

Progress and Future Plan of PandaX

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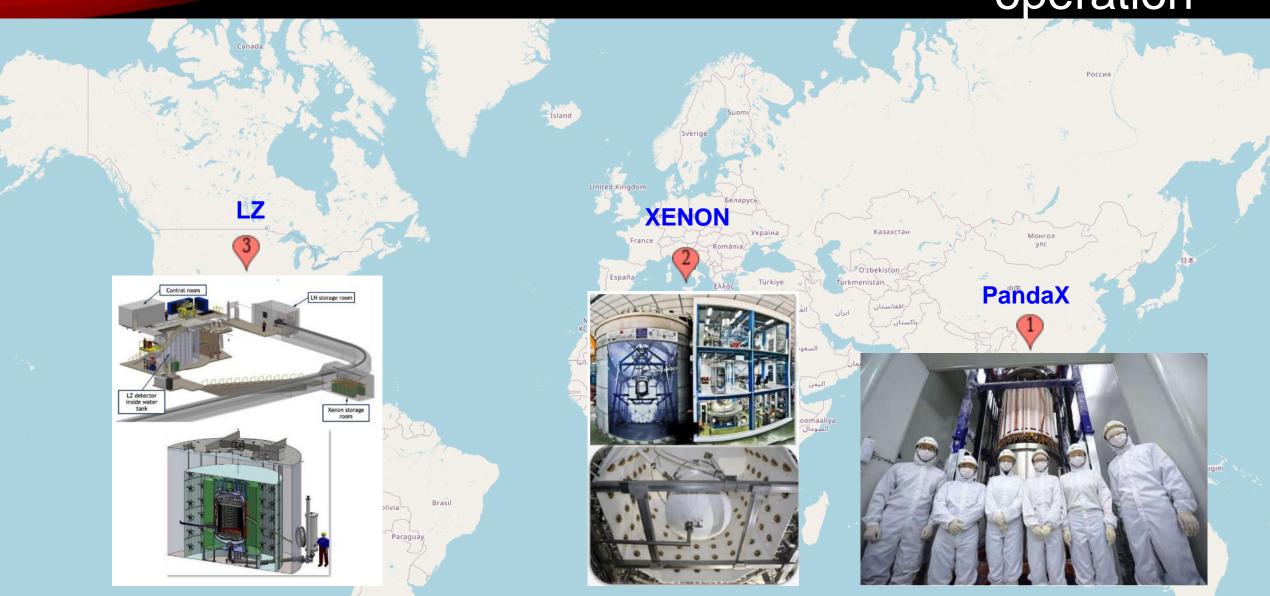




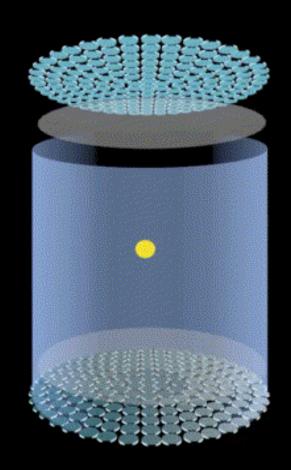
Particle and Astrophysical Xenon observatory



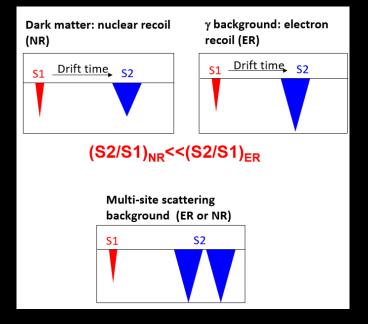
Natural xenon detectors in operation



Gamma electron recoil nuclear recoil



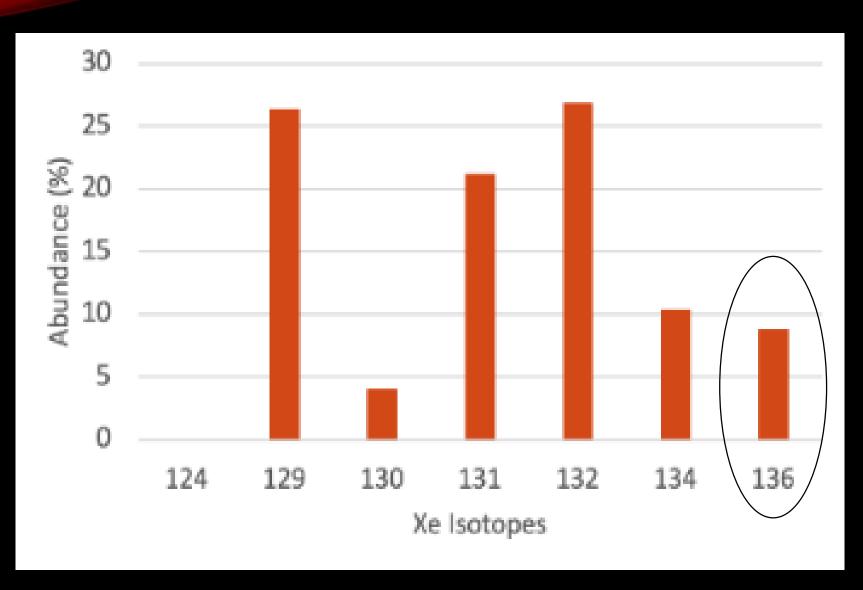
Dual-phase xenon time projection chamber (TPC)



