

Characterizing and removing ER background events in the DEAP-3600 experiment

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On Behalf of the DEAP-3600 Collaboration



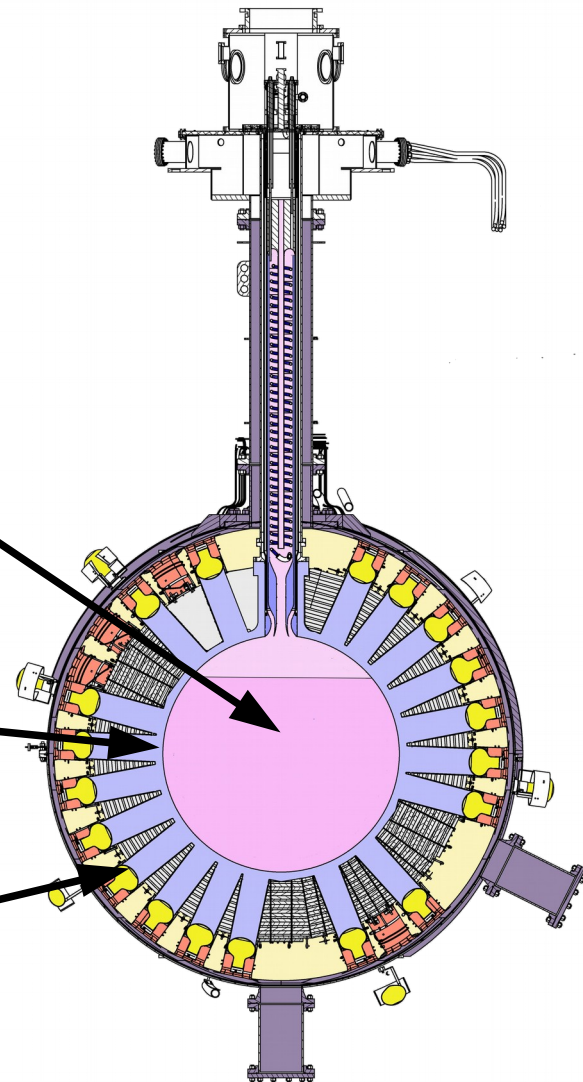
DEAP-3600

- Direct dark matter search experiment
- Detection method:
 - Argon scintillates when excited by interacting particles
 - Good light yield
 - 40,000 photons/MeV
 - Peak emission 128 nm
 - Shifted to 420 nm by TPB
 - Shifted light detected by PMTs

3000 kg LAr Target

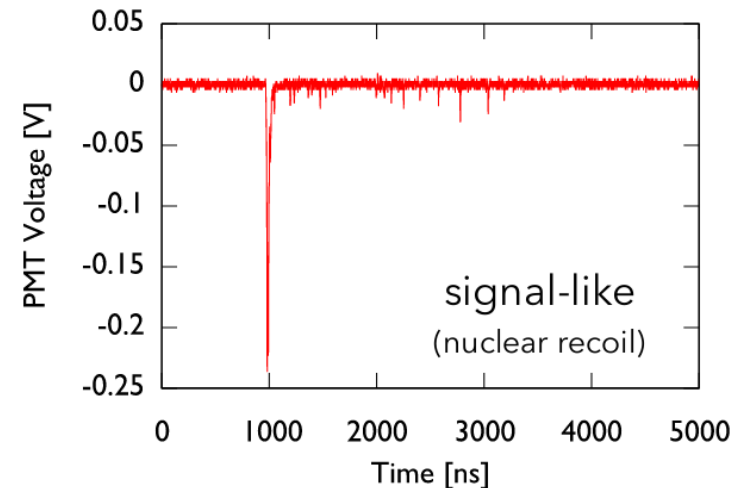
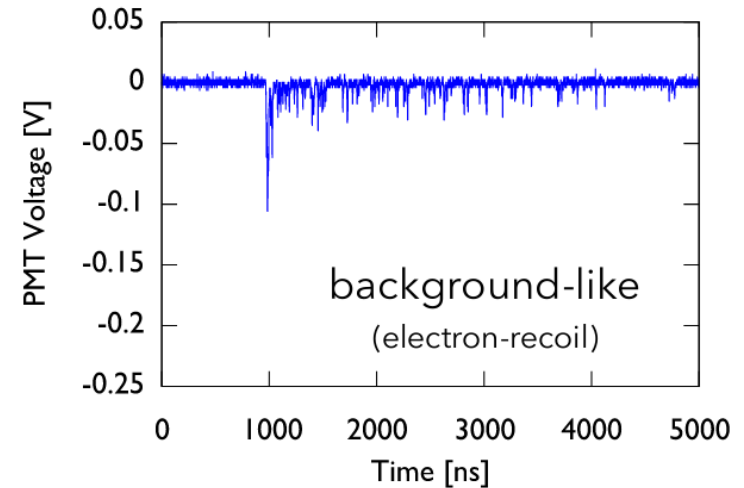
Acrylic vessel – inner surface coated with TPB

