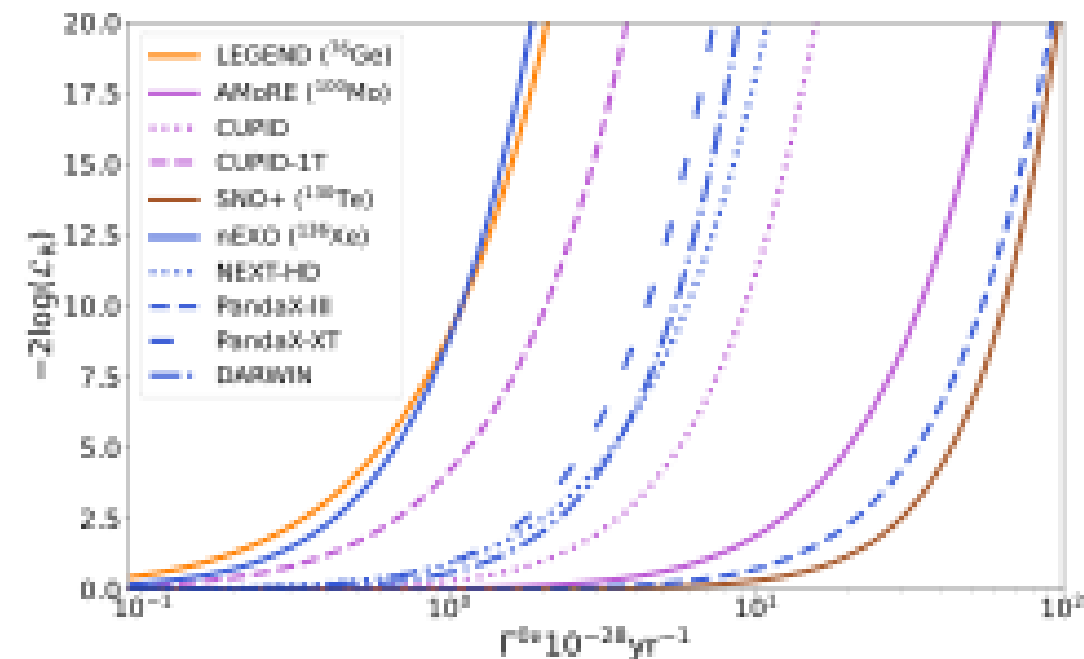
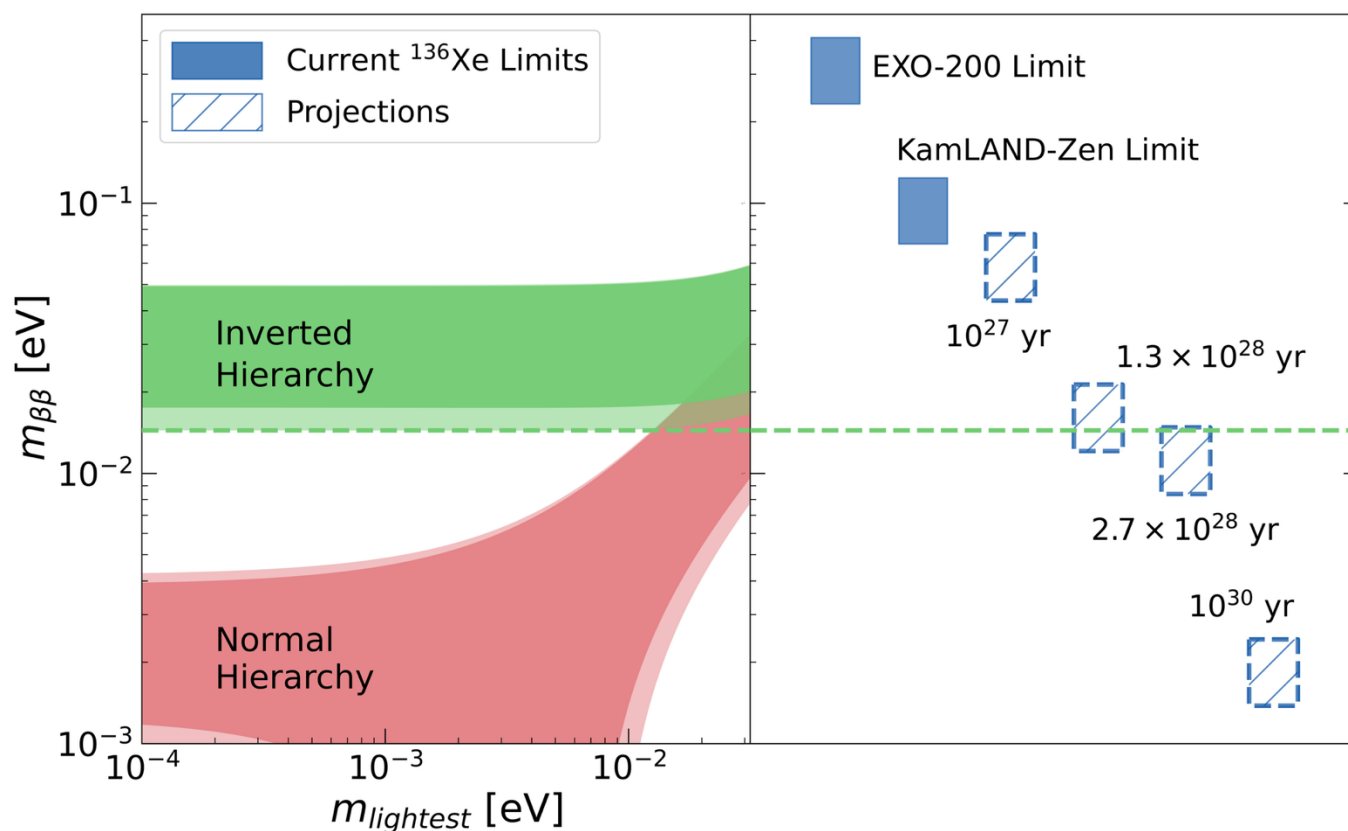


Shickele et al., in prep

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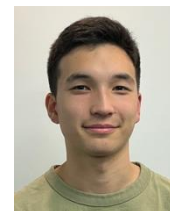
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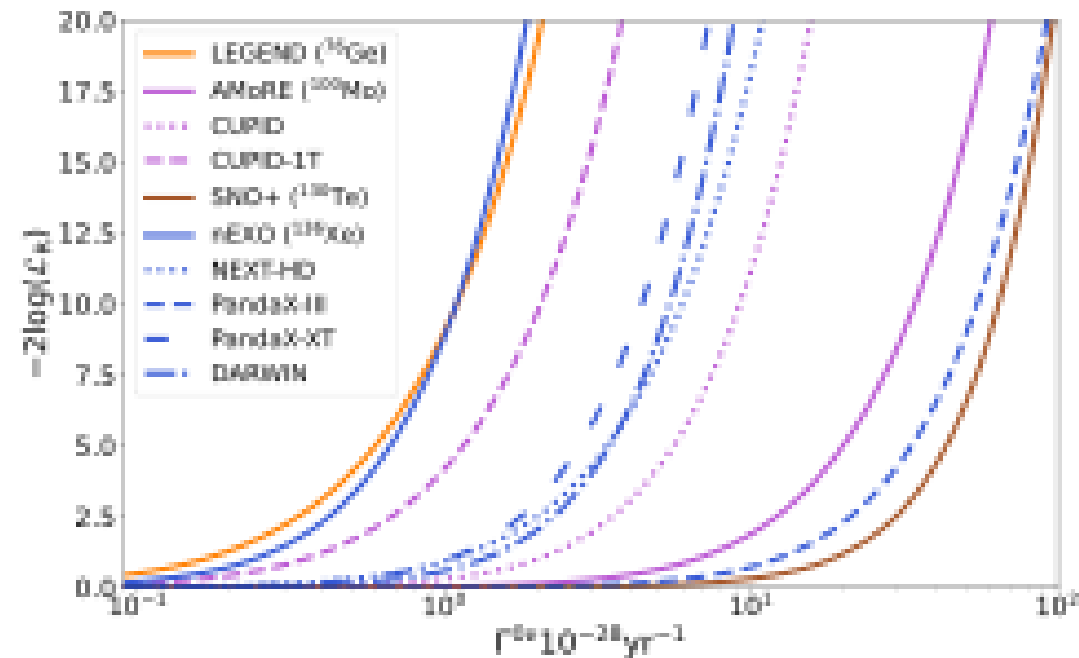
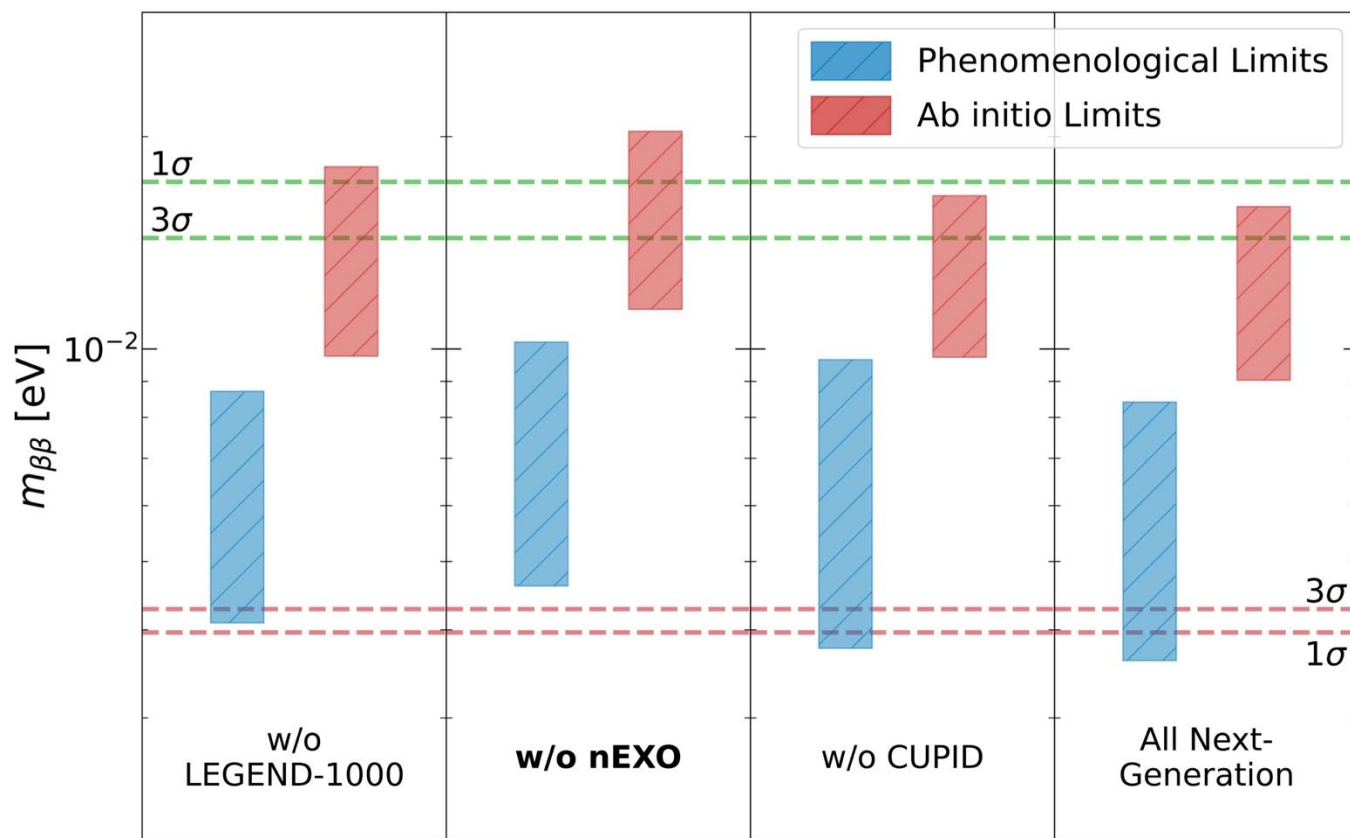