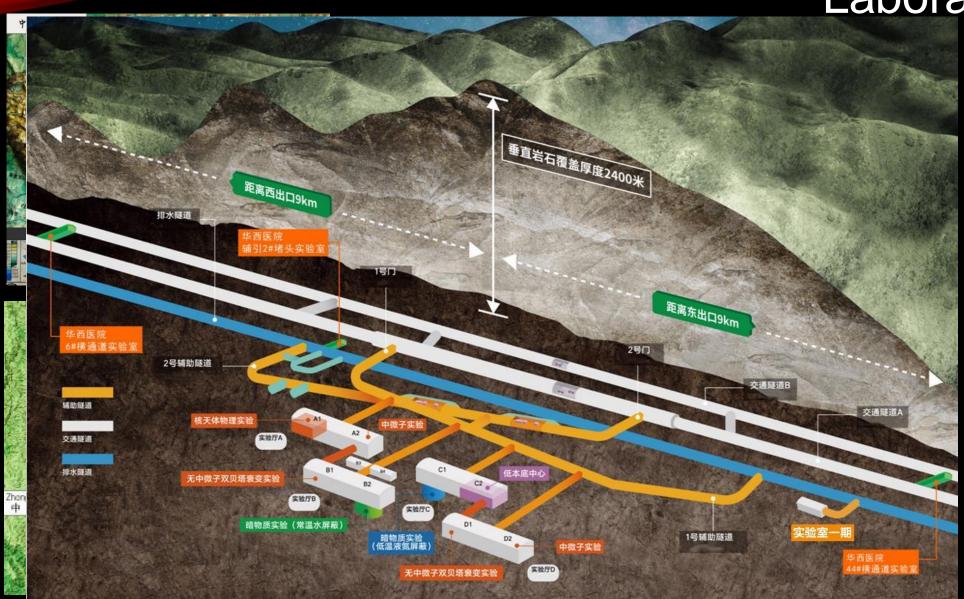
Detector capability:

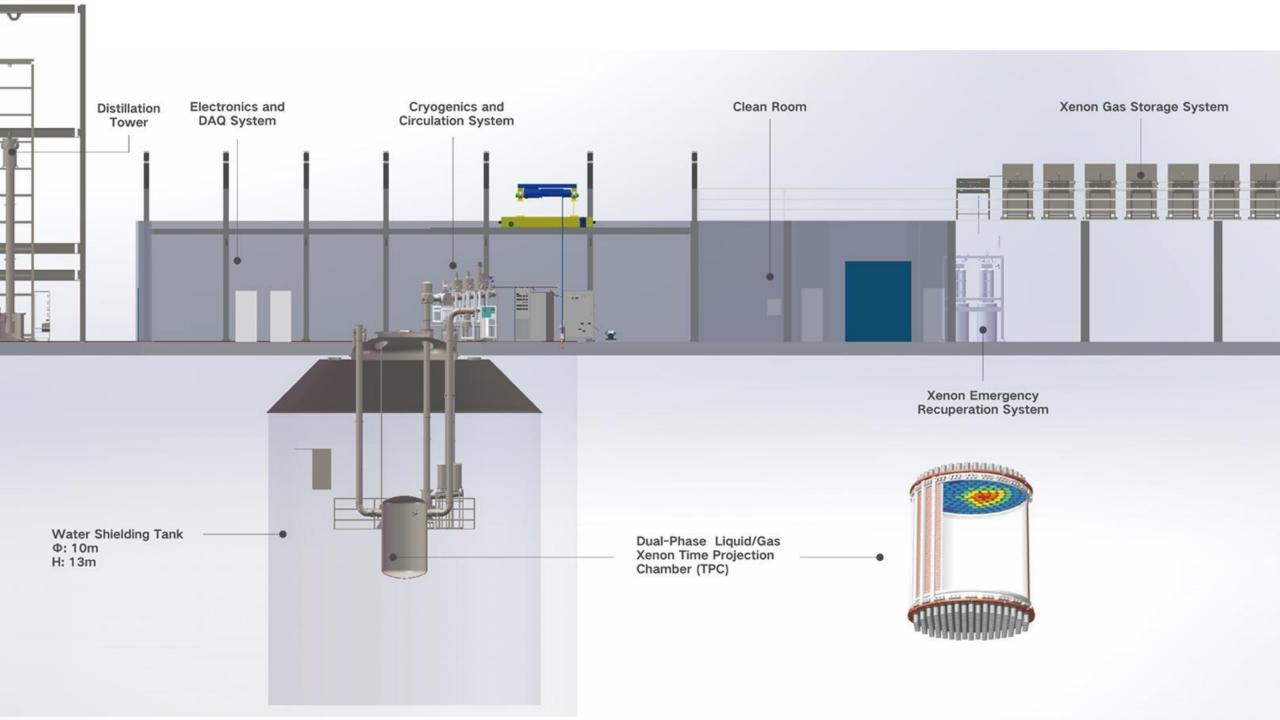
- 3D reconstruction and fiducialization
- Good ER/NR rejection
- Calorimeter capable of seeing a couple of photons/electrons

Natural abundance



China Jinping underground Laboratory





construction





instllation







2020.5 liquid xenon filling





2020/6-2020/11 integration tests







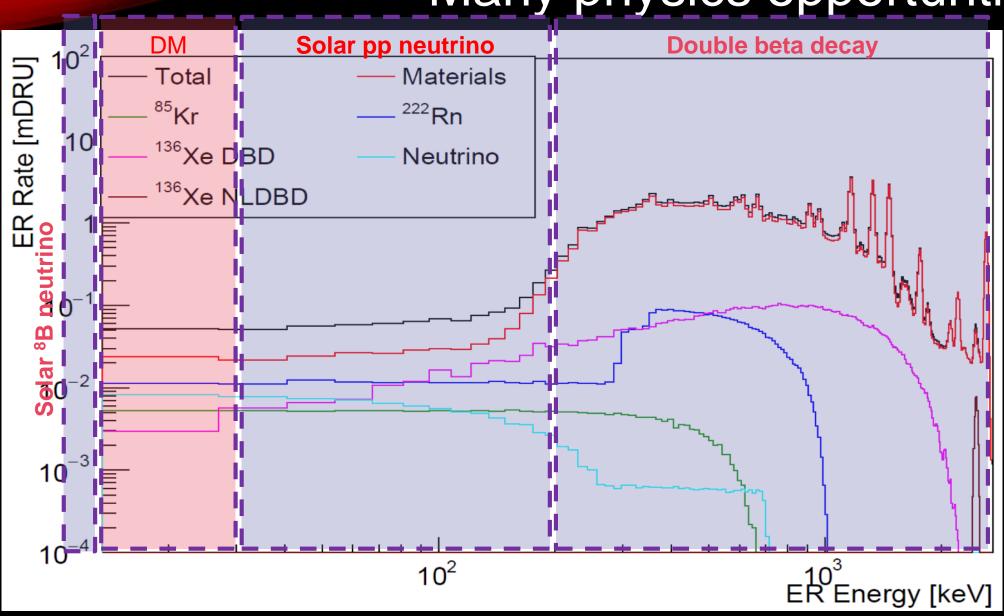
Data taking interleaved with surgeries and lab construction