255 HQE PMTs



Argon Scintillation

- Argon has two dominant excited states
 - Singlet (6 ns)
 - Triplet (1.4 μ s)
- Nuclear Recoils (NR)
i.e. (WIMP, α , n) \rightarrow Singlet State
- Electron Recoils (ER)
i.e. (β , μ , γ) \rightarrow Triplet State
- NR events have much more light earlier in the pulse



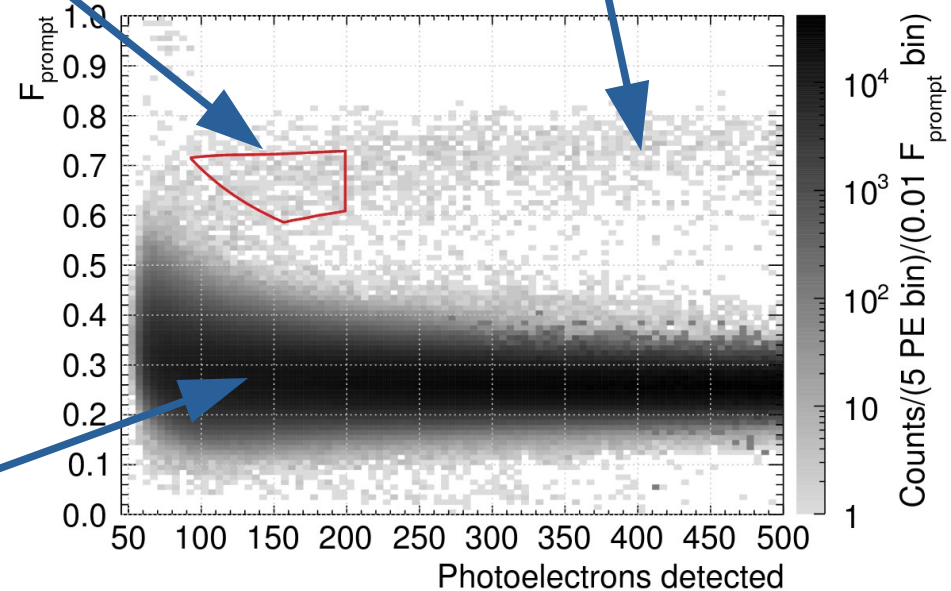
Pulse shape Discrimination

WIMP ROI defined to contain 0.05 ^{39}Ar events for the entire livetime of the dataset

NR band produced by AmBe neutron source

$$F_{\text{prompt}} = \frac{\sum_{t=-28\text{ ns}}^{60\text{ ns}} \text{PE}(t)}{\sum_{t=-28\text{ ns}}^{10\text{ }\mu\text{s}} \text{PE}(t)}.$$

^{39}Ar has activity of 1 Bq/kg of Atmospheric Ar (~3300 Bq in DEAP-3600)



<https://arxiv.org/abs/1902.04048>

Excluding ^{39}Ar events

High rate of ^{39}Ar allows effective model for F_{prompt} distribution of ^{39}Ar

$$F^{\text{ER}}(f, q) = \Gamma(f; \bar{f}, b) \otimes \text{Gauss}(f; \sigma),$$

$$\bar{f}(q) = a_0 + \frac{a_1}{q - a_2} + \frac{a_3}{(q - a_4)^2},$$

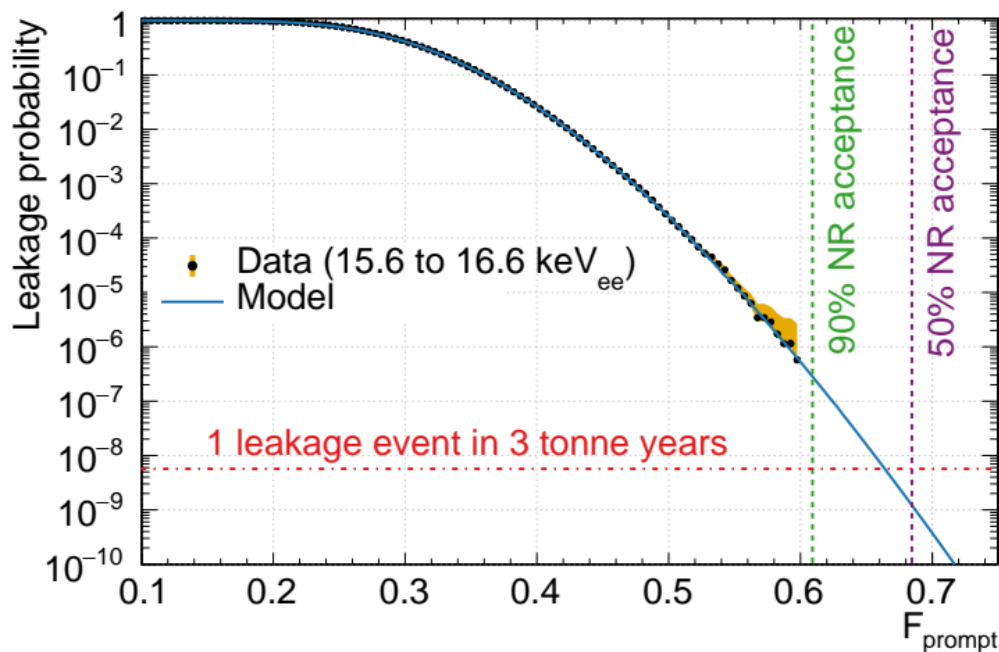
$$b(q) = a_5 + \frac{a_6}{q} + \frac{a_7}{q^2},$$