Shickele et al., in prep

Obtain likelihood functions from experiments – convert decay rates to $m_{\beta\beta}$

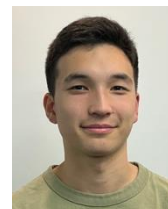
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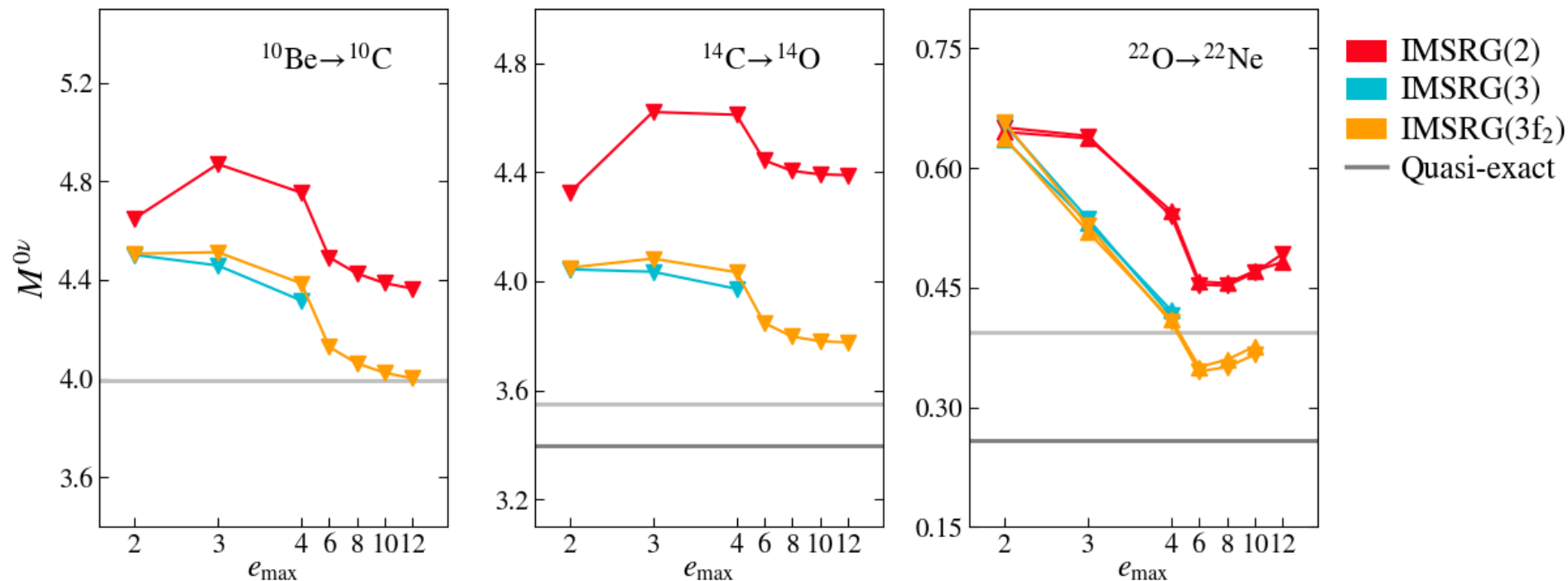


Shickele et al., in prep

Many-body largest source of uncertainty, need some estimation of higher-body operators

Compare with other (quasi-exact) ab initio method

Evaluate next order in IMSRG($2 \rightarrow 3f_2 \rightarrow 3$):



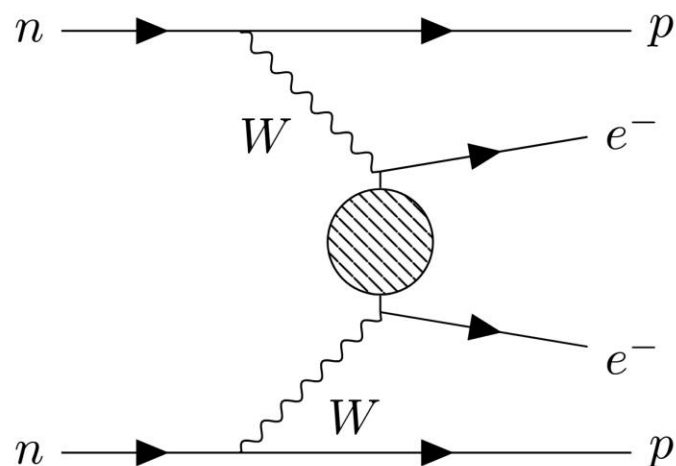
IMSRG(3f₂) very well approximates IMSRG(3)

In all cases: lowers NMEs, reproduces quasi-exact results



Explore other possible exchange mechanisms

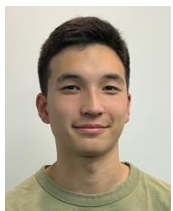
Heavy sterile neutrinos: complicated operator structures



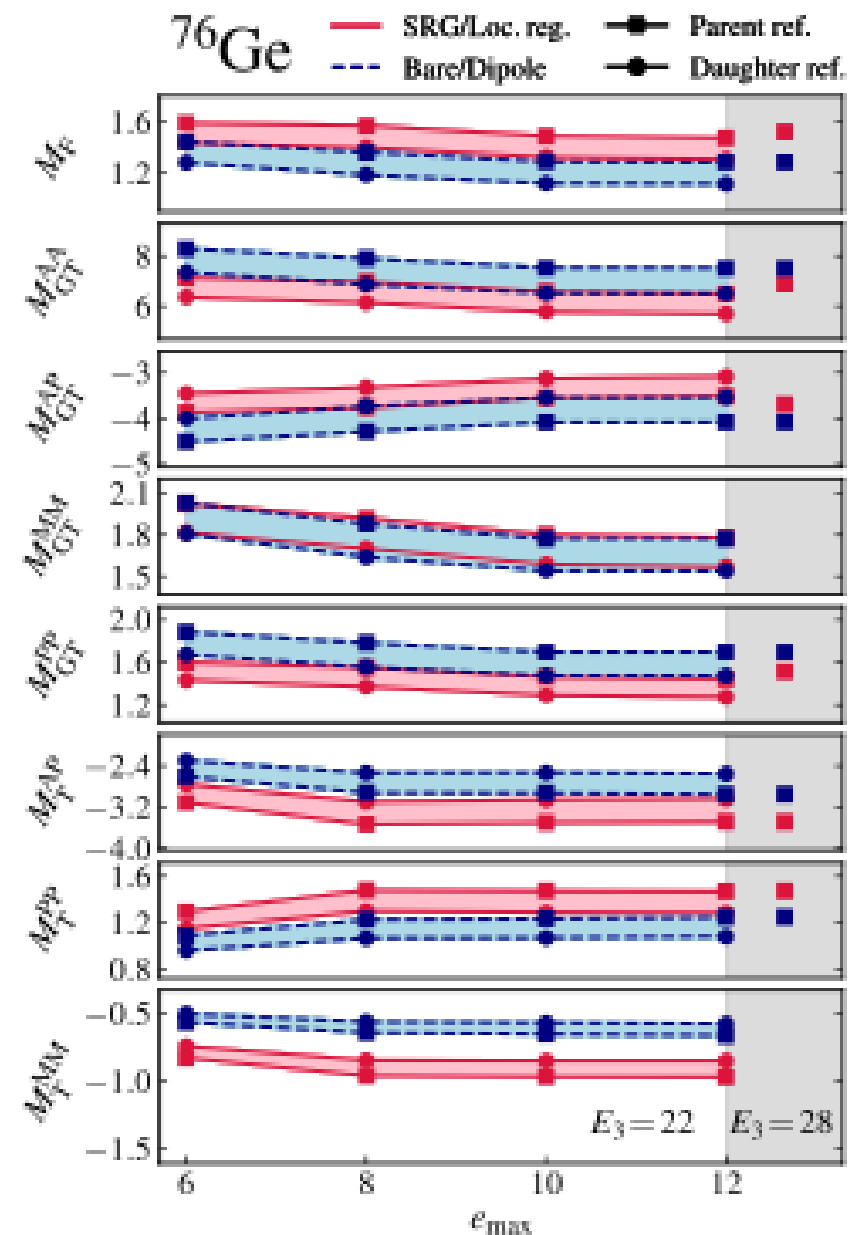
*V. Cirigliano et al.
J. High Energy Phys. (2018)*

$$\frac{1}{T_{1/2}^{0\nu}} = g_A^4 G^{0\nu} \left(|M_L^{0\nu}| \frac{m_{\beta\beta}}{m_e} + |M_H^{0\nu}| \frac{m_p}{M_{\beta\beta}} \right)^2$$

