Superfluid neutron matter with a twist: From particles to matter

George Palkanoglou

## Superfluid neutron matter with a twist: *From particles to matter*

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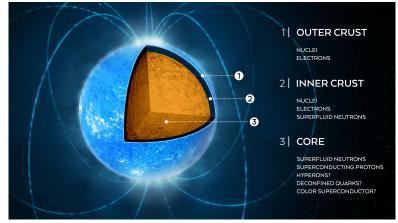
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# Superfluid neutron matter

- Where to find it & why care about it
- How to study it
- Results: How to study infinite superfluid neutron matter with finite systems

## Motivation & physical system

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Credit: Anna L. Watts

#### The best of both worlds

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Ab initio

- No extra assumptions
- Computationally expensive (only smallish N is feasible)

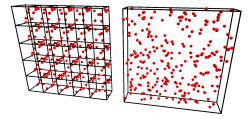
Phenomenology

- Easier to implement
- Uncontrolled approximations

Phenomenology can guide *ab initio Ab initio* can constrain phenomenology

#### Twisted Boundary Conditions

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Credit: Nawar Ismail

$$\begin{aligned} |\psi(\mathbf{r}_1 + L\hat{\mathbf{x}}, \dots, \mathbf{r}_N)|^2 &= |\psi(\mathbf{r}_1, \dots, \mathbf{r}_N)|^2 \\ \psi(\mathbf{r}_1 + L\hat{\mathbf{x}}, \dots, \mathbf{r}_N) &= \frac{e^{i\theta_x}\psi(\mathbf{r}_1, \dots, \mathbf{r}_N) \end{aligned}$$

#### Twisted Boundary Conditions

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$$|\psi(\mathbf{r}_1 + L\hat{\mathbf{x}}, \dots, \mathbf{r}_N)|^2 = |\psi(\mathbf{r}_1, \dots, \mathbf{r}_N)|^2$$

#### **Periodic Boundary Conditions**

#### Twisted boundary conditions

$$\psi \left( \mathbf{r}_{1} + L\hat{x}, \mathbf{r}_{2}, \dots, \mathbf{r}_{N_{0}} \right) = e^{i\theta_{x}} \psi \left( \mathbf{r}_{1}, \dots, \mathbf{r}_{N_{0}} \right)$$
$$\Downarrow$$
$$\mathbf{k} = \frac{2\pi}{L} \left( \mathbf{n} + \frac{\boldsymbol{\theta}}{2\pi} \right)$$

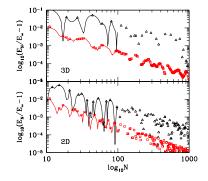
#### Twist-averaged boundary conditions

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Averaging over  $\boldsymbol{\theta}$  can reduce finite-size effects

$$\left\langle \hat{F} \right\rangle = \int \frac{d^{3}\boldsymbol{\theta}}{\left(2\pi\right)^{3}} \left\langle \psi(\boldsymbol{\theta}) \right| \hat{F} \left| \psi(\boldsymbol{\theta}) \right\rangle$$



C. Lin, F. H. Zong, and D.M. Ceperley Phys. Rev. E **64**, 016702 (2001)

## The pairing Hamiltonian

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 $\hat{\mathcal{H}} = \sum_{\mathbf{k}\sigma} \epsilon_{\mathbf{k}} \hat{c}^{\dagger}_{\mathbf{k}\sigma} \hat{c}_{\mathbf{k}\sigma} + \sum_{\mathbf{k}\mathbf{l}} V_{\mathbf{k}\mathbf{l}} \hat{c}^{\dagger}_{\mathbf{k}\uparrow} \hat{c}^{\dagger}_{-\mathbf{k}\downarrow} \hat{c}_{-\mathbf{l}\downarrow} \hat{c}_{\mathbf{l}\uparrow}$ 