$$\sigma(q) = a_8 + \frac{a_9}{q} + \frac{a_{10}}{q^2},$$

q is number of PE

Total leakage can be obtained via:

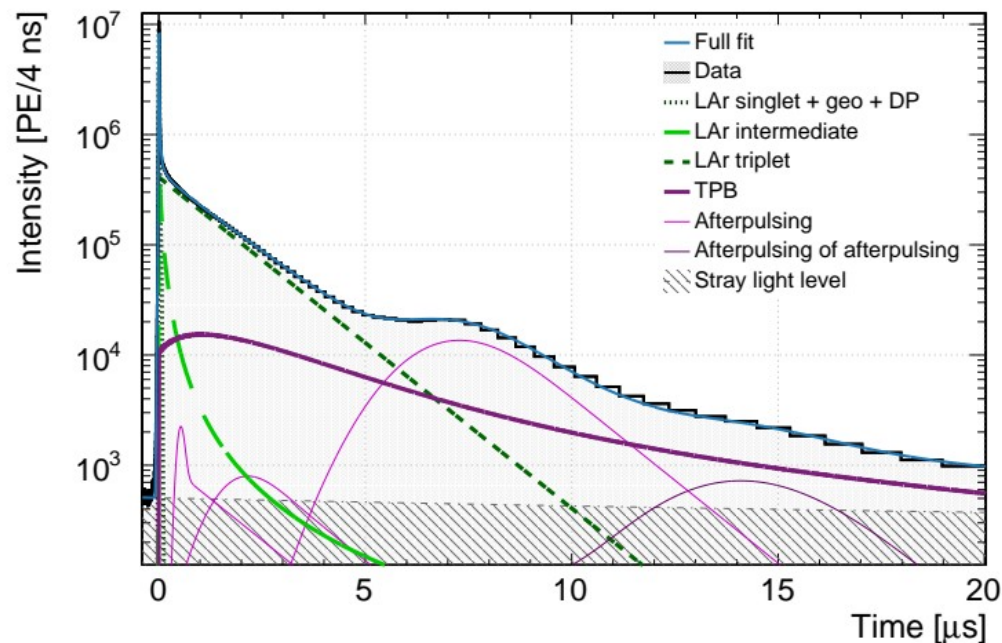
$$L = T \sum_q R(q) \times P_{\text{Leak}}(f, q)$$