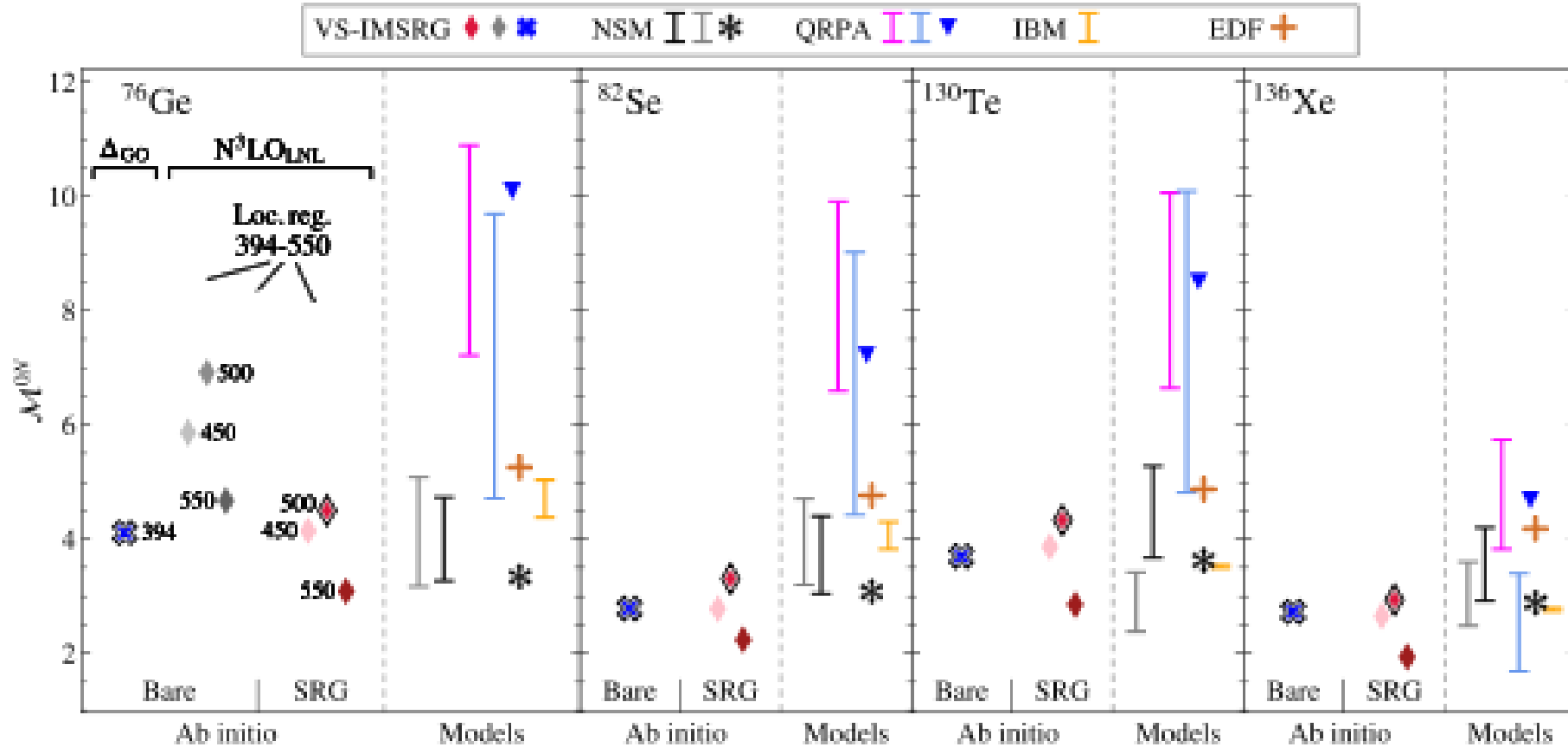
Converged heavy NMEs for all major candidates



Todd, Shickele et al., PRL (under review)



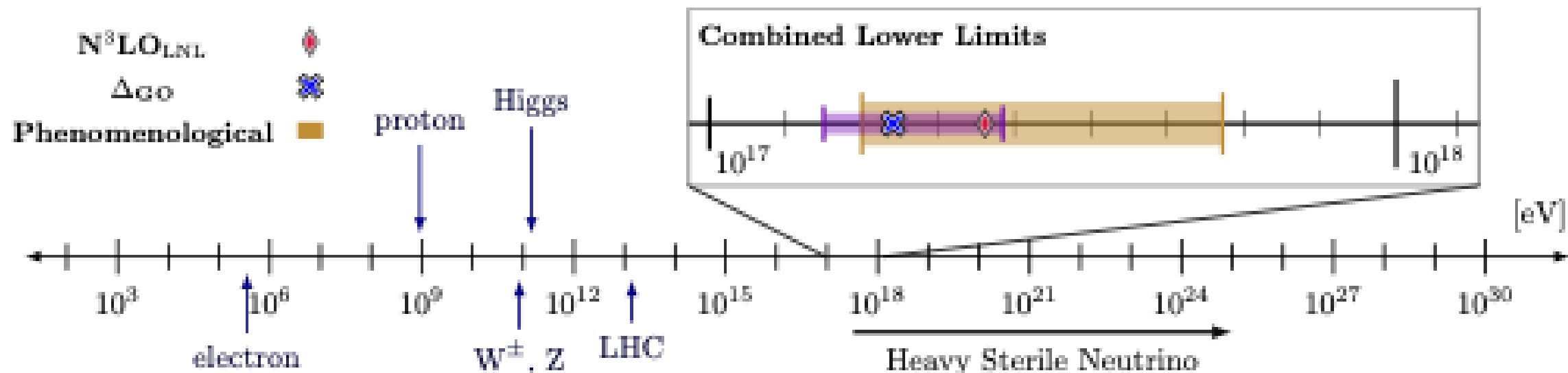
Generally smaller values (and spread) than nuclear models



Todd, Shickele et al., PRL (under review)

Use ab initio $0\nu\beta\beta$ decay to obtain lower limits for sterile neutrino mass

Generally smaller values (and spread) than nuclear models



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Use ab initio $0\nu\beta\beta$ to obtain lower limits for sterile neutrino mass

Results probe particle masses beyond the LHC: new collaborations w/ particle theory @TRIUMF

(Former)
Postdoctoral
Fellows

Future of Ab Initio Theory



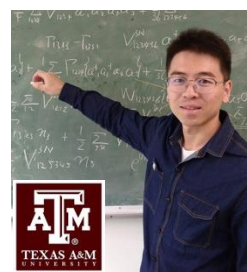
Ragnar Stroberg



Takayuki Miyagi



Lotta Jokiniemi

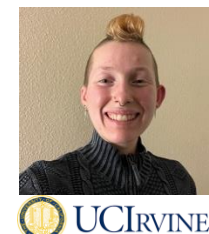


Baishan Hu

$0\nu\beta\beta$ Decay



Antoine Belley



Isabella Ginnette

Structure + Astrophysics



Kanting Motimele



Maude Larivière

Symmetry Violation



Jose Munoz



Jack Pitcher

AI/Machine Learning

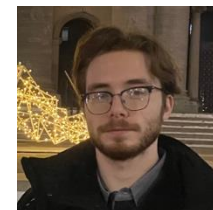


Emily Love

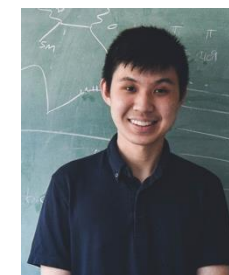


Matt Martin

CKM Unitarity



Alex Todd



Michael Liudeng

Atomic Systems



Gaurav Tenkila



Hrishi Patel



Vijay Chand

Just the beginning:
join our revolution!
jholt@triumf.ca

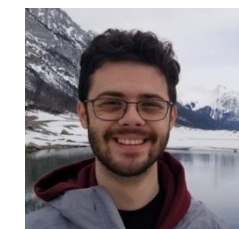
Dark Matter/ ν Scattering



Sam Leutheusser



Jose Padua



Mathieu Bruneault

Construct predictive posterior distribution for NMEs – first case ^{76}Ge

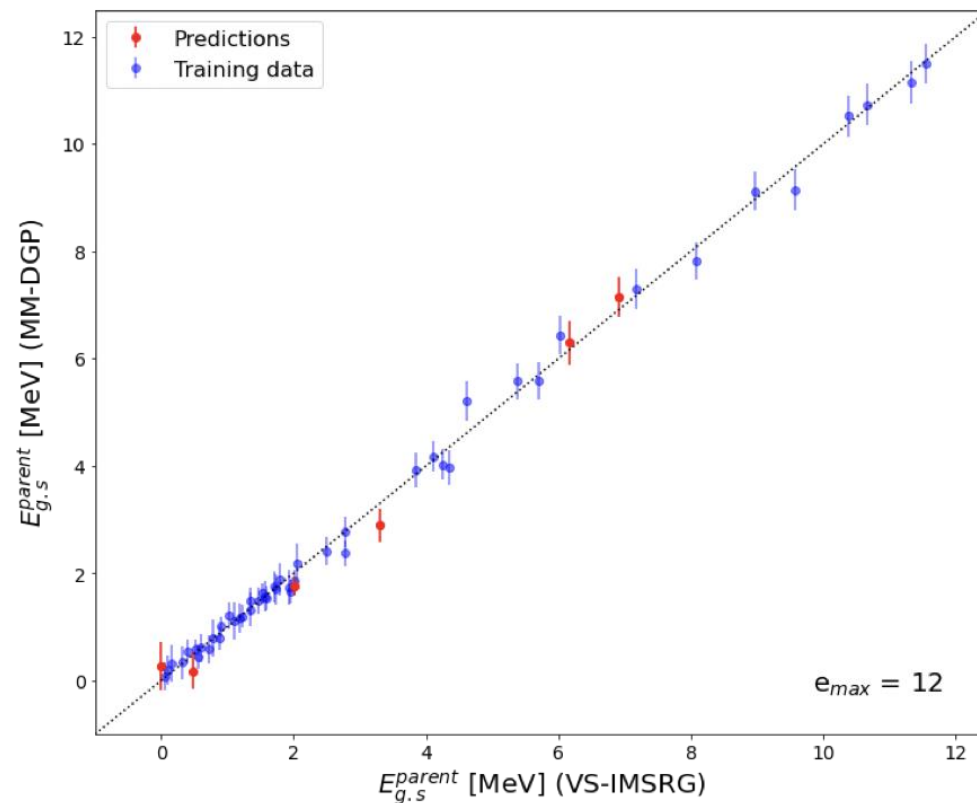
Consider all physics contributing to systematic error

$$y = y_{\text{MM-DDGP}} + \epsilon_{\text{emulator}} + \epsilon_{\text{EFT}} + \epsilon_{\text{many-body}} + \epsilon_{\text{operator}}$$