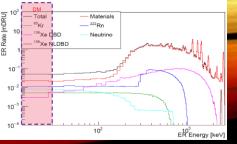
Time	Activity
2020/11- 2021/04	commissioning (Run0) 95 days
2021/07- 2021/10	tritium removal xenon distillation, gas flushing, etc.
2021/11- 2022/05	physics run (Run1) 164 days
2022/09- 2023/12	CJPL B2 hall renovation xenon recuperation, detector upgrade
2024/01- 2024/07	resuming operation
Current status	physics data taking (Run2)



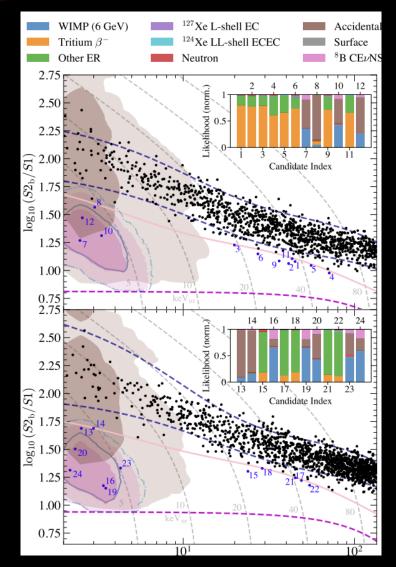


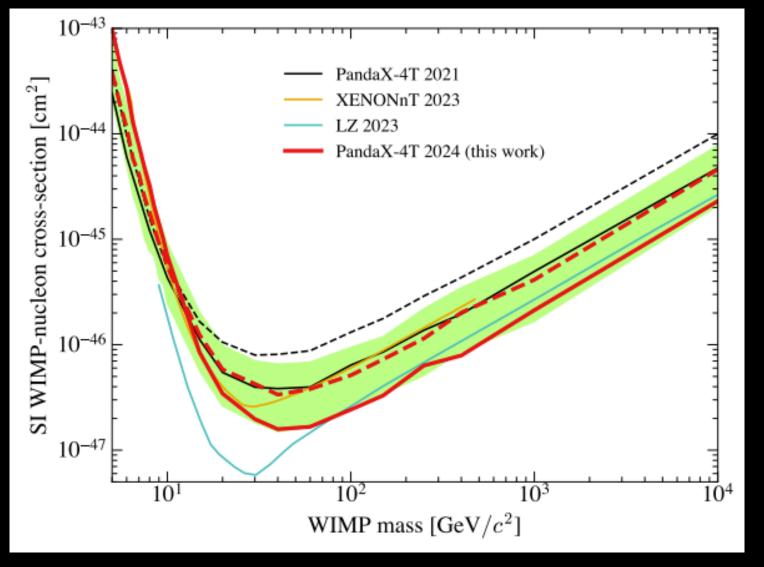
Many physics opportunties

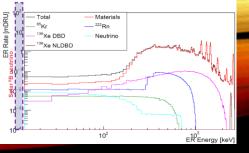




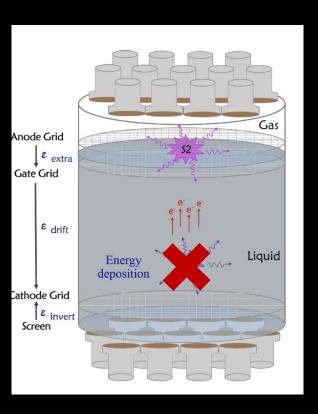
1.5 tonne*year, blinded analysis, Run0+1