F_{prompt} distribution in lowest energy bin used in WIMP search

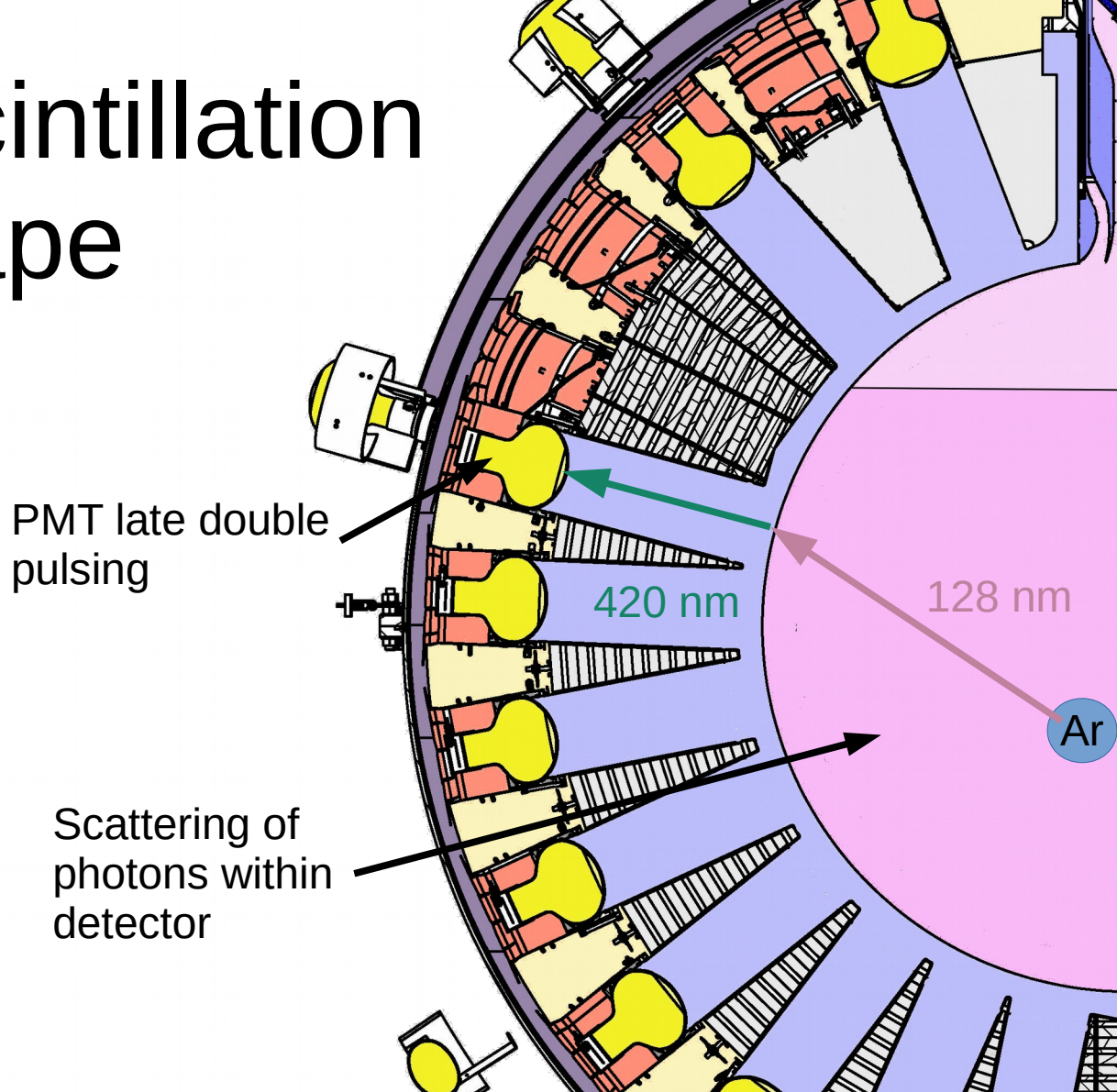
Measuring the Scintillation Pulse Shape

- Fprompt provides power to discriminate ER from NR events
- Constructing a scintillation model
 - Allows ongoing monitoring of the detector
 - Motivates design choices for future detectors

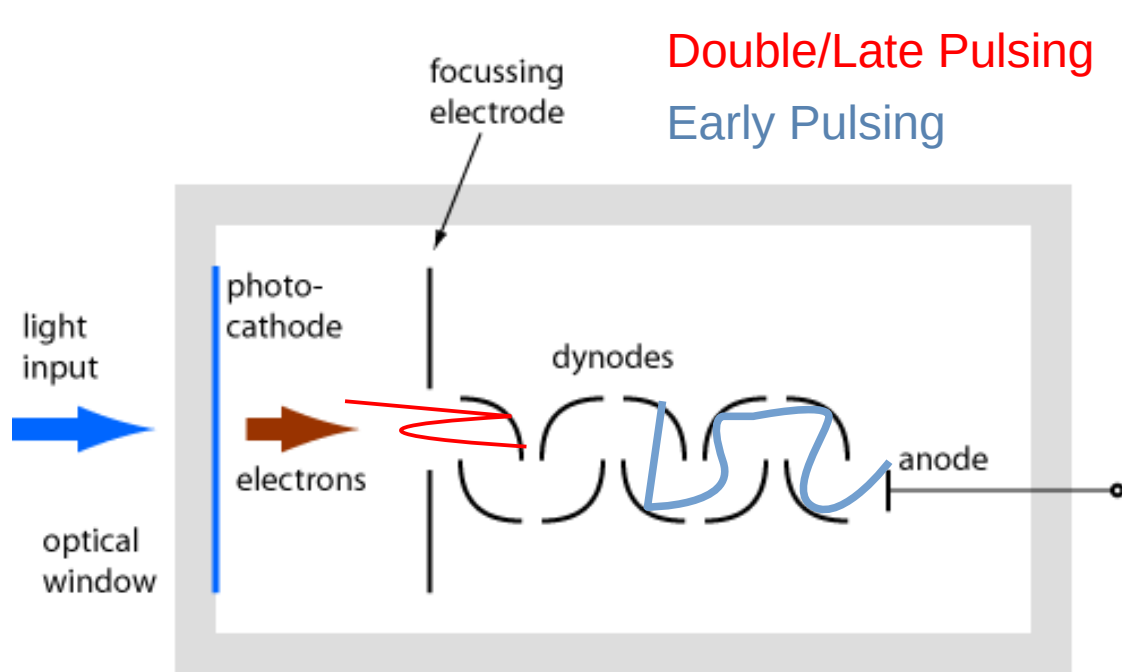


Modelling the Scintillation Pulse Shape

- Double exponential structure of LAr scintillation
- Instrumental Effects:
 - TPB Re-emission
 - PMT effects