Use 8100+ samplings of NN+3N forces - weighted by np $^1\text{S}_0$ phase shift

Solve many-body problem with MM-DGP

Emulator: from process itself



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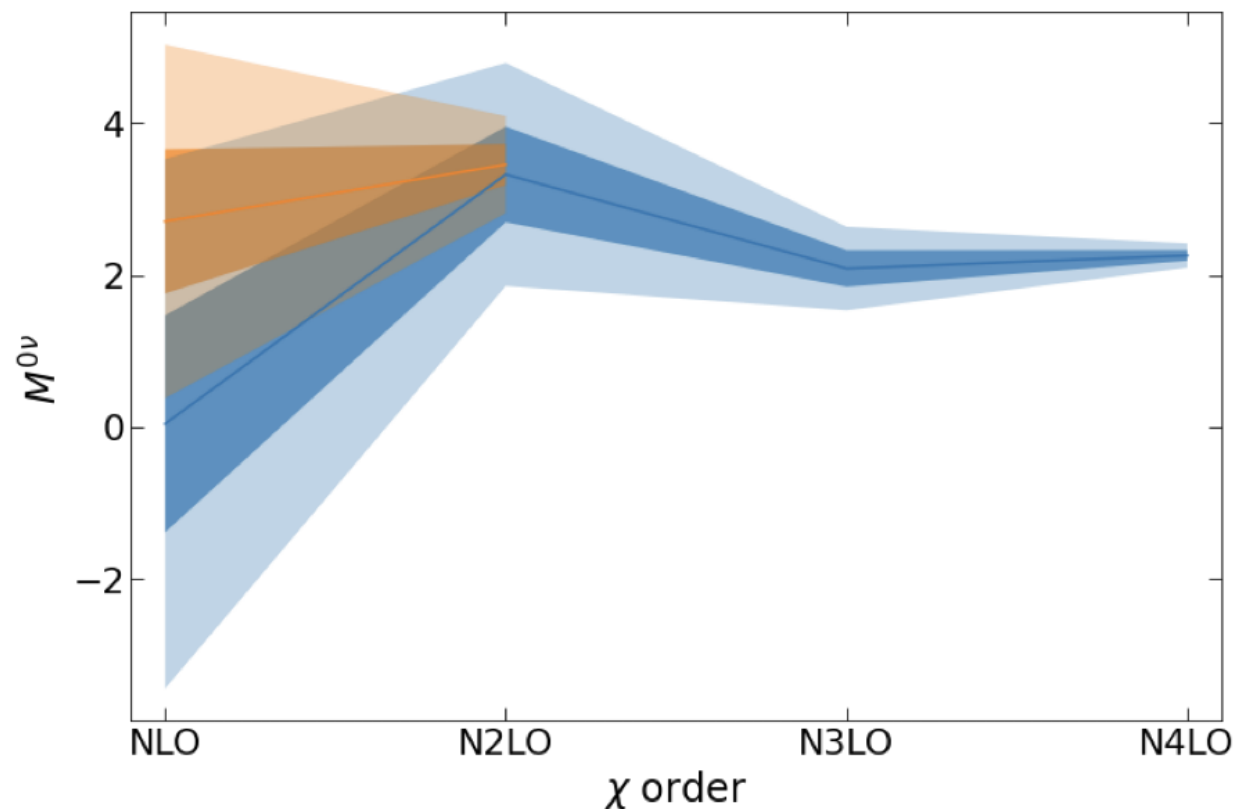
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EFT: Truncation in nuclear forces

Belley et al, arXiv:2308.15634



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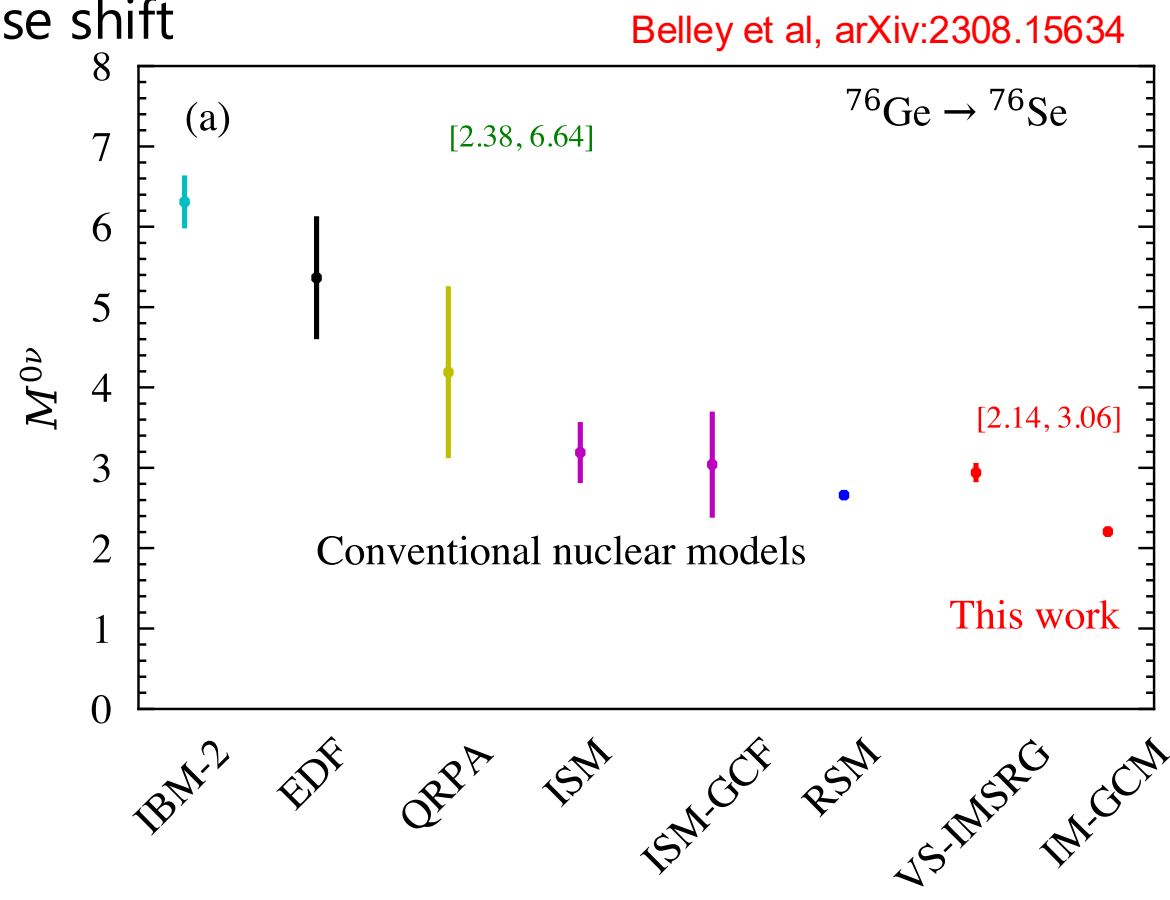
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Belley et al, arXiv:2308.15634