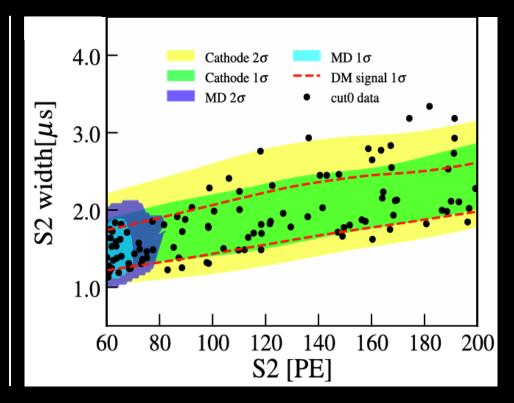


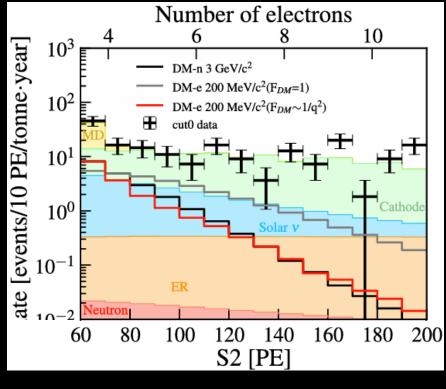


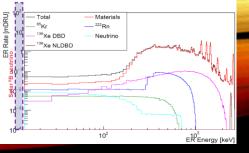


S2-only approach

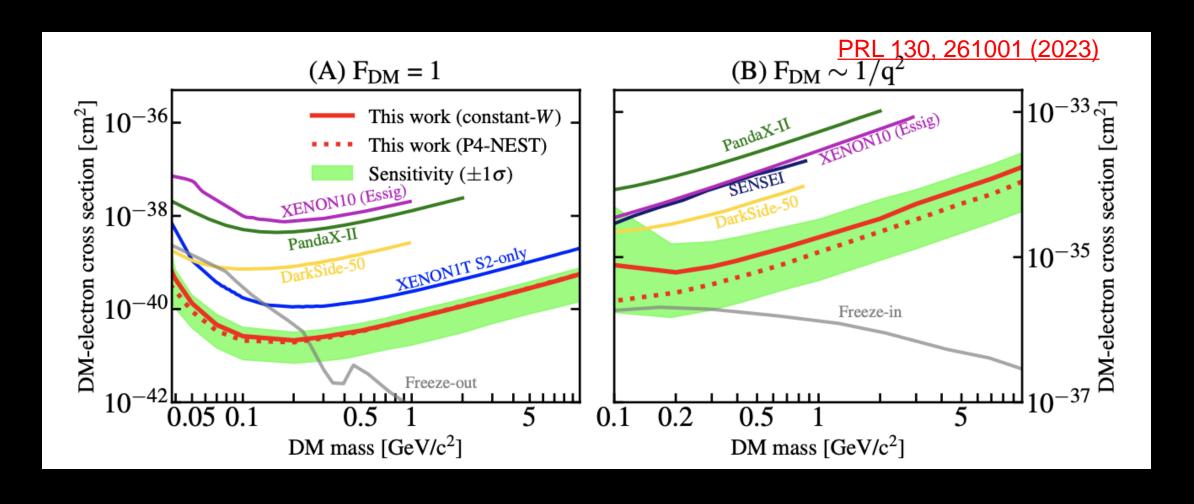


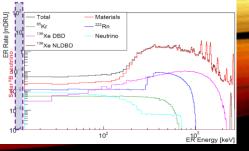






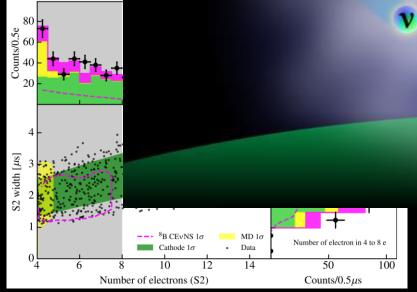
Tight limits on DM-e scattering with S2-only analysis

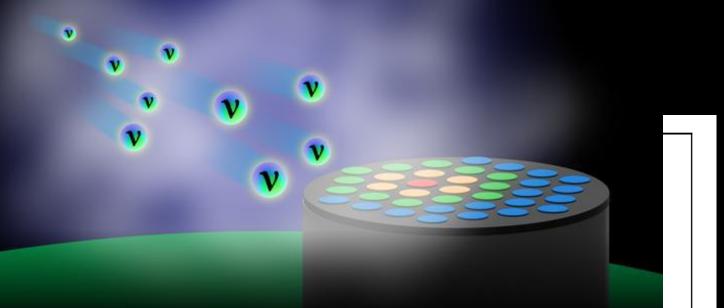




B8 neutrino CEvNS

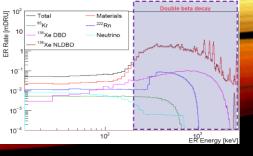
- Greatly benefit from S2-only data (~ton×year)
- Best-fit Po oventer 75+20/62 ank/) and 2 5+1 2/61 62 naired)
- Reject ba
- PRL 133,
- XENONn



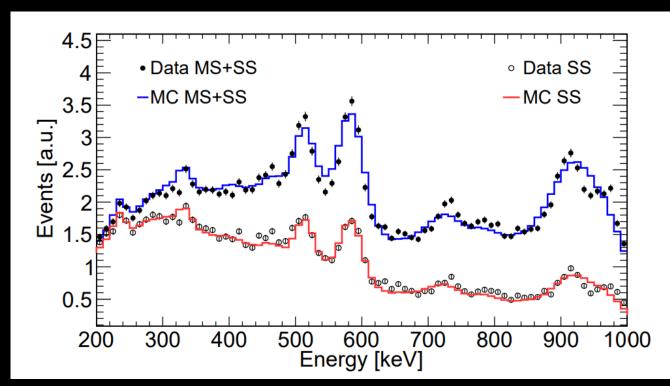


Solar 8 B ν flux [$\times 10^{6}$ cm $^{-2}$ s $^{-1}$]

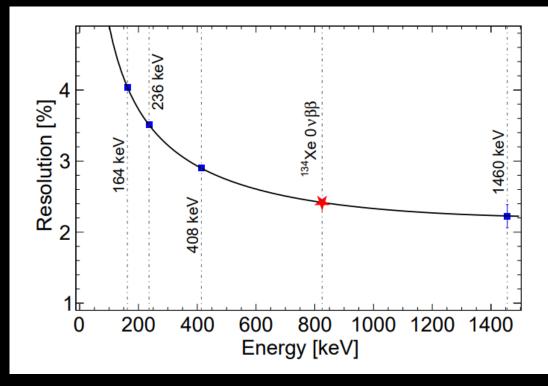
Physics Magazine highlights of the year 2024