Instrumental Effects: PMT/Detector Effects



Double/Late Pulsing

Early Pulsing

Prompt pulsing
15 ns width due to
scattering of light

$$I_{\text{geo}}(t) = \underbrace{\nu_{\text{DET}} \cdot \text{Gaus}(t, \mu_{\text{DET}}, \sigma_{\text{DET}})}_{\text{Prompt pulsing}} + \underbrace{\nu_{\text{DP}} \cdot \text{Gaus}(t, \mu_{\text{DP}}, \sigma_{\text{DP}})}_{\text{Double/Late pulsing}}$$

Double/Late pulsing
Peak ~58 ns late

<https://www.rp-photonics.com/photomultipliers.html>

$$I_{\text{EP}}(t) = R_{\text{EP}} \cdot (I_{\text{LAr}}(t - t_{\text{EP}}) \otimes I_{\text{TPB}}(t - t_{\text{EP}}) \otimes I'_{\text{geo}}(t - t_{\text{EP}}))$$

Argon Scintillation

$$I_{\text{LAr}}(t) = \underbrace{\frac{R_s}{\tau_p} e^{-t/\tau_p}}_{\text{Singlet emission}} + \underbrace{\frac{1 - R_s - R_t}{(1 + t/\tau_{\text{rec}})^2} \frac{1}{\tau_{\text{rec}}}}_{\text{Intermediate emission}} + \underbrace{\frac{R_t}{\tau_t} e^{-t/\tau_t}}_{\text{Triplet emission}}$$

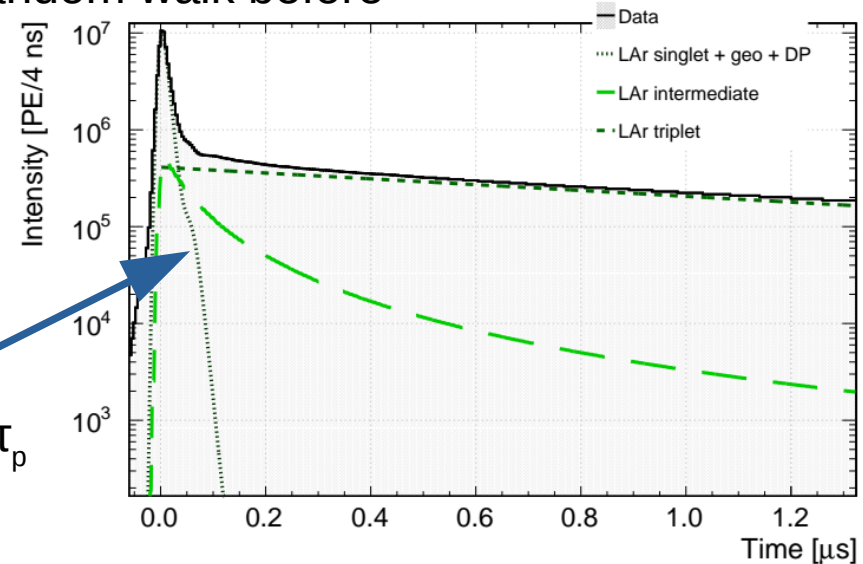
Singlet emission
– decay time
modified to reflect
TPB response in
fit

Intermediate emission – electrons scattered far
away from ions → random walk before
recombination

$$I_{\text{LAr}}(t) \otimes I_{\text{geo}}(t)$$

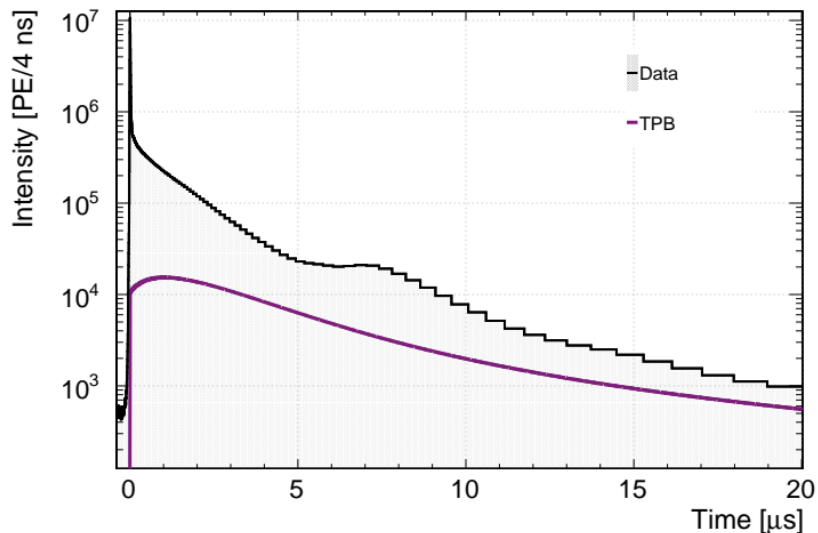
Double pulsing produces
shoulder in singlet peak