Use 8100+ NN+3N forces - weighted by np 1S_0 phase shift

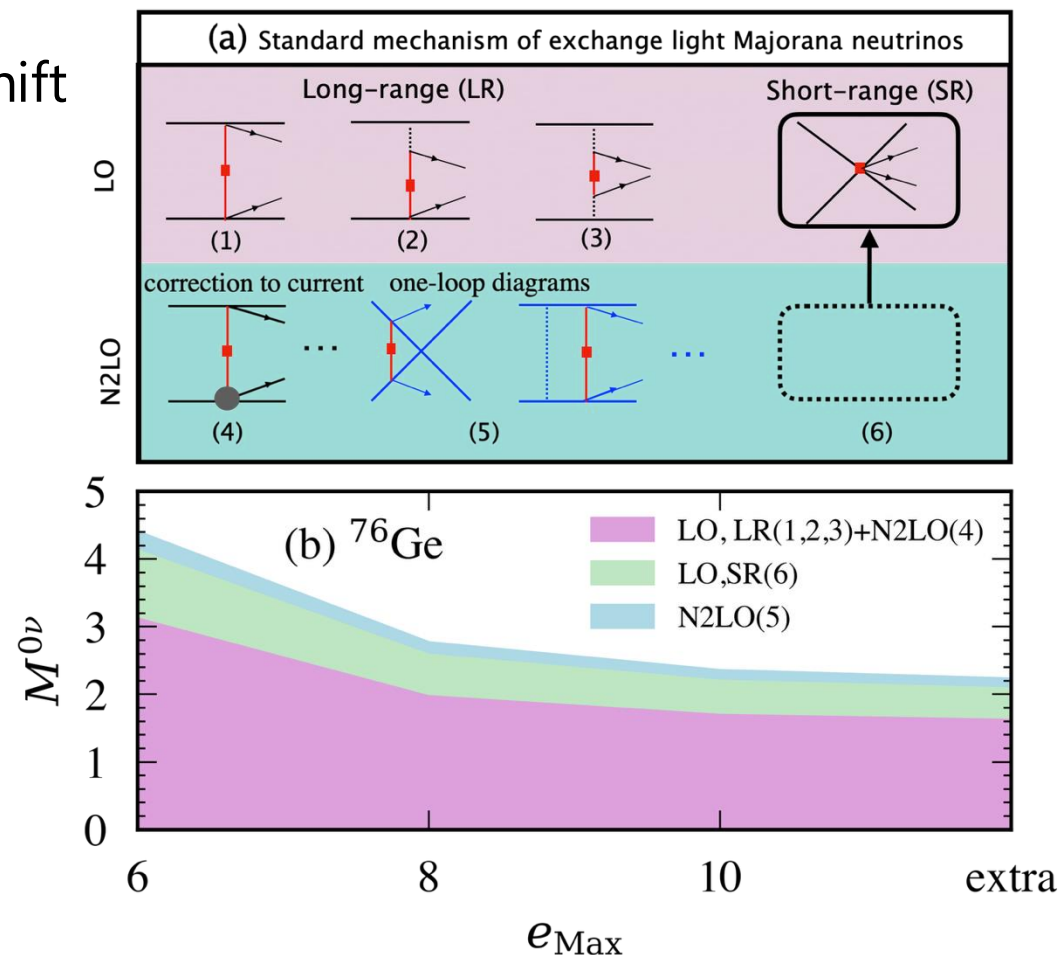
Solve many-body problem with MM-DGP

Emulator: from process itself

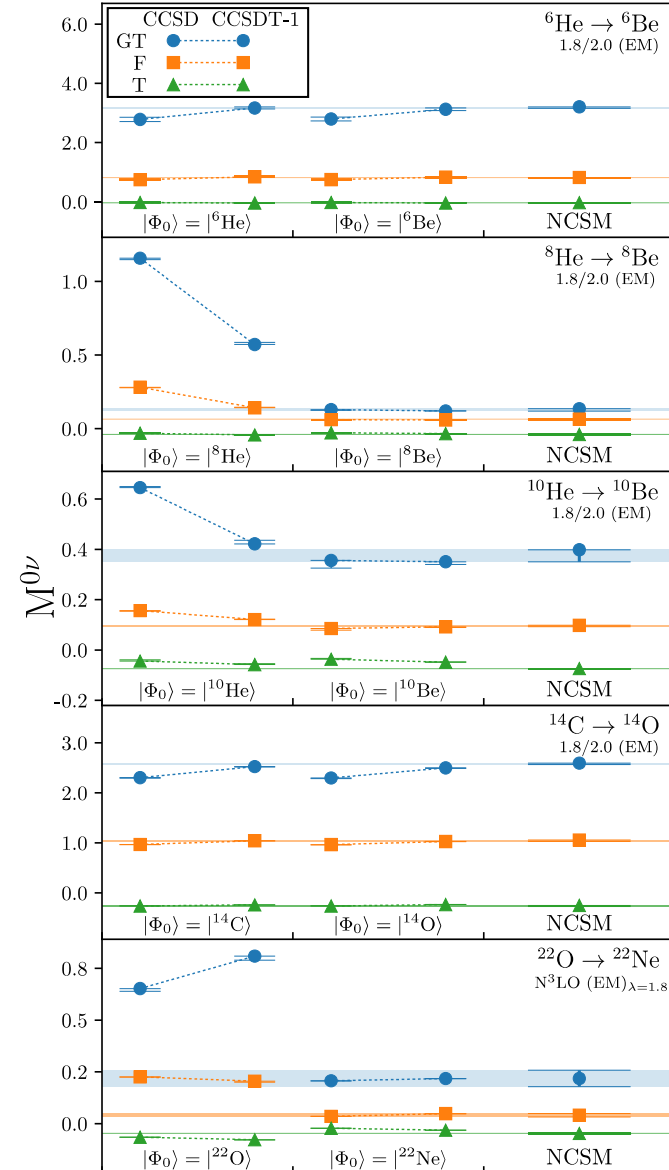
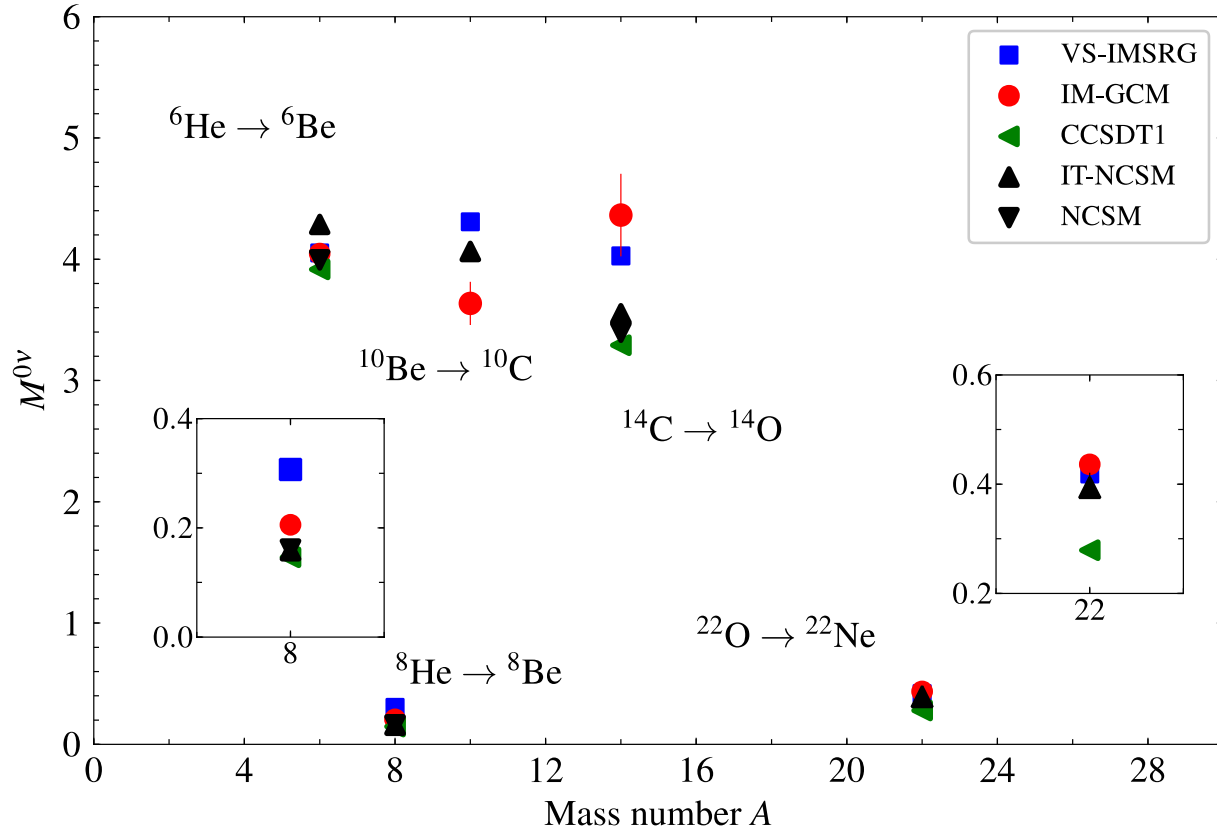
EFT: Truncation in nuclear forces

Many-body: Truncation in many-body physics

Operator: higher order contributions to operator



Benchmark with **quasi-exact ab initio methods** in light systems: $A=6-22$



Reasonable to good agreement in all cases

Pursue true double-beta decay candidates!

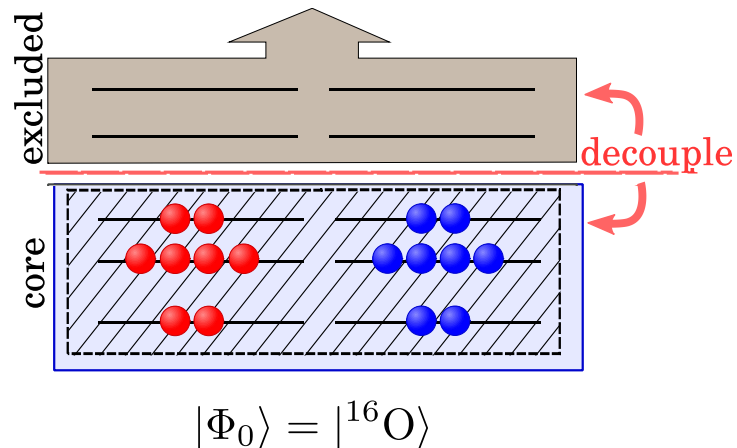
Explicitly construct unitary transformation from sequence of rotations

$$U = e^{\Omega} = e^{\eta_n} \dots e^{\eta_1} \quad \eta = \frac{1}{2} \arctan \left(\frac{2H_{\text{od}}}{\Delta} \right) - \text{h.c.}$$

$$\tilde{H} = e^{\Omega} H e^{-\Omega} = H + [\Omega, H] + \frac{1}{2} [\Omega, [\Omega, H]] + \dots$$

All operators truncated at two-body level IMSRG(2)
IMSRG(3) in progress

Step 1: Decouple core



Can we achieve accuracy
of large-space methods?

$$\langle \tilde{\Psi}_n | P \tilde{H} P | \tilde{\Psi}_n \rangle \approx \langle \Psi_i | H | \Psi_i \rangle$$

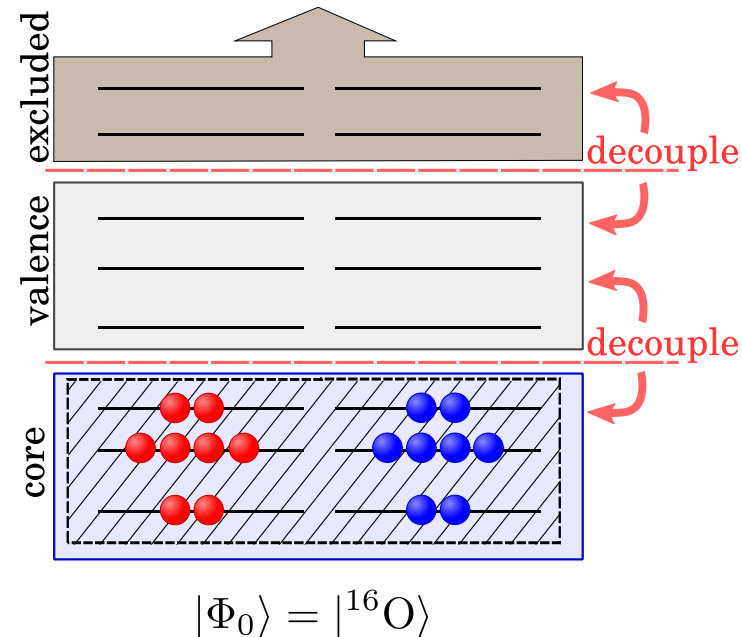
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Tsukiyama, Bogner, Schwenk, PRC 2012
 Morris, Parzuchowski, Bogner, PRC 2015