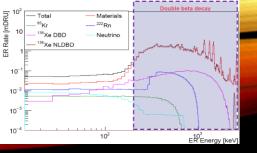
Calibrating PandaX at MeV



Th232 gamma calibration data

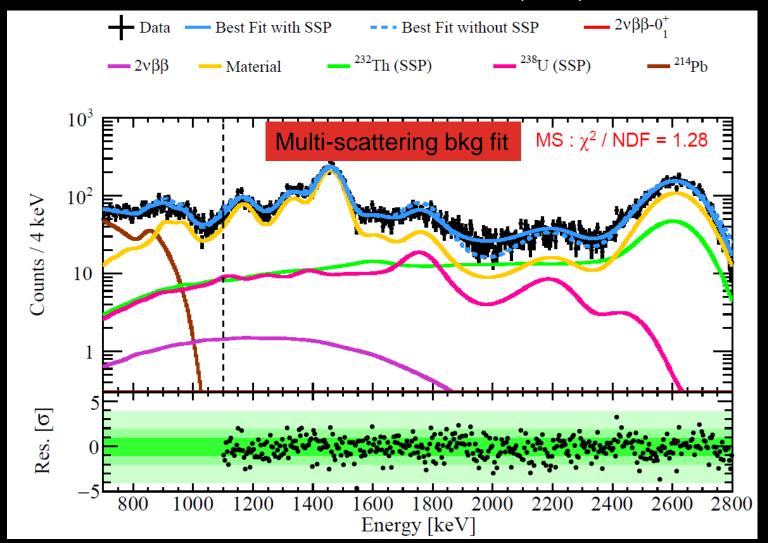


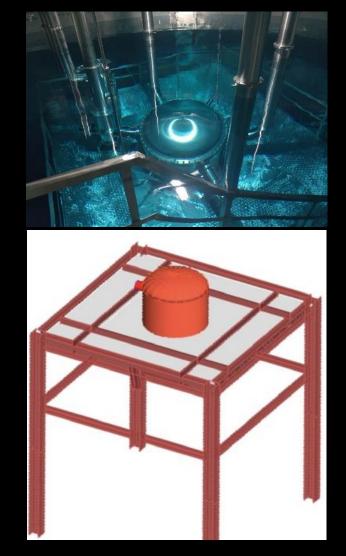
Internal gamma peaks

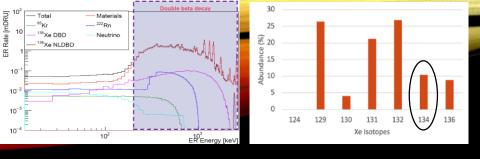


Understanding background at MeV

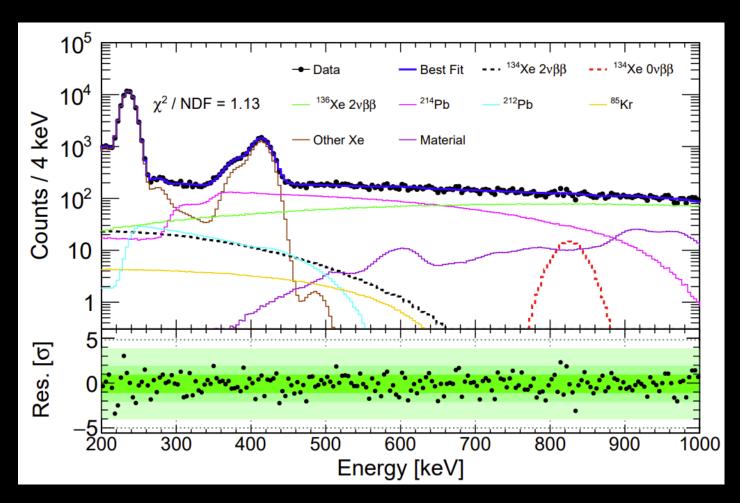
JHEP 05 (2025) 089, arXiv:2502.03017

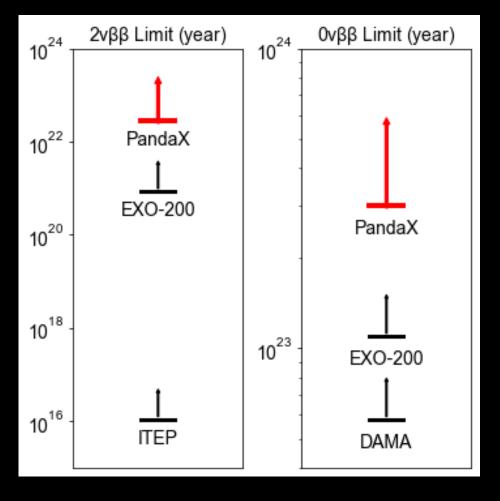




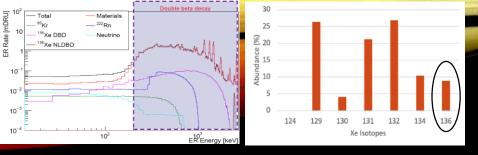


¹³⁴Xe: a unique blessing

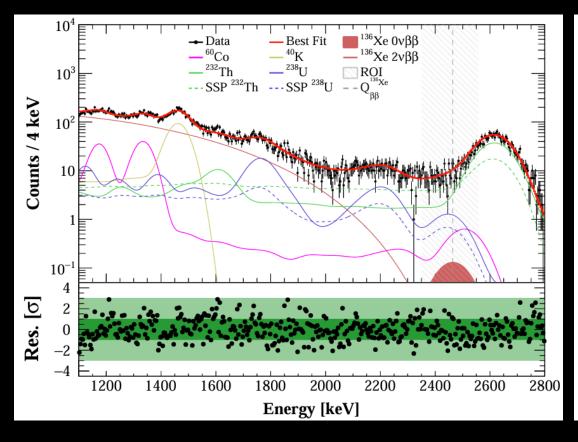


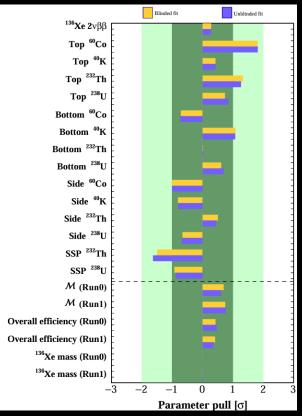


PRL 132, 152502 (2024)



Search for Xe136 0vDBD

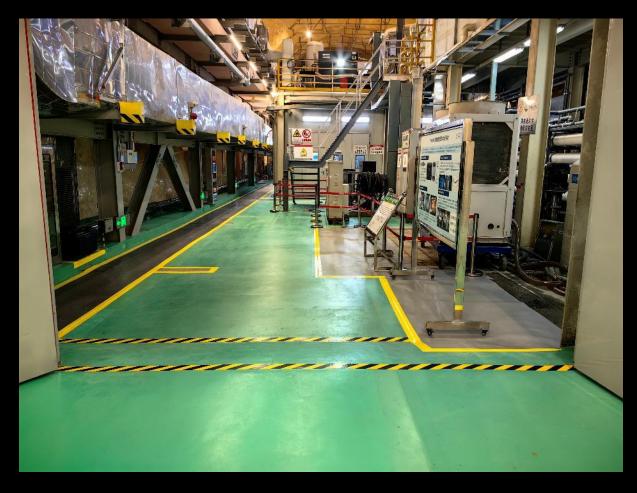




Science Bulletin Vol. 70,11 (2025) arXiv:2412.13979

Lifetime $> 2.1 \times 10^{24}$ year 90% CL $m_{\beta\beta}$ <0.6-1.4 eV

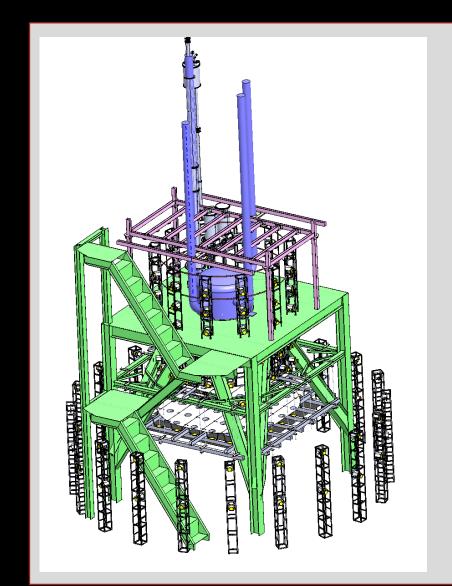
2024: Shiny again!