Assumes fast TPB response is absorbed into I_{geo} and τ_p



Instrumental Effect: Delayed TPB Re-emission

$$I_{\text{TPB}}(t) = \underbrace{(1 - R_{\text{TPB}})\delta(t)}_{\text{Prompt re-emission modelled by delta function}} + \underbrace{\frac{R_{\text{TPB}} \cdot N_{\text{TPB}} \cdot e^{-2t/\tau_T}}{1 + A_{\text{TPB}}[Ei(-\frac{t+t_a}{\tau_T}) - Ei(-\frac{t_a}{\tau_T})]^2(1 + t/t_a)}}_{\text{Delayed TPB emission using Voltz and Laustriat model}}$$

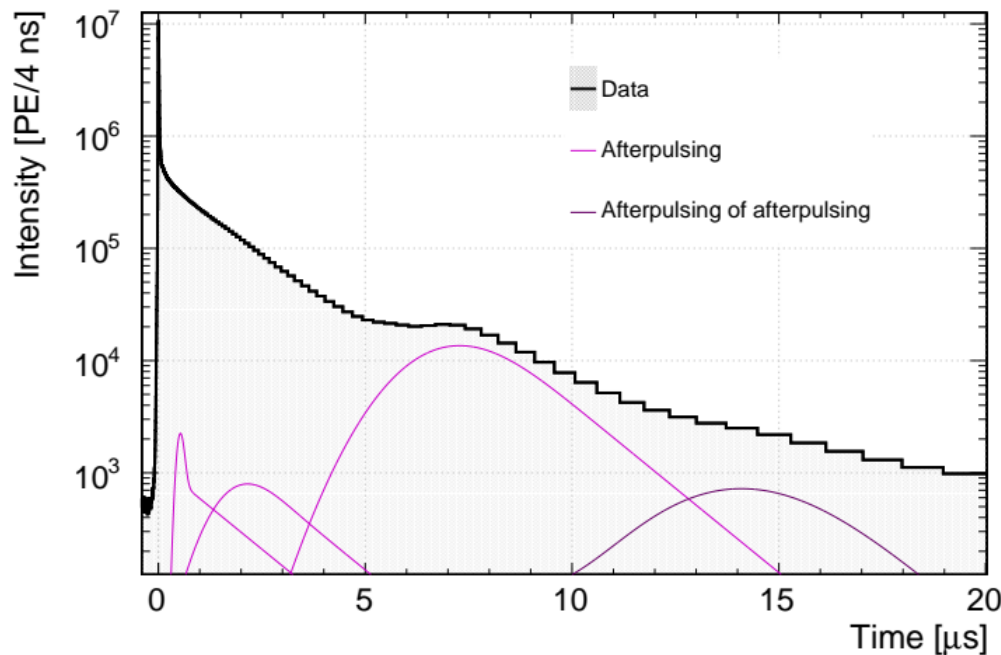


Delayed TPB emission using Voltz and Laustriat model (<https://arxiv.org/abs/1804.06895v1>)

$$I_{\text{LAr}}(t) \otimes I_{\text{TPB(Late)}}(t)$$

Delayed emission caused by Ar scintillation photons

Instrumental Effect: PMT Afterpulsing



$$I_{AP}(t) = \sum_{i=1}^3 \nu_{APi} \cdot \text{Gaus}(t, \mu_{APi}, \sigma_{APi})$$

$$\mu_{APi} = 0.5, 1.7, 6.3 \mu\text{s}$$

$$I_{APofAP}(t) = I_{AP}(t) \otimes I_{AP}(t)$$

Like TPB re-emission, AP is in response to detected LAr photons

$$I_{LAr}(t) \otimes [I_{AP}(t) + I_{APofAP}(t)]$$

Instrumental Effect: Significantly delayed TPB emission from previous events

Sum up all contributions of events before Δt

Number of PE in window due to event $\Delta t'$ before

$$\bar{I}_{\text{stray}}(\Delta t) = \frac{\int_{\Delta t}^{\infty} N_p(\Delta t') d\Delta t'}{\int_{\Delta t}^{\infty} N_{ev}(\Delta t') d\Delta t'}$$