Step 1: Decouple core

Step 2: Decouple valence space

Can we achieve accuracy
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$$\langle \tilde{\Psi}_n | P \tilde{H} P | \tilde{\Psi}_n \rangle \approx \langle \Psi_i | H | \Psi_i \rangle$$

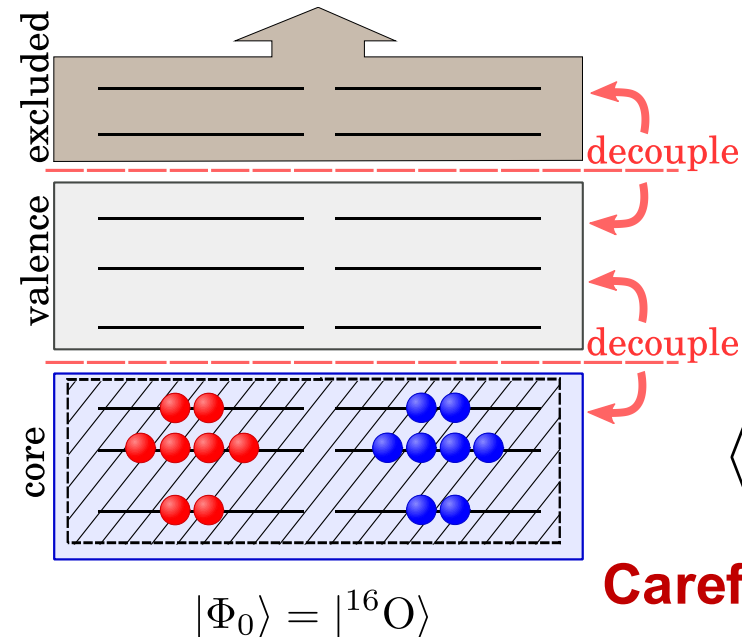
$\langle P H P \rangle$	$\langle P H Q \rangle \rightarrow 0$
$\langle Q H P \rangle \rightarrow 0$	$\langle Q H Q \rangle$

Explicitly construct unitary transformation from sequence of rotations

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$$\tilde{H} = e^{\Omega} H e^{-\Omega} = H + [\Omega, H] + \frac{1}{2} [\Omega, [\Omega, H]] + \dots$$

$$\tilde{\mathcal{O}} = e^{\Omega} \mathcal{O} e^{-\Omega} = \mathcal{O} + [\Omega, \mathcal{O}] + \frac{1}{2} [\Omega, [\Omega, \mathcal{O}]] + \dots$$



Step 1: Decouple core

Step 2: Decouple valence space

Step 3: Decouple additional operators

$$\langle \tilde{\Psi}_n | P \tilde{H} P | \tilde{\Psi}_n \rangle \approx \langle \Psi_i | H | \Psi_i \rangle$$

$$\langle \tilde{\Psi}_n | P \tilde{M}_{0\nu} P | \tilde{\Psi}_n \rangle \approx \langle \Psi_i | M_{0\nu} | \Psi_i \rangle$$

Careful benchmarking essential

$\langle P H P \rangle$	$\langle P H Q \rangle \rightarrow 0$
$\langle Q H P \rangle \rightarrow 0$	$\langle Q H Q \rangle$

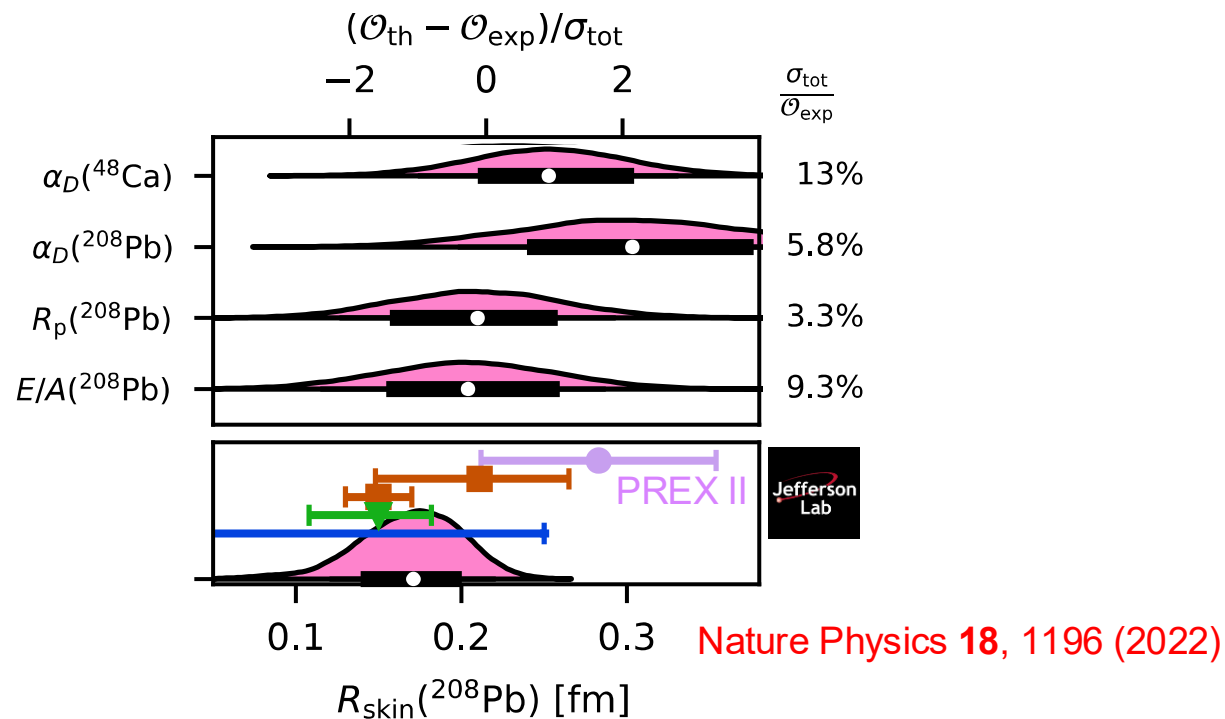
Combine ab initio + machine learning advances

I: History Matching: 34 “non-implausible” interactions

II: Calibration/importance sampling: ^{48}Ca properties

III: Validation: ^{208}Pb properties from ab initio

IV: Prediction: PPD for neutron skin



$R_{\text{skin}}(^{208}\text{Pb}) = 0.14 - 0.20 \text{ fm}$ – consistent with PREXII

Ab Initio Neutron Skin of ^{208}Pb

News & views

Nuclear physics

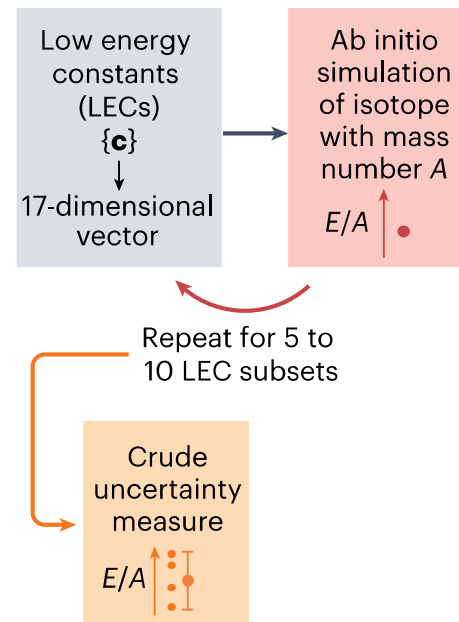
<https://doi.org/10.1038/s41567-022-01782-x>

A historic match for nuclei and neutron stars

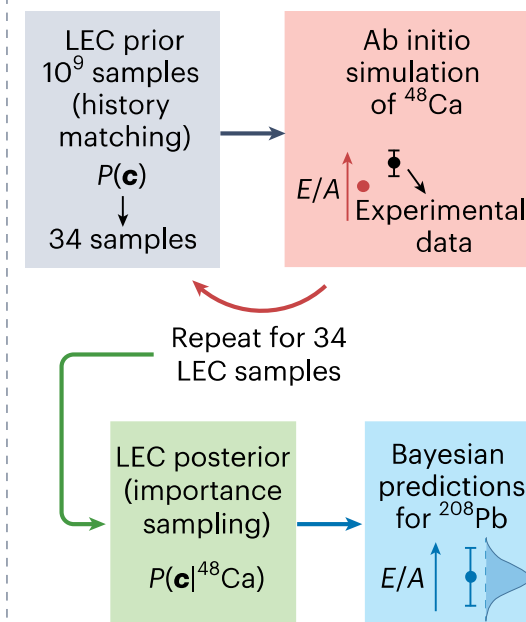
Arnau Rios

 Check for updates

Uncertainty estimation in previous work



Uncertainty estimation in this work



Combine TRIUMF/ORNL/Chalmers advances!