Newly implemented water veto detector

- installed 270 8inch PMTs in late 2023
- inner layer: 2 m radius, about 1 m to the outer surface of the LXe cryostat
- outer layer: 5 m radius
- Seeing muons, seeing neutrons!



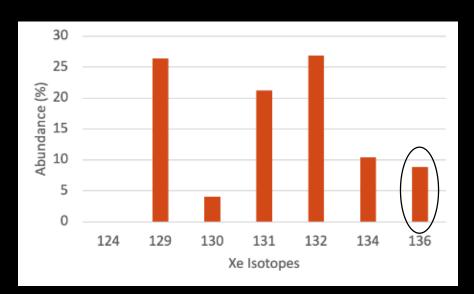




PandaX-xT

Sensitivity $\propto \sqrt{\frac{B}{Mt}}$

B = bkg rate in ROI per unit target mass



A multi-ten ton liquid xenon project at CJPL-II

- Searching for DM-nucleon interactions to neutrino floor, a decisive test on WIMP paradigm
- 2. A sensitive test on Majorana nature of neutrinos using the 0vDBD of ¹³⁶Xe, covering inverted neutrino MO parameter space
- Detecting low-energy neutrinos from the Sun and other astrophysical and terrestrial origins, and exploring other ultra-rare signals.

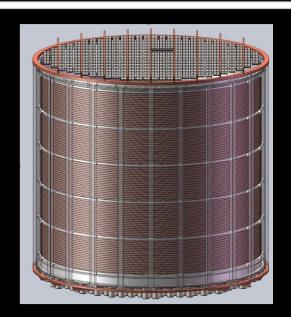
Next Generation Xe

PandaX-xT: step-wise strategy to a 40-ton-scale LXe observatory at CJPL

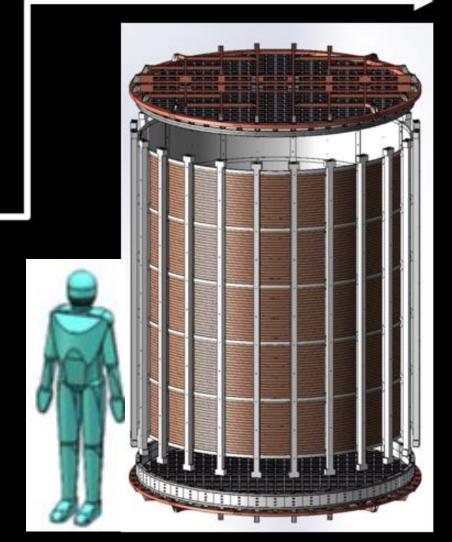
PandaX-20T (2027, mostly funded)

PandaX-4T

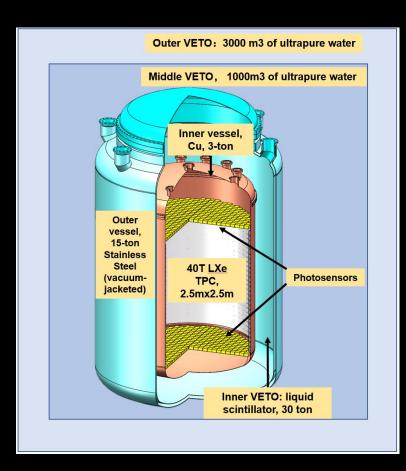


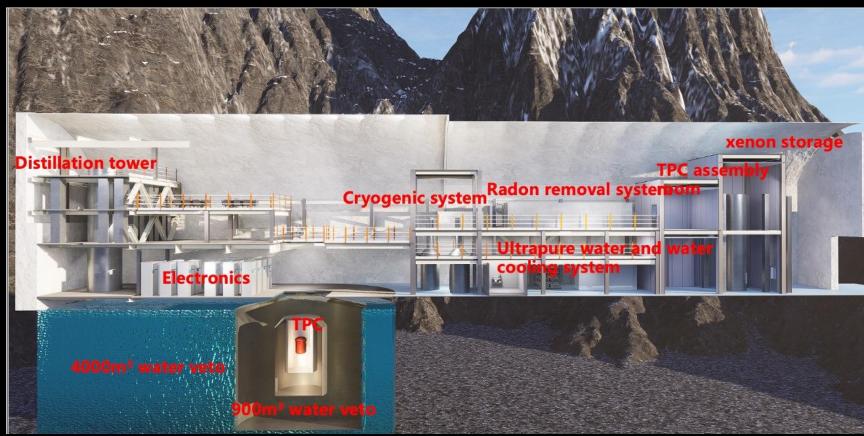


PandaX-40T



PandaX-20/40T @ CJPL





Ultrapure copper measurement

• Two batches of Aurubis NOSV copper from Germany manufacture were sampled

and screened

- Preparation methods:
 - UTEVA/TEVA resin exchange
 - Directly dilute dissolved solution
- Measurement devices:
 - PerkinElmerNexION5000 and Agilent7900 ICPMS



Resin exchange

Dilute dissolved solutions

Aurubis id	Preparation method	Measured Devices	Time	Th-232(ppt)	U-238(ppt)
#1-no.14	Dilution	SJTU ICPMS1	2025.6.12	2.18 ± 5.97	26.08 ± 14.77
#1-no.15	Dilution	PKU ICPMS	2025.6.23	4.6 ± 0.6	4.4 ± 0.3
#1-no.16	Resin exchange	SJTU ICPMS2	2025.7.21	<1	1.14 ± 0.01
#1-no.17	Dilution	SJTU ICPMS1	2025.7.27	22.4 ± 49.9	5.5±5.1
#1-no.18	Resin exchange	SJTU ICPMS1	2025.9.29	<0.74	<2.28
#2-no.6	Dilution	PKU ICPMS	2025.6.23	6.0 ± 0.5	5.1 ± 0.5
#2-no.7	Resin exchange	SJTU ICPMS2	2025.7.21	<1	1.55 ± 0.33