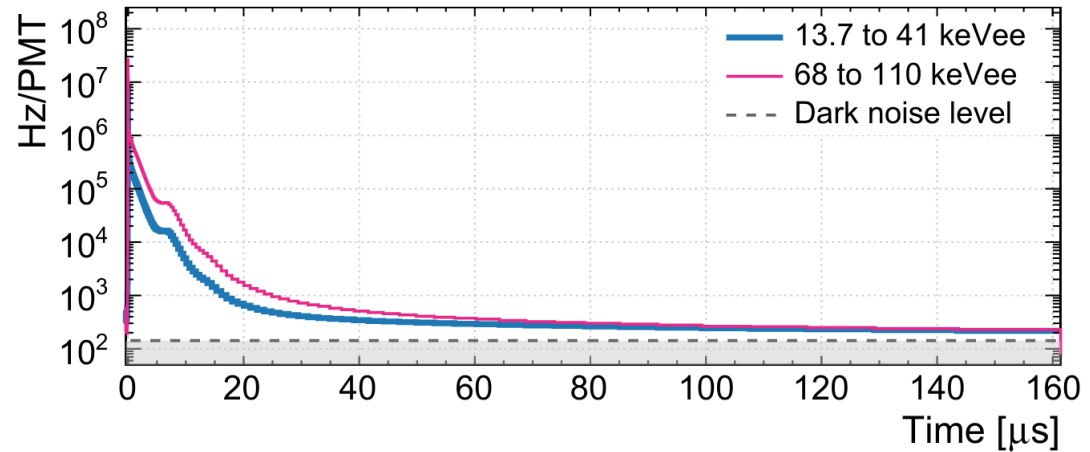
Total number of events with $\Delta t'$

$$\eta \cdot \bar{I}_{\text{stray}}(\underbrace{\Delta T_{\text{cut}} + 1.6\mu s}_{} + t)$$

Minimum cut time between events (200 μs)

Correction as we evaluate N_p in pre-trigger window

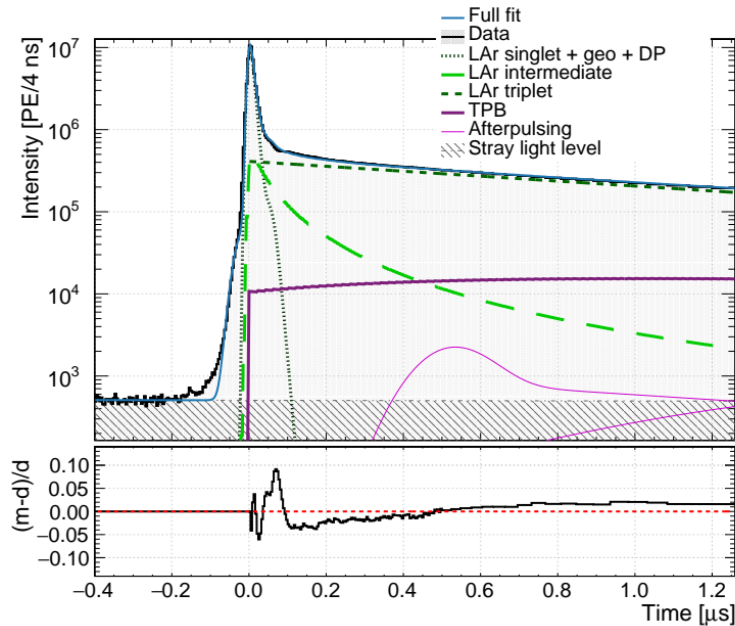
Additional light at late times thought to be caused by very long lived TPB component



At later times in the event window stray light still significantly higher than dark noise