I: History Matching confronted with $A=2,3,4$ data + ^{16}O

10^9 calculations spanning EFT parameter space at N^2LO

34 non-implausible interactions

II: Calibration use ^{48}Ca E/A , $E(2^+)$, R_p , dipole polarizability

Importance resampling – statistically weight interactions

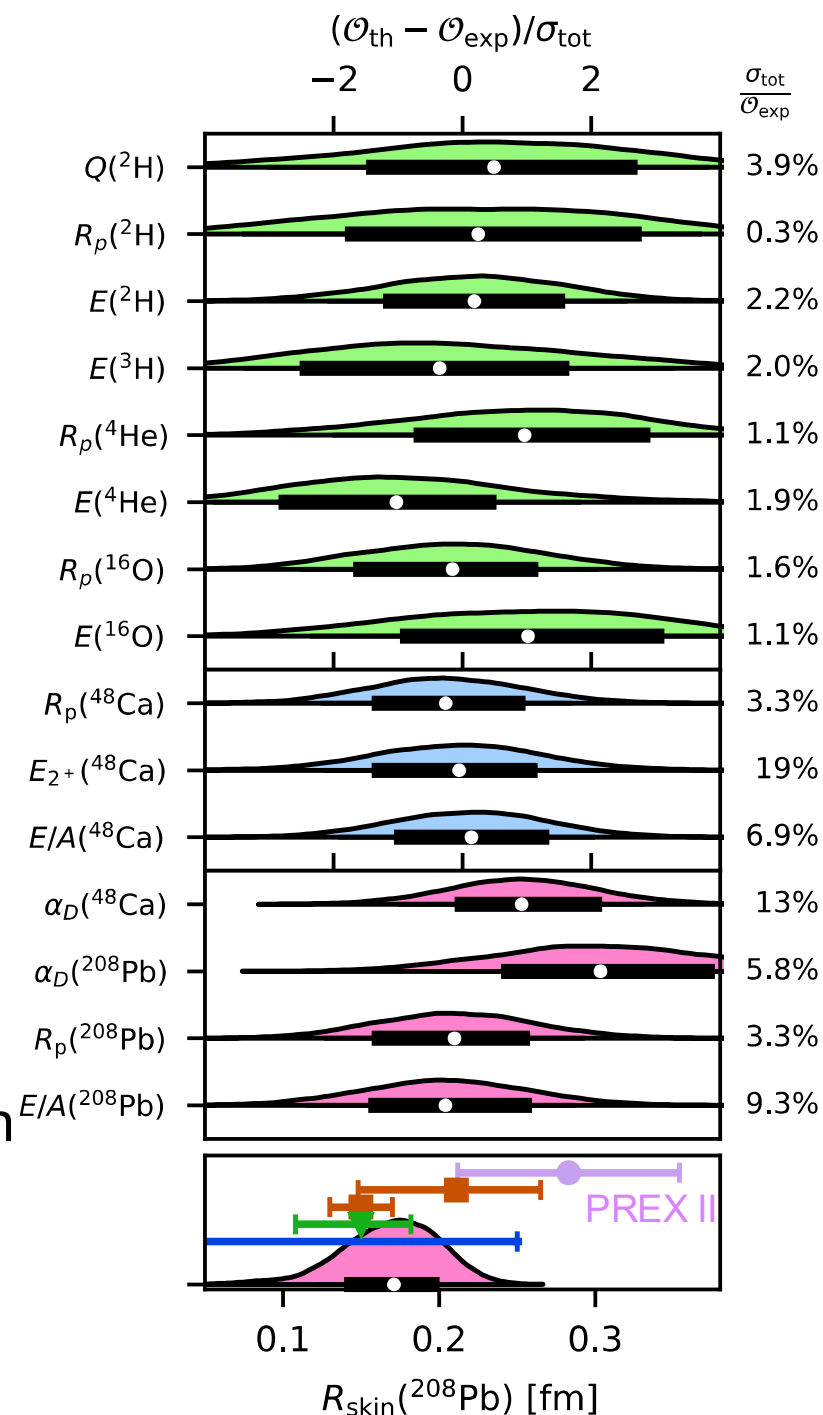
III: Validation ^{208}Pb E/A , R_p + $^{48}\text{Ca}/^{208}\text{Pb}$ DP from ab initio

Clear quality description of data

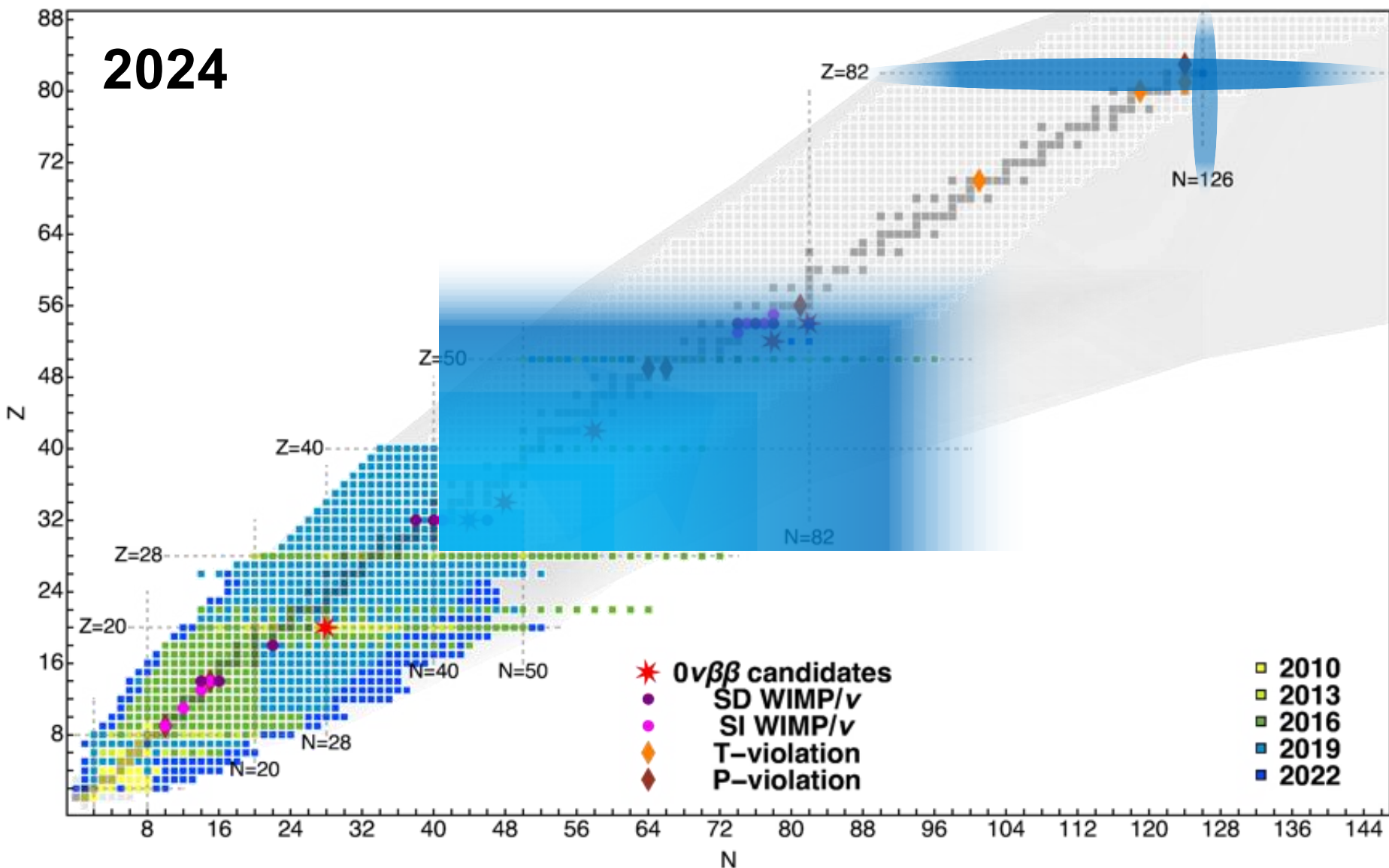
IV: Prediction - posterior predictive distribution for neutron skin

$R_{\text{skin}}(^{208}\text{Pb}) = 0.14\text{--}0.20\text{fm}$ (68% credible level)

Consistent(ish) with extracted **PREXII result**



(Not so) Bold predictions: Five new direction in the next 5 years



Nucleosynthesis

Ab initio input for r-process

Dark Matter/ ν Scattering

Ab initio structure functions

Symmetry Violation

Schiff/anapole moments

CKM Unitarity (V_{ud})

Superaligned Fermi

Atomic/molecular/QC(?)

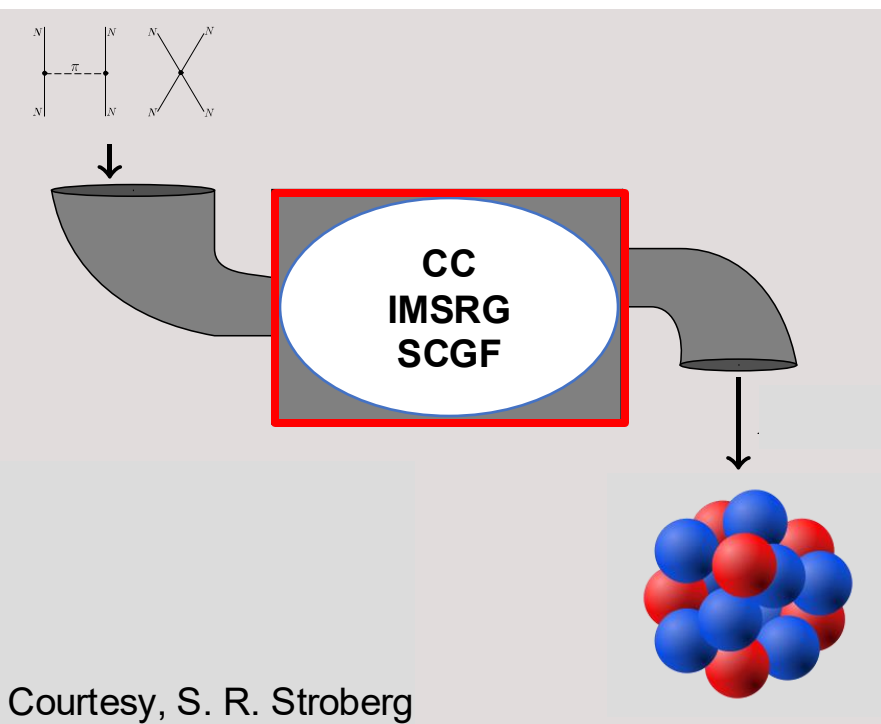
AI/Machine Learning

New emulators

Aim of modern nuclear theory: develop unified *first-principles* picture of structure and reactions