## BCS and PBCS theories for neutron matter

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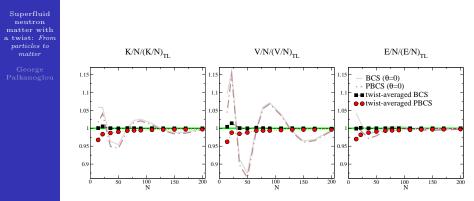
PBCS:

fixed-N

$$\begin{split} |\psi_N\rangle &= \int_0^{2\pi} \frac{d\phi}{2\pi} e^{-i\frac{N}{2}\phi} \prod_{\mathbf{k}} (u_{\mathbf{k}} + e^{i\phi} v_{\mathbf{k}} \hat{c}^{\dagger}_{\mathbf{k}\uparrow} \hat{c}^{\dagger}_{-\mathbf{k}\downarrow}) |0\rangle \\ & \to \boxed{\Delta(N) = E(N+1) - \frac{1}{2} \left[ E(N) + E(N+2) \right]} \end{split}$$

for even N

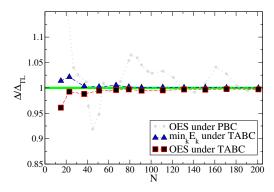
#### Twist-averaged energy



G. Palkanoglou and A. Gezerlis, Universe 2021, 7(2), 24

#### Twist-averaged pairing gap

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G. Palkanoglou and A. Gezerlis, Universe 2021, 7(2), 24

$$\Delta(N) = E(N+1) - \frac{1}{2} \left[ E(N) + E(N+2) \right]$$

#### Conclusions

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- $\bigstar$  Twist-averaged boundary conditions work well for superfluid systems
- $\bigstar$  We now have a better prescription on how to approximate infinite superfluid neutron matter with a finite system

## Acknowledgements

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#### Acknowledgements:

Alex Gezerlis

Computational Resources:

- NERSC
- SHARCNET



#### Conclusions

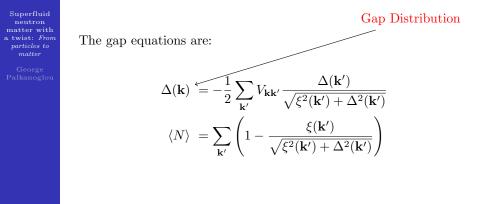
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- $\bigstar$  Twist-averaged boundary conditions work well for superfluid systems
- ★ We now have a better prescription on how to approximate infinite superfluid neutron matter with a finite system

# Thank you

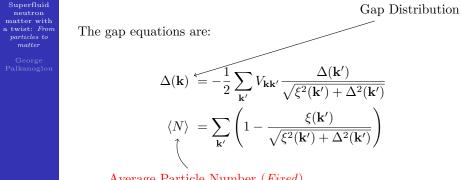
#### BCS Theory and the gap Equations



where:

$$\xi(\mathbf{k}) = \frac{\hbar^2}{2m_n} |\mathbf{k}|^2 - \mu$$

#### BCS Theory and the gap Equations

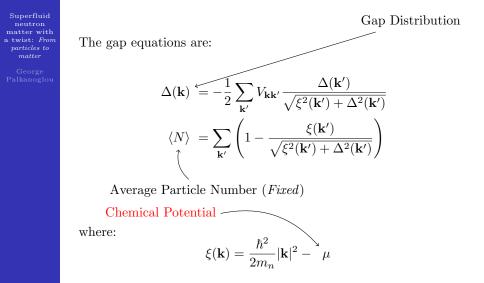


Average Particle Number (*Fixed*)

where:

$$\xi(\mathbf{k}) = \frac{\hbar^2}{2m_n} |\mathbf{k}|^2 - \mu$$

#### BCS Theory and the gap Equations



#### The solution of the BCS gap Equations

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$$\begin{split} u_{\mathbf{k}}^2 &= \frac{1}{2} \left( 1 + \frac{\xi(\mathbf{k})}{\sqrt{\xi^2(\mathbf{k}) + \Delta^2(\mathbf{k})}} \right) \\ v_{\mathbf{k}}^2 &= \frac{1}{2} \left( 1 - \frac{\xi(\mathbf{k})}{\sqrt{\xi^2(\mathbf{k}) + \Delta^2(\mathbf{k})}} \right) \end{split}$$

where:

 $v_{\mathbf{k}}^2+u_{\mathbf{k}}^2=1$ 

#### Odd and even particle numbers

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$$\Delta(N) = E(N+1) - \frac{1}{2} \left[ E(N) + E(N+2) \right]$$

$$\begin{aligned} |\psi_{\phi}\rangle &= \prod_{\mathbf{k}} (u_{\mathbf{k}} + e^{i\phi} v_{\mathbf{k}} \hat{c}^{\dagger}_{\mathbf{k}\uparrow} \hat{c}^{\dagger}_{-\mathbf{k}\downarrow}) |0\rangle \\ & (\text{even systems}) \\ \left|\psi_{\phi}^{\mathbf{b}\gamma}\right\rangle &= \hat{c}^{\dagger}_{\mathbf{b}\gamma} \prod_{\mathbf{k}\neq\mathbf{b}} (u_{\mathbf{k}} + e^{i\phi} v_{\mathbf{k}} \hat{c}^{\dagger}_{\mathbf{k}\uparrow} \hat{c}^{\dagger}_{-\mathbf{k}\downarrow}) |0\rangle \\ & (\text{odd systems}) \end{aligned}$$

BCS is formulated in a Grand Canonical Ensemble.

## PBCS - the Projected Energy

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George Palkanoglou The energy of the projected states is:

$$\begin{split} E_{N} =& 2\sum_{\mathbf{k}}\epsilon_{\mathbf{k}}v_{\mathbf{k}}^{2}\frac{R_{1}^{1}(\mathbf{k})}{R_{0}^{0}} + \sum_{\mathbf{kl}}V_{\mathbf{kl}}u_{\mathbf{k}}v_{\mathbf{k}}u_{l}v_{l}\frac{R_{1}^{2}(\mathbf{kl})}{R_{0}^{0}} \ ,\\ E_{N+1} =& 2\sum_{\mathbf{k}}\epsilon_{\mathbf{k}}v_{\mathbf{k}}^{2}\frac{R_{1}^{2}(\mathbf{bk})}{R_{0}^{1}(\mathbf{b})} + \\ &+ \sum_{\mathbf{kl}}V_{\mathbf{kl}}u_{\mathbf{k}}v_{\mathbf{k}}u_{l}v_{l}\frac{R_{1}^{3}(\mathbf{bkl})}{R_{0}^{1}(\mathbf{b})} + \frac{\hbar^{2}}{2m_{n}}|\mathbf{b}|^{2} \ . \end{split}$$

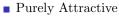
## The Potential

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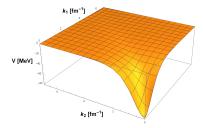
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#### The Modified Poschl-Teller Potential:

$$V(r) = -\lambda(\lambda - 1)\frac{\hbar^2}{m_n}\frac{q^2}{\cosh^2(qr)}$$



Finite Range



#### The residuum integrals

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George Palkanoglou The residuum integrals:

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$$\begin{aligned} R_n^m(\mathbf{k}_1\mathbf{k}_2\dots\mathbf{k}_N)(M) &= \\ &= \int_0^{2\pi} \frac{d\phi}{2\pi} e^{-iM\phi} e^{in\phi} \prod_{\mathbf{k}\neq\mathbf{k}_1,\mathbf{k}_2,\dots,\mathbf{k}_m} (u_\mathbf{k}^2 + e^{i\phi}v_\mathbf{k}^2) \end{aligned}$$

where 
$$M = \frac{N}{2}$$