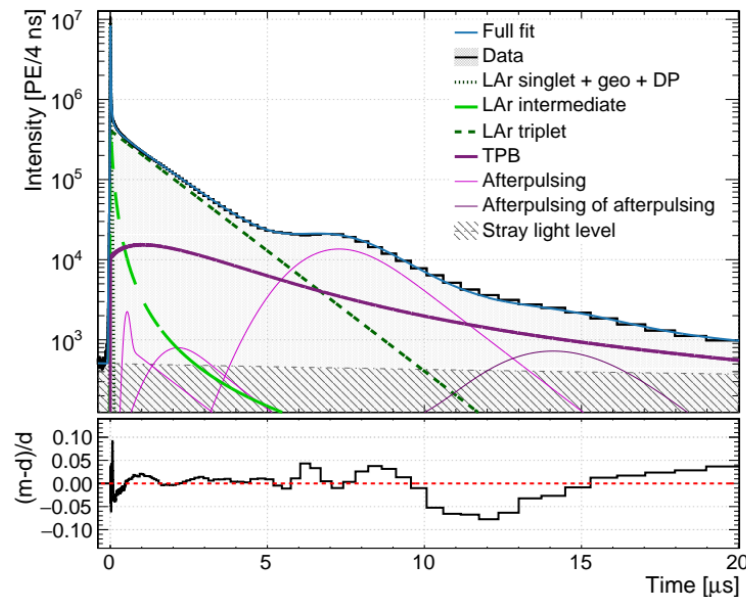
The Full Model

Model agrees
within 10% of data



LAr scintillation + TPB
late light

$$I_{PS}(t) = \underbrace{\mathbf{I}_0 \cdot \left(I_{LAr}(t) \otimes I_{TPB}(t) \otimes I_{geo}(t) \right)}_{\text{LAr scintillation + TPB late light}} + \underbrace{\mathbf{I}_0 \cdot R_{EP} \cdot (I_{LAr}(t - t_{EP}) \otimes I_{TPB}(t - t_{EP}) \otimes I'_{geo}(t - t_{EP}))}_{\text{Early Pulsing}} + \underbrace{\eta \cdot \bar{I}_{stray}(\Delta T_{cut} + 1.6\mu s + t)}_{\text{Stray light}} + \underbrace{\mathbf{I}_0 \cdot R_{EP} \cdot (I_{LAr}(t - t_{EP}) \otimes I_{TPB}(t - t_{EP}) \otimes I'_{geo}(t - t_{EP}))}_{\text{Afterpulsing}}$$

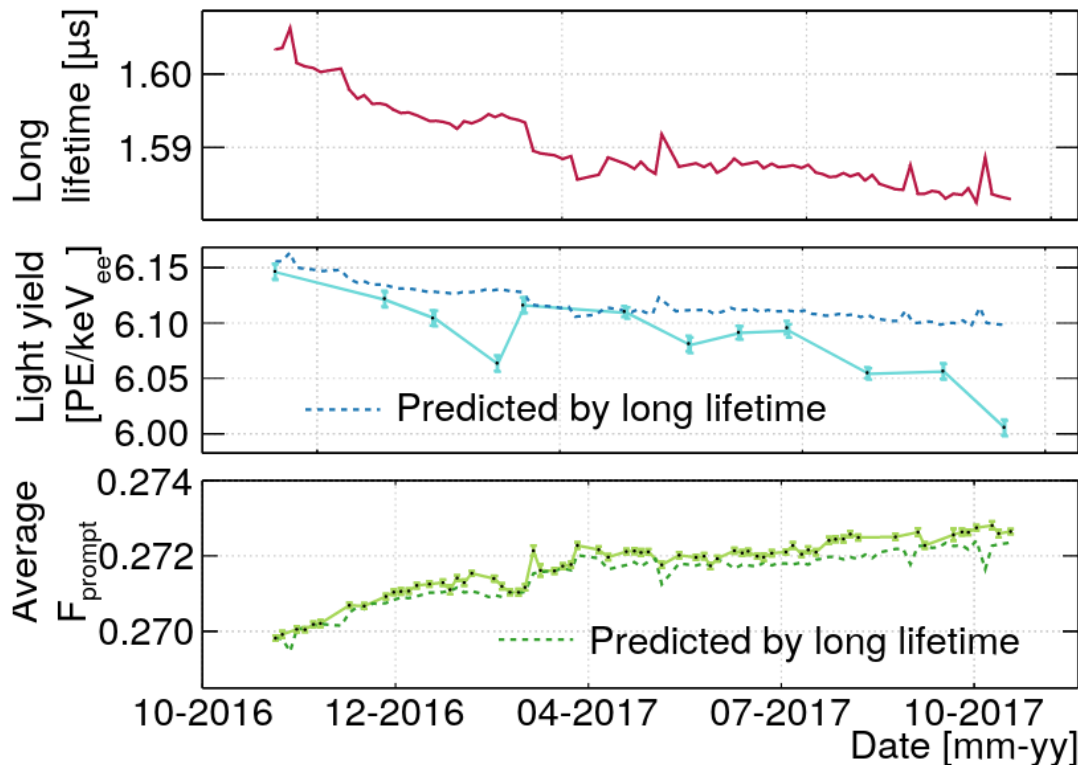


Afterpulsing

Monitoring the Pulse Shape

- High rate of ^{39}Ar means stability of detector can be measured in periods < 1 day
- Throughout analysis for 231 days dataset long lifetime was stable within 1 %

Long lifetime includes detector effects
– not just LAr triplet livetime



Detector response to 2.61 MeV γ -rays in LAr
<https://arxiv.org/abs/1902.04048>

Conclusions

- ^{39}Ar backgrounds can be effectively removed by evaluating the fraction of light in the start of the event window
- A detailed model including long lived TPB emission has been developed to model the response to ^{39}Ar decays
 - Allows monitoring of the detector during running i.e. variations in the AP rate and Ar purity without dedicated calibration runs
 - Motivates the design choices for a future experiment using LAr and TPB

For more details see: <https://arxiv.org/abs/2001.09855>