1.2 ton of copper shipped from Hamburg to Shanghai

TPC development

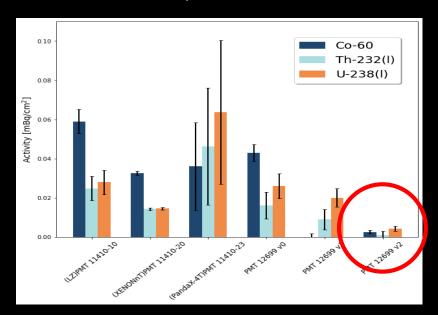


- Field cage: Kapton film with copper embedded
- Supporting structure: low radioactivity PMMA (developed in JUNO) or PTFE
- 400 kg prototype TPC under construction @TDLI
- 2.6 m electrodes built and tested

43-ton array

New photosensors

- 10 m² top/bottom coverage
- Hamamatsu R12699 (2"x2"), iterated two versions (v1/v2)
- Dark noise <10 Hz/channel, QE @178 nm > 30%
- arXiv:2412.10830, NIMA



Liquid xenon storage and handling



Each storage tank: 6-m³ (maximum 18 ton) of LXe Cryogenic pump: filling/emptying detector @ 1.5 ton/hour.

Cryogenics and recirculation



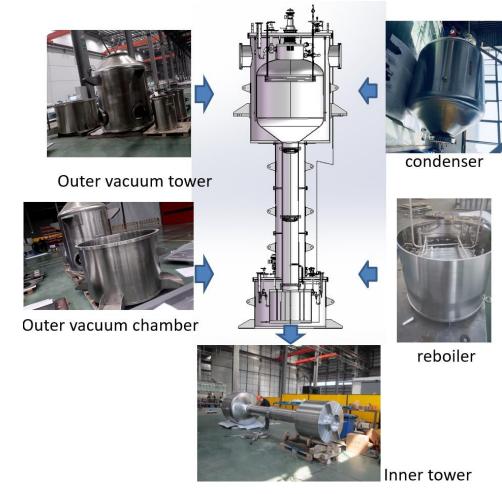
- Flexible with redundancy
- Cooling Power:
 - ≥ 1500W@178K (GM cooler)
- + others
- Online purification speed:
- 3.5-ton/day for gas, 8-ton/day for liquid

Distillation towers



Kr distillation tower: 30 kg/hour.

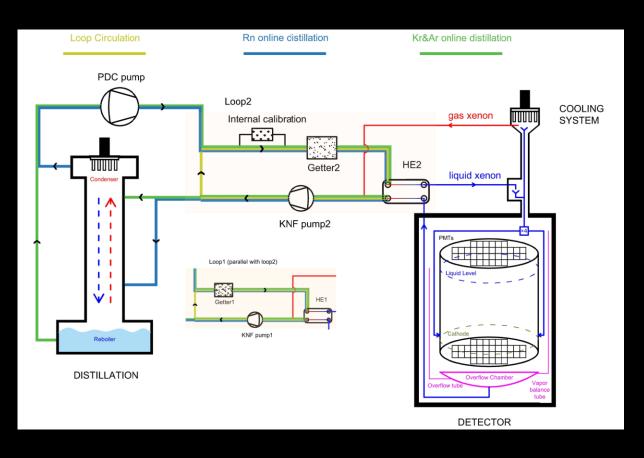
Target Kr concentration in Xe: 0.01 ppt

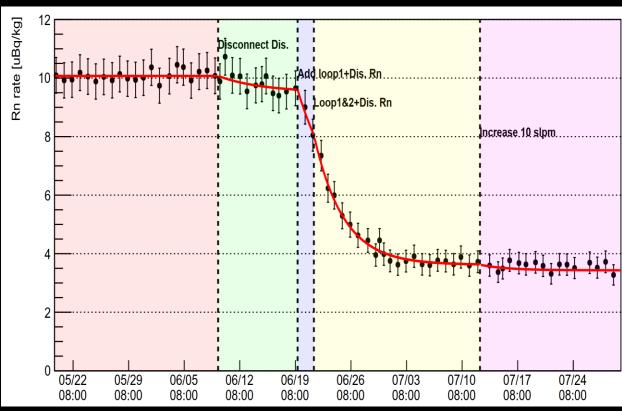


Rn distillation tower: 850 kg/hour.

Target: 0.5 μBq/kg

Rn online suppression studies with PandaX-4T





With online distillation, Rn222 level reached 3.6 uBq/kg

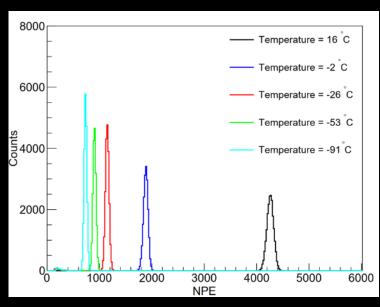
Inner Veto R&D

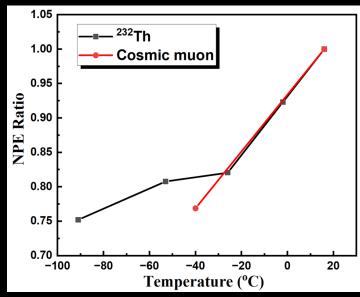
- LAB-based liquid scintillator becomes semi-transparent and viscous at liquid xenon temperature (about -100 °C)
 - pressure balance: like refrigerant fluid (HFE-7000) used in EXO, make thinner liquid xenon container
 - veto: use wavelength shifting fibers for light detection, as in LiquidO