(Approximately) solve nonrelativistic Schrödinger equation

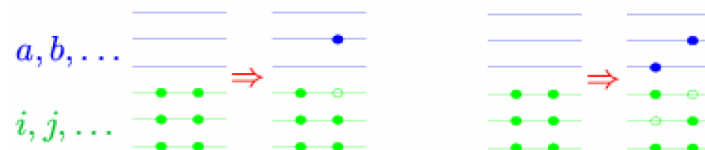
$$H\psi_n = E_n\psi_n$$



Courtesy, S. R. Stroberg

Ab Initio Cheat Sheet (polynomial scaling methods)

CC: Coupled cluster theory



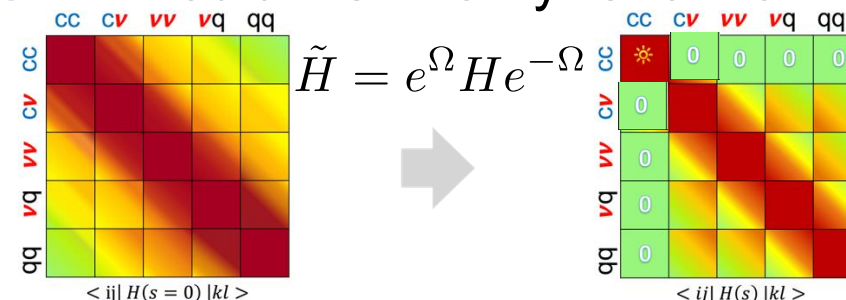
$$E = \langle \Phi | \bar{H} | \Phi \rangle$$

$$0 = \langle \Phi_i^a | \bar{H} | \Phi \rangle$$

$$0 = \langle \Phi_{ij}^{ab} | \bar{H} | \Phi \rangle$$

$$\bar{H} \equiv e^{-T} H e^T = (H e^T)_c$$

IMSRG: In-medium similarity renormalization group



SCGF: Self-consistent Green's function

$$\begin{array}{c} | \\ | \\ | \end{array} = \begin{array}{c} | \\ | \\ | \end{array} + \begin{array}{c} \bigcirc \\ \bigcirc \end{array} \Sigma = \begin{array}{c} \bigcirc \\ \bigcirc \end{array} + \begin{array}{c} \bigcirc \\ \bigcirc \end{array} + \dots$$

Global Ab Initio Calculations: Limits of Nuclear Existence



Ragnar Stroberg

Featured in Physics

Editors' Suggestion

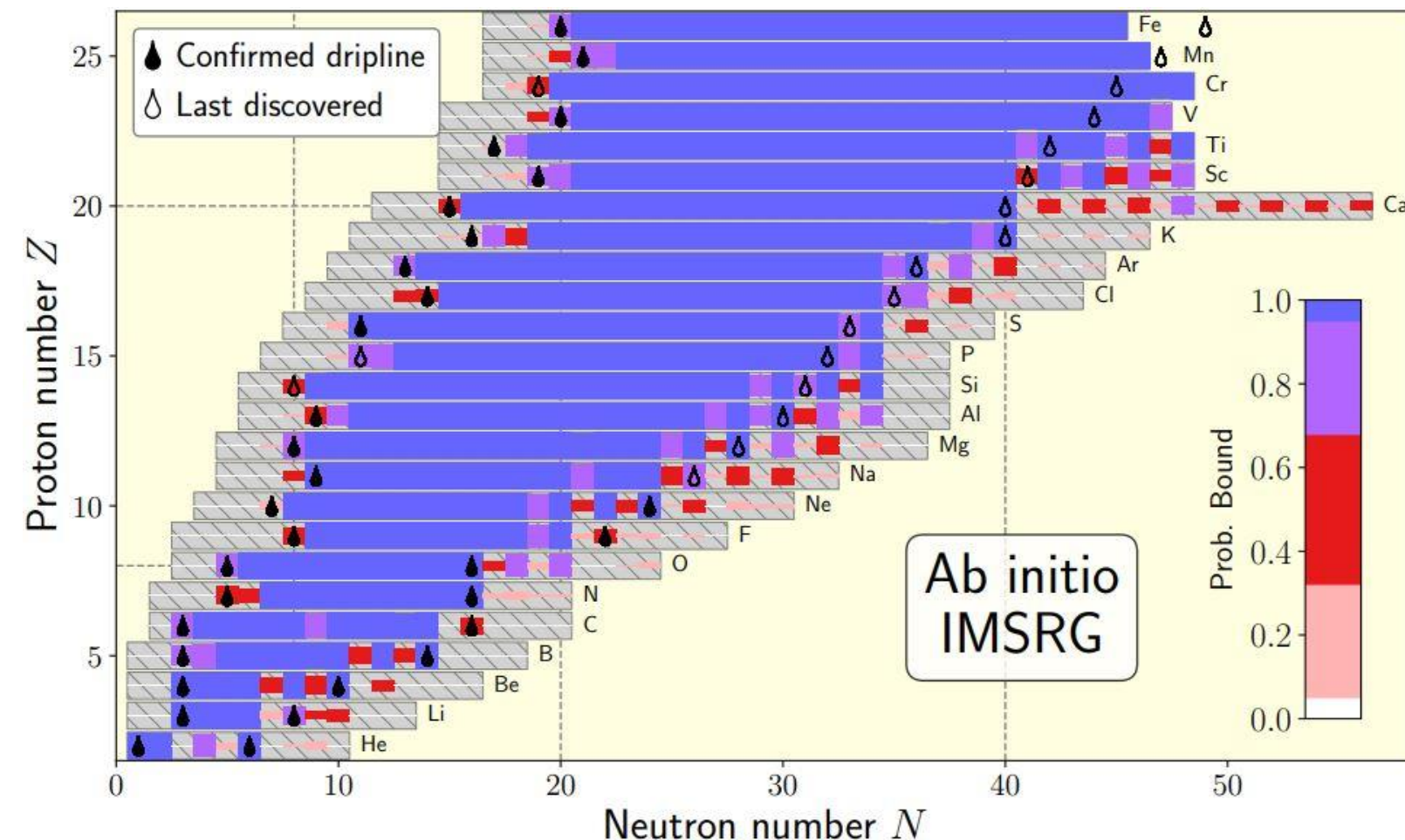
Ab Initio Limits of Atomic Nuclei



S. R. Stroberg, J. D. Holt, A. Schwenk, and J. Simonis
Phys. Rev. Lett. **126**, 022501 – Published 12 January 2021

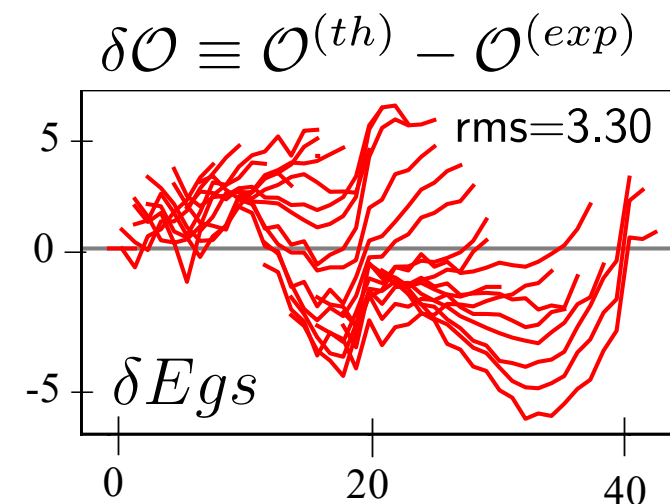
PhysICS See synopsis: [Predicting the Limits of Atomic Nuclei](#)

Ab initio calculations of ~700 nuclei from He to Fe



Known drip lines predicted within uncertainties

Ab initio guide for neutron-rich driplines



Input H fit to 2,3,4-body
Not biased towards existing data

Featured in Physics

Editors' Suggestion



Ab Initio Limits of Atomic Nuclei

S. R. Stroberg, J. D. Holt, A. Schwenk, and J. Simonis
Phys. Rev. Lett. **126**, 022501 – Published 12 January 2021

PhysICS See synopsis: [Predicting the Limits of Atomic Nuclei](#)

Ab Initio Neutron Skin of ^{208}Pb Linked to Neutron Star Properties

nature
physics

ARTICLES

<https://doi.org/10.1038/s41567-022-01715-8>

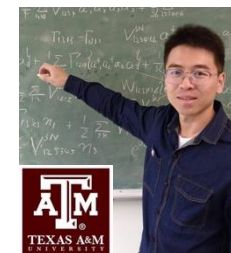
OPEN

Ab initio predictions link the neutron skin of ^{208}Pb to nuclear forces

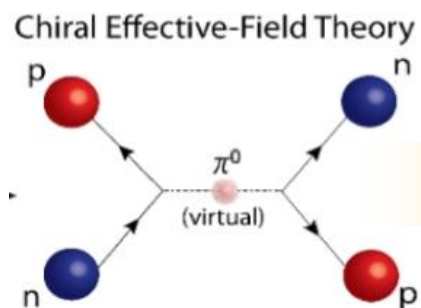
Baishan Hu^{1,11}, Weiguang Jiang^{2,11}, Takayuki Miyagi^{1,3,4,11}, Zhonghao Sun^{5,6,11}, Andreas Ekström², Christian Forssén², Gaute Hagen^{1,5,6}, Jason D. Holt^{1,7}, Thomas Papenbrock^{5,6}, S. Ragnar Stroberg^{8,9} and Ian Vernon¹⁰



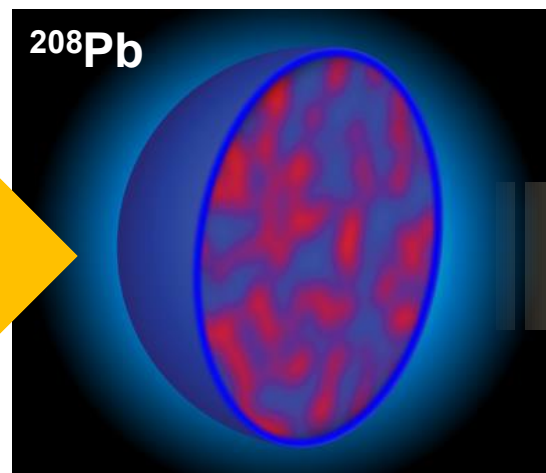
Takayuki Miyagi



Baishan Hu



Ab Initio



Nuclear
EOS