Transporting Xenon to CJPL



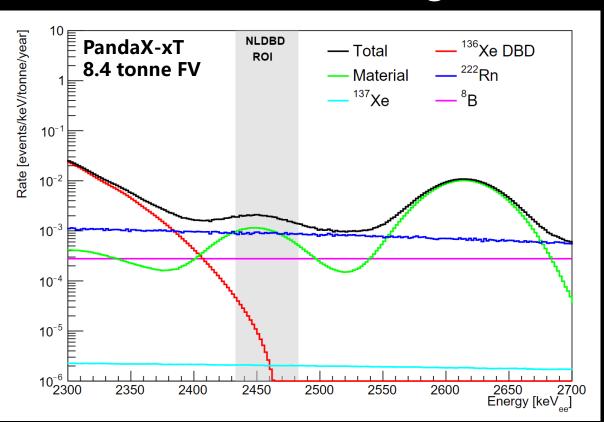


"Baseline" background

	Component	Background assumption
1	Photosensors	U/Th/K = 0.02/0.01/1.5 mBq/cm ² Hamamatsu R12699 prototype
2	Inner vessel	U/Th/K 5 ppt (g/g)
3	Kr	0.01 ppt (mol/mol)
4	Rn222	0.5 uBq/kg

10^{3} Rate [1/keV/tonne/year] Materials Solar neutrino 10^{2} 10^{-1} 10^{-2} 20 40 60 80 100 Energy [keV

Today's signal, tomorrow's background

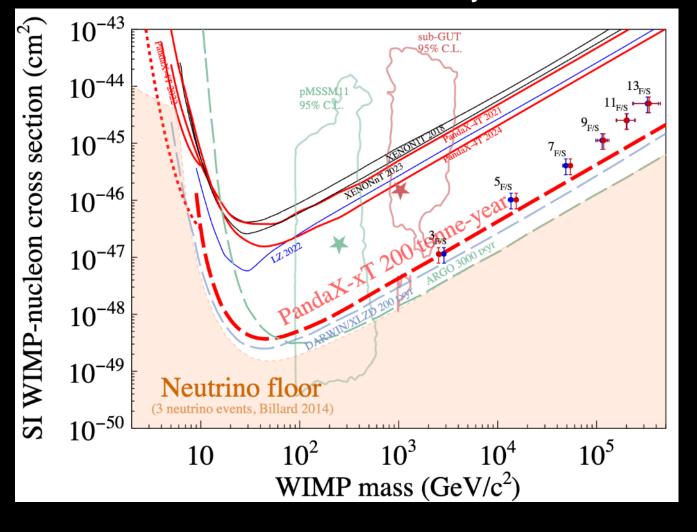


Low energy: dominated by ²²²Rn, pp neutrino, ¹³⁶Xe (34.2 ton FV)

High energy: dominated by ²²²Rn, ⁸B neutrino, and ²³⁸U in material (8.4 ton FV)

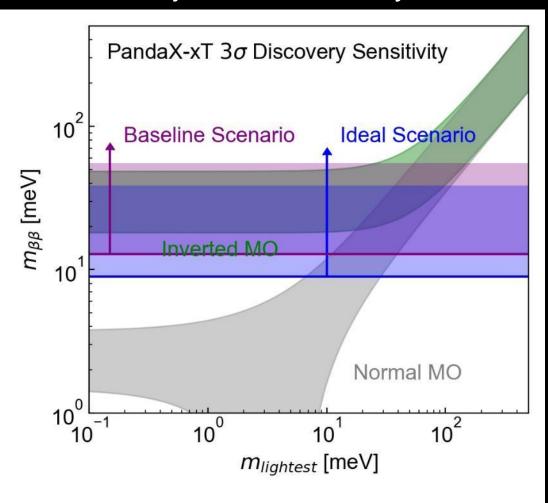


Dark matter sensitivity



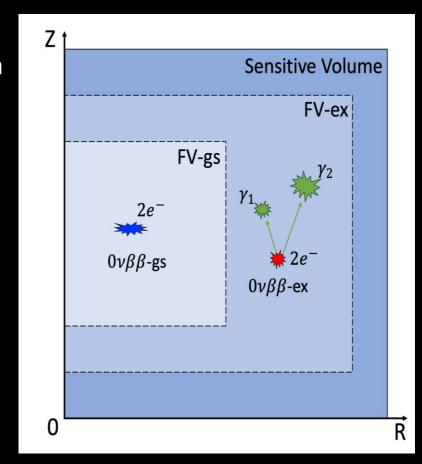
Physics reach

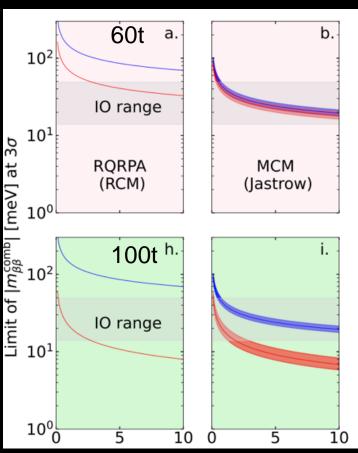
Majorana sensitivity



Unique advantage of natural xenon detector

- Large Liquid Xe TPC offers a unique detector response to search for 0vββ-ex of ¹³⁶Xe
 - Larger FV
 - Lower background rate
- Combining $0\nu\beta\beta$ -gs and $0\nu\beta\beta$ -ex improves the sensitivity to $m_{\beta\beta}$
- Depends on how lucky we are with the NME
- The bigger FV => the more favorable the combined approach with natural xenon





arXiv:2508.17413, Ding, Han, Wang, Yao

Complementarity with XLZD

- Friendly competition good for the science
- Data taking in parallel => increase the world data for global analysis
- Different experimental sites/detector design features/background
 Cross checks
- Learn from each other's lessons and technological breakthrough
- Joint force in a general way

PandaX-xT first open meeting (Apr. 7-11, 2025)





- A few groups in Russian, Japan, and UK would like to participate in the development.
- Some NSFC-CJPL international cooperation "incubator" funding

Summary

- PandaX-4T has developed a very rich physics program in DM and Neutrinos(!)
 - Technically: developing more and more sensitive detectors
 - Physics-wise: stronger connections with theorists
- A pragmatic approach with a stage-wise upgrade to PandaX-xT aiming for VERY exciting physics
 - Next stage: PandaX-20T
- Highly welcome collaborators!
- Hiring postdoc and students!